

The Higgs and Naturalness:

Where do we stand after the LHC Run I?

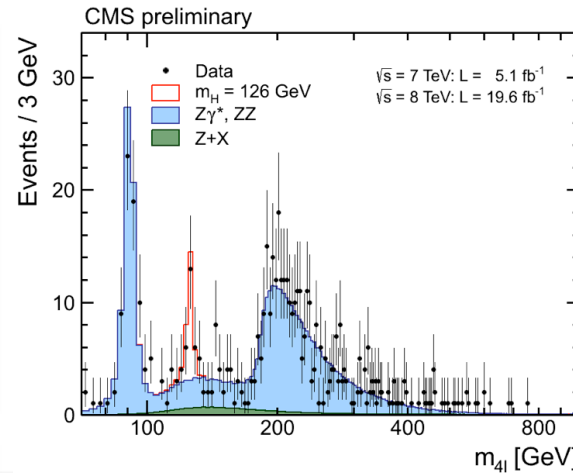
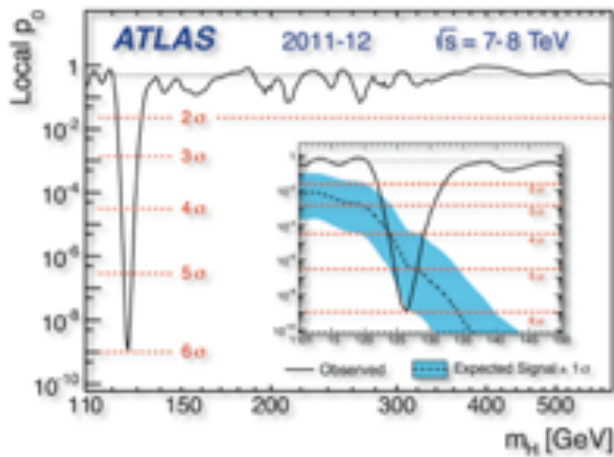
Tony Gherghetta

University of Minnesota

Johns Hopkins Workshop, Heidelberg, Germany, July 22, 2014

Based on recent work with *James Barnard, Ben von Harling, Anibal Medina, Tirtha Sankar Ray, Michael Schmidt*

Higgs discovery - Run I



Higgs Bosons — H^0 and H^\pm

PDG 2013

H^0 Mass $m = 125.9 \pm 0.4$ GeV

H^0 signal strengths in different channels $[\mu]$

Combined Final States = 1.07 ± 0.26 ($S = 1.4$)

WW^* Final State = 0.88 ± 0.33 ($S = 1.1$)

ZZ^* Final State = $0.89^{+0.30}_{-0.25}$

$\gamma\gamma$ Final State = 1.65 ± 0.33

$b\bar{b}$ Final State = $0.5^{+0.8}_{-0.7}$

$\tau^+\tau^-$ Final State = 0.1 ± 0.7

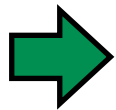
Higgs potential:

$$V(h) = -\mu_h^2 |H|^2 + \lambda_h |H|^4$$

$$\langle H \rangle = \frac{1}{\sqrt{2}}(v + h)$$

$$v^2 = \frac{\mu_h^2}{\lambda_h} \simeq (246 \text{ GeV})^2$$

$$m_h^2 = 2\lambda_h v^2 \simeq (126 \text{ GeV})^2$$



$$\mu_h^2 \simeq (89 \text{ GeV})^2$$

$$\lambda_h \simeq 0.13$$

Standard Model and nothing else...?

"In this field, almost everything is already discovered, and all that remains is to fill a few unimportant holes."

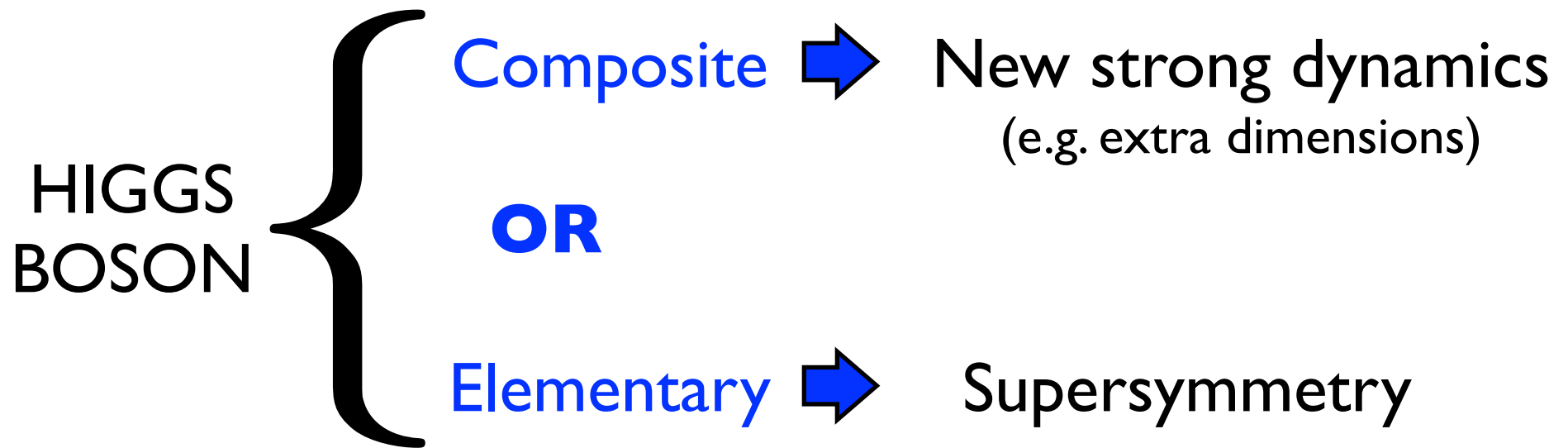
[Max Planck's PhD adviser, Philipp von Jolly, 1878]



A few “holes” in the Standard Model

- Planck/weak scale hierarchy? ($\mu_h \ll M_P$)
- Fermion mass hierarchy? Neutrino masses?
- Dark matter?
- Baryon asymmetry?
- Strong CP problem?
- GUTS? Inflation?
- UV completion of gravity?
- Cosmological constant?

NATURAL explanations of ~ 126 GeV Higgs



How “natural” are these two possibilities after Run 1?

1. Composite Higgs

- Higgs is pseudo Nambu-Goldstone boson [Georgi, Kaplan '84]

$$G \rightarrow H \quad \text{at scale } f \quad \text{where} \quad H \supset SO(4) \sim SU(2)_L \times SU(2)_R$$

- Partially composite top $\mathcal{L} = \lambda_L t_L \mathcal{O}_R + \lambda_R t_R \mathcal{O}_L$
[Kaplan '91; Agashe, Contino, Pomarol '04]

$$m_t \sim \lambda_L \lambda_R v \quad \text{where} \quad \lambda_{L,R} \sim \left(\frac{\Lambda}{\Lambda_{UV}} \right)^{\dim \mathcal{O}_{L,R} - \frac{5}{2}} \quad \Rightarrow \quad \dim \mathcal{O}_{L,R} \sim \frac{5}{2}$$

Higgs potential

$$V(h) = -\mu_h^2 |H|^2 + \lambda_h |H|^4 \quad \text{where} \quad \mu_h^2 \sim \frac{g_{SM}^2}{16\pi^2} g_\rho^2 f^2 \quad \lambda_h \sim \frac{g_{SM}^2}{16\pi^2} g_\rho^2$$

$$\text{EWSB} \left(\langle H \rangle = \frac{v}{\sqrt{2}} \right) \quad v^2 = \frac{\mu_h^2}{\lambda_h} \quad \Rightarrow$$

$$\text{Tuning: } \Delta^{-1} \sim \frac{v^2}{f^2} \lesssim 10\%$$

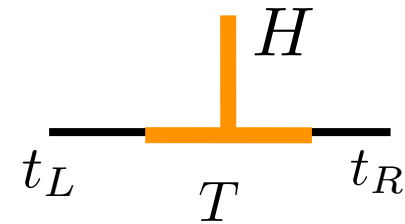
$$(v = 246 \text{ GeV}, f \gtrsim 750 \text{ GeV})$$

[See also Panico, Redi, Tesi, Wulzer 1210.7114]

Higgs mass:

$$m_h^2 \simeq \frac{N_c}{\pi^2} m_t^2 \frac{m_T^2}{f^2} = g_T^2$$

m_T = fermion resonances (EM charges 5/3, 2/3, -1/3)



$m_h \sim 126 \text{ GeV} \rightarrow m_T < m_\rho$

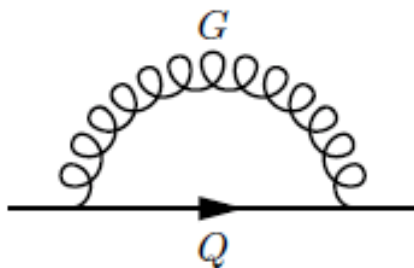
light fermion resonances

[Matsedonskyi, Panico, Wulzer '12]

[Marzocca, Serone, Shu '12; Pomarol, Riva '12]

e.g. LHC limit: $m_{T_{5/3}} \gtrsim 770 \text{ GeV}$

But, there are also gluon partners



gluon partner coupling

$$i\Sigma(p) = \frac{16\pi}{3} \alpha_G \int \frac{d^4k}{(2\pi)^4} \gamma^\mu \frac{1}{\not{k} - m_Q} \gamma^\nu \frac{\eta_{\mu\nu}}{(k-p)^2 - m_G^2}$$

where $\alpha_G \simeq \frac{g_3^2}{4\pi} \ln \left(\frac{\Lambda_{UV}}{\Lambda_{IR}} \right) \approx 3$

gluon partner mass

Contribution to Higgs mass:

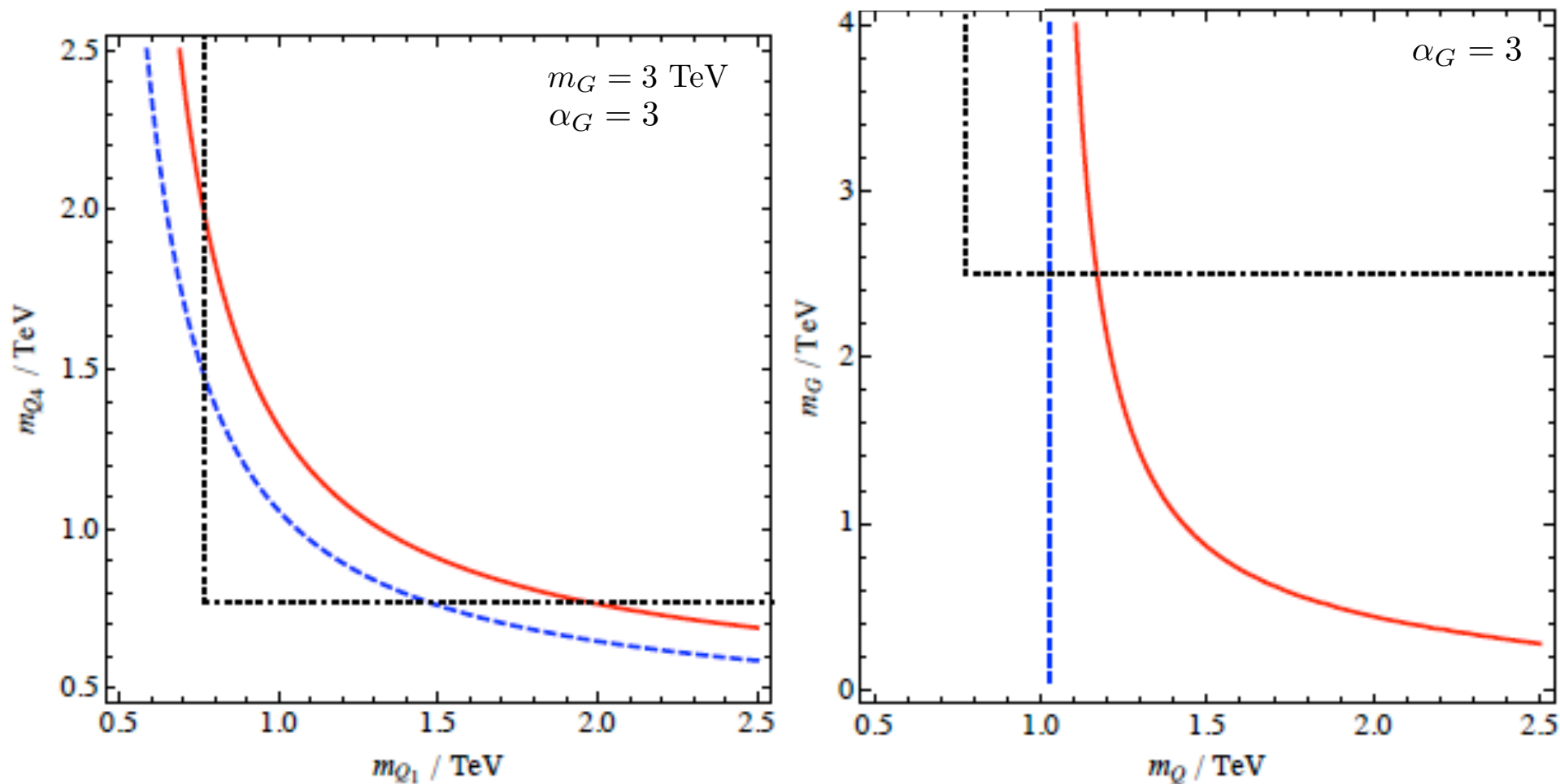
$$m_h^2 \simeq \frac{8N_c v^2}{f^4} \int \frac{d^4 p_E}{(2\pi)^4} \left[\frac{1}{p_E^2} |M(p_E^2)|^2 + \frac{1}{4} \Pi_L^h(p_E^2)^2 + \Pi_R^h(p_E^2)^2 \right]$$

$M, \Pi_{L,R}^h = 2\text{-point functions}$

→ Gluon partner correction to Higgs mass is negative

Contours of $m_h = 126$ GeV

[Barnard, TG, Medina, Sankar Ray 1307.4778]



For $\alpha_G = 3$ and $m_G = 3 \text{ TeV}$ obtain $\approx 10\%$ correction



Eases tension of having light top partners!

UV description of Composite Higgs models

1. Where does global symmetry and spontaneous breaking come from?
2. How do you get a partially composite top?

Possible approach

- AdS/CFT -- D-brane engineering
- Supersymmetric (e.g. Seiberg duality)

[Caracciolo, Parolini, Serone 1211.7290]

} Involves scalars

Look for description *without* elementary scalars...

UV description of SO(6)/SO(5) model [Barnard, TG, Sankar Ray 1311.6562]

$$SO(6)/SO(5) \sim SU(4)/Sp(4) = \underbrace{\mathbf{2}}_{\text{Higgs doublet}} \text{ of } SU(2)_L + \mathbf{1} \text{ singlet}$$

[See also Ferretti, Karateev 1312.5330]
[Ferretti 1404.7137]

Introduce new strong gauge group $Sp(2N_c)$
with 4 Weyl fermion flavors ψ^a ($a = 1, \dots, 4$) \rightarrow $SU(4)$ global symmetry

SU(4) gauged NJL model $\mathcal{L}_{\text{int}} = \frac{\kappa_A}{2N_c} (\psi^a \psi^b)(\bar{\psi}_a \bar{\psi}_b) + \frac{\kappa_B}{8N_c} [\epsilon_{abcd} (\psi^a \psi^b)(\psi^c \psi^d) + \text{h.c.}]$

For $\xi = \frac{1}{4\pi^2} (\kappa_A + \kappa_B) \Lambda^2 > 1$ $SU(4) \rightarrow Sp(4)$ with $\dim \psi\psi \gtrsim 1$

Large anomalous dimension

Top partners $(\psi^a \chi \psi^b)$ or $(\psi^a \tilde{\chi} \psi^b)$

$$\mathcal{L} = \lambda_L t_L \mathcal{O}_R + \lambda_R t_R \mathcal{O}_L$$

$$\dim \mathcal{O}_{L,R} = \dim \psi \chi \psi \approx \dim \psi \psi + \frac{3}{2} \gtrsim \frac{5}{2}$$

	$Sp(2N_c)$	$SU(4)$	$SU(3)_c \times U(1)$
ψ	\square	$\mathbf{4}$	$\mathbf{1}_0$
χ	\boxplus	$\mathbf{1}$	$\mathbf{3}_{+2/3}$
$\tilde{\chi}$	\boxminus	$\mathbf{1}$	$\mathbf{\bar{3}}_{-2/3}$

\rightarrow Allows for order-one top Yukawa coupling

$\xi \gg \underbrace{\sqrt{\alpha}}_{\text{Sp}(2N_c) \text{ gauge coupling}}$ \rightarrow Top partners are naturally lighter than uncolored partners!

Sp(2Nc) gauge coupling

2. Elementary Higgs

Supersymmetric Standard Model

EWWSB $V(h) = -\mu_h^2 |H|^2 + \lambda_h |H|^4$ ($m_A \gg m_Z, \tan \beta \gg 1$)

SUSY {

want large $m_{\tilde{t}}$ [Haber, Hempfling '91] [Ellis, Ridolfi, Zwirner '91]

$$\lambda_h = \frac{1}{4}(g^2 + g'^2) + \frac{3}{32\pi^2} y_t^4 \left[\ln \frac{m_{\tilde{t}}^2}{m_t^2} + \frac{X_t^2}{m_{\tilde{t}}^2} \left(1 - \frac{X_t^2}{12m_{\tilde{t}}^2} \right) \right]$$

where $X_t = A_t - \mu \cos \beta$

want small $m_{\tilde{t}}$

$$\mu_h^2 \simeq |\mu|^2 - \frac{3y_t^2}{4\pi^2} m_{\tilde{t}}^2 \log \frac{\Lambda_{mess}}{m_{\tilde{t}}} \quad (\tan \beta \gg 1)$$

$\lambda_h \simeq 0.13$ ➡ $m_{\tilde{t}} \gtrsim 1 \text{ TeV}$ ($X_t \sim m_{\tilde{t}}$)

$\mu_h^2 \simeq (89 \text{ GeV})^2$ ➡ $m_{\tilde{t}} \lesssim 400 \text{ GeV}$ (“natural”)

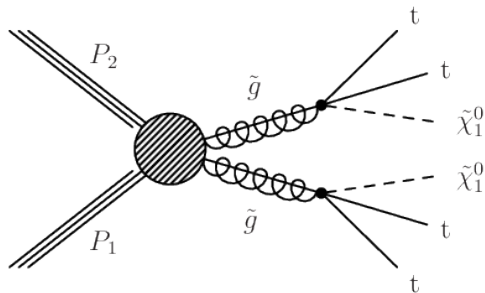
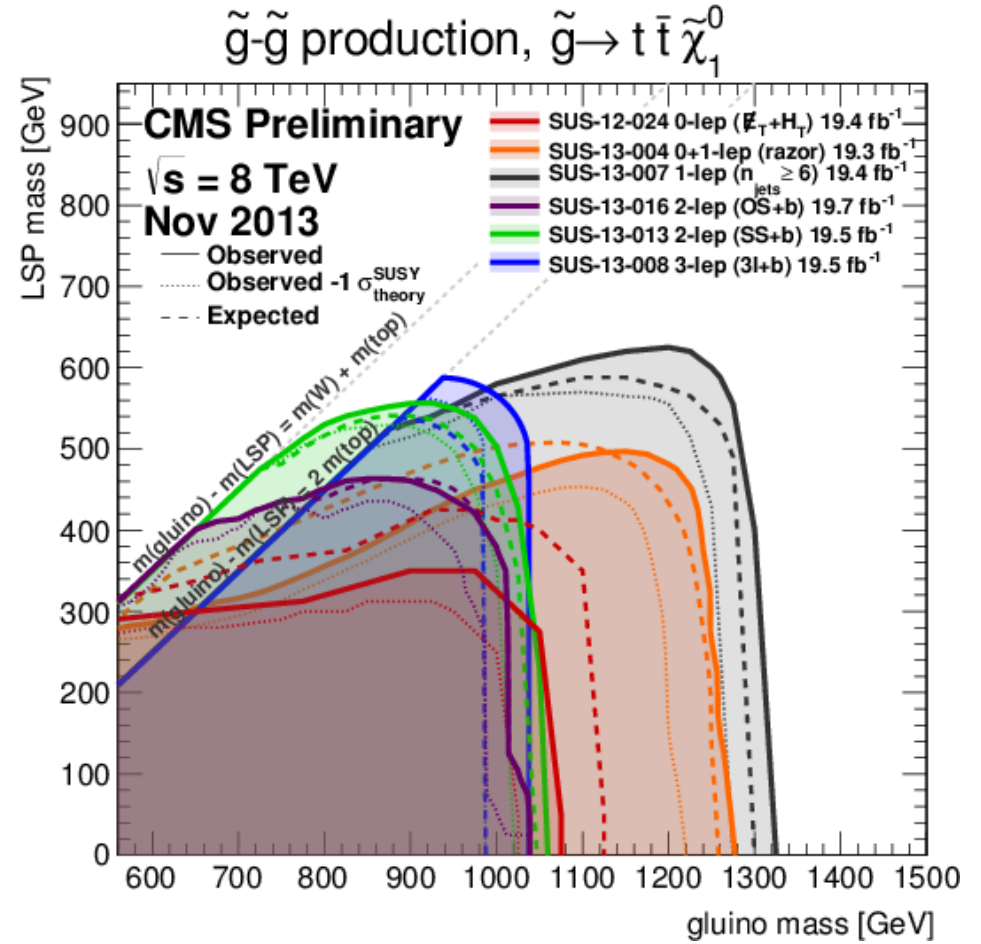
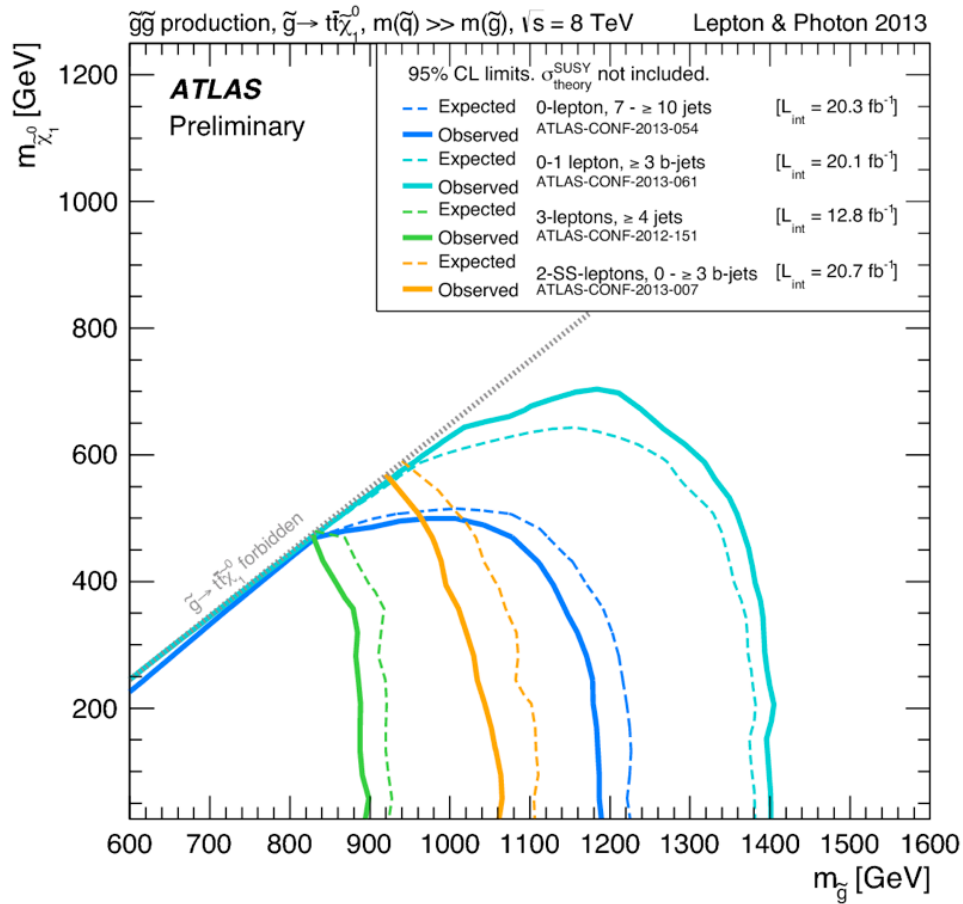
⚠

Also: $\delta m_{\tilde{t}}^2 = \frac{2g_s^2}{3\pi^2} m_{\tilde{g}}^2 \log \frac{\Lambda_{mess}}{m_{\tilde{g}}}$ ➡ $m_{\tilde{g}} \lesssim 2m_{\tilde{t}}$

Natural SUSY = $m_{\tilde{t}_{1,2}}, m_{\tilde{b}_L} \lesssim 500 - 700 \text{ GeV}$
(minimal requirement from EWSB) $m_{\tilde{g}} \lesssim 900 \text{ GeV} - 1.5 \text{ TeV}$
 $\mu \lesssim 200 - 350 \text{ GeV}$
 All other superparticles are heavy $\gtrsim 2 \text{ TeV}$

[Kats, Meade, Reece, Shih '11; Brust, Katz, Lawrence, Sundrum '11; Essig, Izaguirre, Kaplan, Wacker '11; Papucci, Ruderman, Weiler '11]

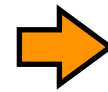
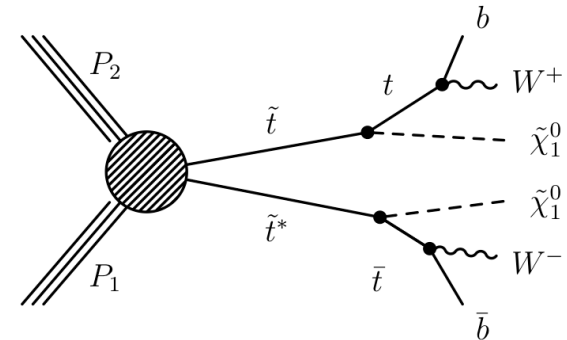
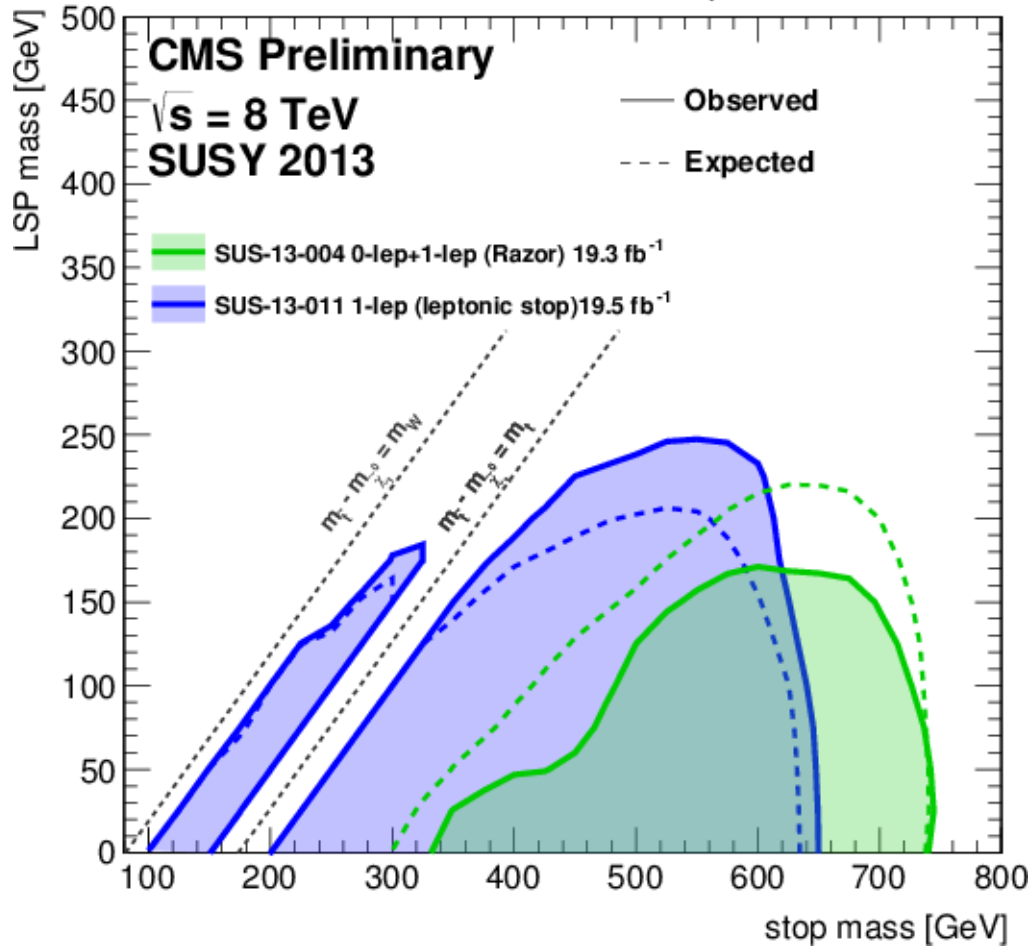
Natural SUSY LHC limits:



$$m_{\tilde{g}} \gtrsim 1400 \text{ GeV}$$

Direct stop production

$\tilde{t}\text{-}\tilde{t}$ production, $\tilde{t} \rightarrow t \tilde{\chi}_1^0$

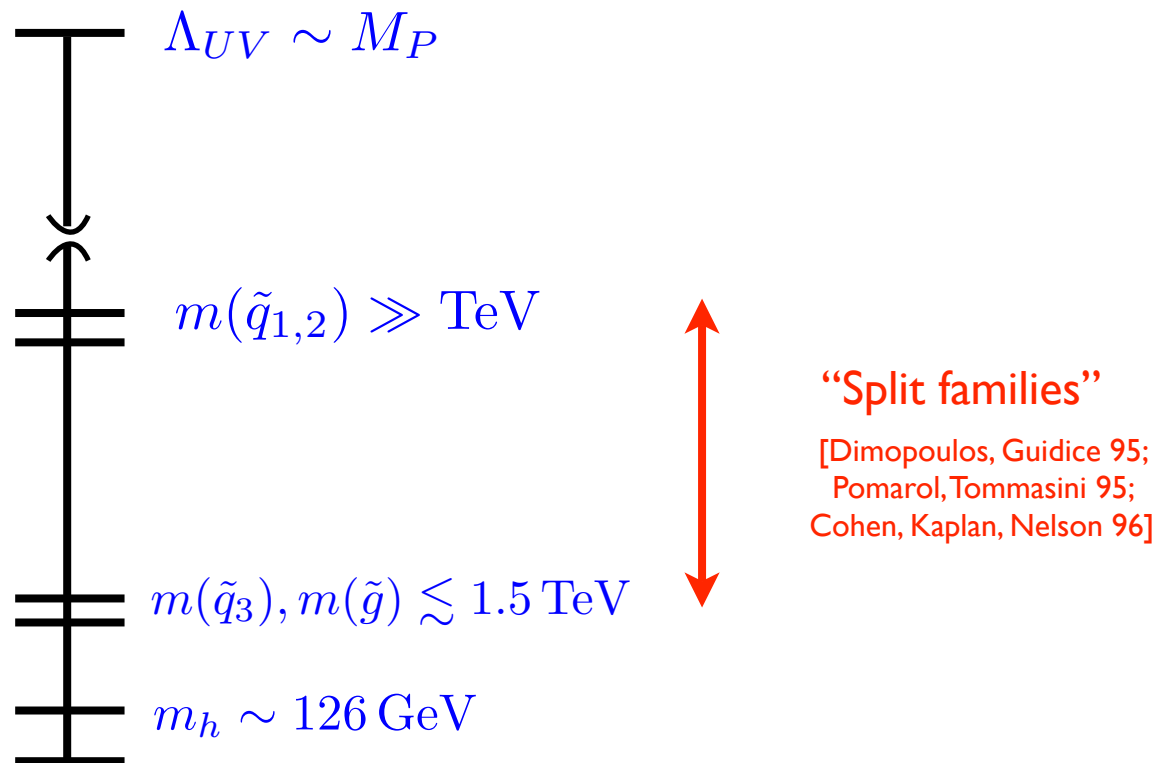


$$m_{\tilde{t}_1} \gtrsim 650 - 740 \text{ GeV}$$

A consistent *natural* SUSY scenario based on Run1:

(i) weakly-coupled Higgs (~ 126 GeV)

(ii) $m(\tilde{q}_{1,2}) \gg m(\tilde{g}), m(\tilde{q}_3)$ SUSY breaking is **flavour dependent!**



However, *natural SUSY* requires a **new** contribution to Higgs quartic coupling:

$$\lambda_h \rightarrow \underbrace{\lambda_h^{(min)}}_{\substack{\text{difficult to naturally accommodate} \\ m_h \sim 126 \text{ GeV}}} + \delta\lambda_h$$

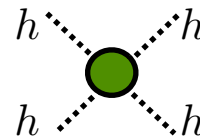
Possibilities:

(i) **NMSSM** $W = \lambda S H_u H_d$ \Rightarrow $\delta\lambda_h \simeq \frac{4\lambda^2}{\tan^2 \beta}$

(ii) **DMSSM** Extra gauge group \Rightarrow $\delta\lambda_h \simeq g^2(1 + \Delta)$
 $\Delta \sim \mathcal{O}(1)$

(iii) **Strong dynamics** $\Delta\mathcal{L} = \xi H \mathcal{O}$

[TG, Pomarol | 107.4697]



\Rightarrow $\delta\lambda_h \simeq \epsilon^4 \frac{16\pi^2}{N}$

How natural is “*natural SUSY*”?

To minimize tuning:

(i) Low messenger scale $\Lambda_{mess} = 20 \text{ TeV}$

$$\log \frac{\Lambda_{mess}}{m_{\tilde{t}}} \sim 3$$

(ii) Add new contribution to Higgs quartic coupling

No need for heavy stop, A-term



(scale-invariant) NMSSM

$$W_{\text{NMSSM}} = \lambda S H_u H_d + \frac{\kappa}{3} S^3 \quad S = \text{singlet}$$

$$V = (m_{H_d}^2 + \lambda^2 S^2) |H_d|^2 + (m_{H_u}^2 + \lambda^2 S^2) |H_u|^2 + \lambda^2 |H_d H_u|^2 + m_S^2 |S|^2 + \kappa^2 |S|^4 \\ + [(a_\lambda S + \lambda \kappa S^2) H_u H_d + \frac{a_\kappa}{3} S^3 + h.c.] + V_D$$

Higgs mass: $m_h^2 = m_Z^2 \cos^2 2\beta + \lambda^2 v^2 \sin^2 2\beta$

Naturalness constraints

[Agashe, Cui, Franceschini 1209.2115]
 [TG, von Harling, Medina, Schmidt 1212.5243]

(i) Electroweak VEV tuning

$$\lambda^2 v^2 = 2 \frac{(a_\lambda v_S + \lambda \kappa v_S^2)}{\sin 2\beta} - \hat{m}_{H_u}^2 - \hat{m}_{H_d}^2 - 2\lambda^2 v_S^2$$

where $\hat{m}_{H_{u,d}}^2 \equiv m_{H_{u,d}}^2 + \frac{d}{dv_{u,d}^2} V_1$

 Large λ helps

(ii) Higgs mass tuning $m_h^2 = m_Z^2 \cos^2 2\beta + \lambda^2 v^2 \sin^2 2\beta + \delta m_{h,mix}^2$

$$\delta m_{h,mix}^2 \simeq -\lambda^2 v^2 \frac{\left(\frac{\lambda}{\kappa} - \sin 2\beta \left[1 + \frac{a_\lambda}{2\lambda \kappa v_S} \right] \right)^2}{1 + \frac{a_\kappa}{4\kappa^2 v_S} + \frac{a_\lambda \sin 2\beta v^2}{8\kappa^2 v_S^3}}$$

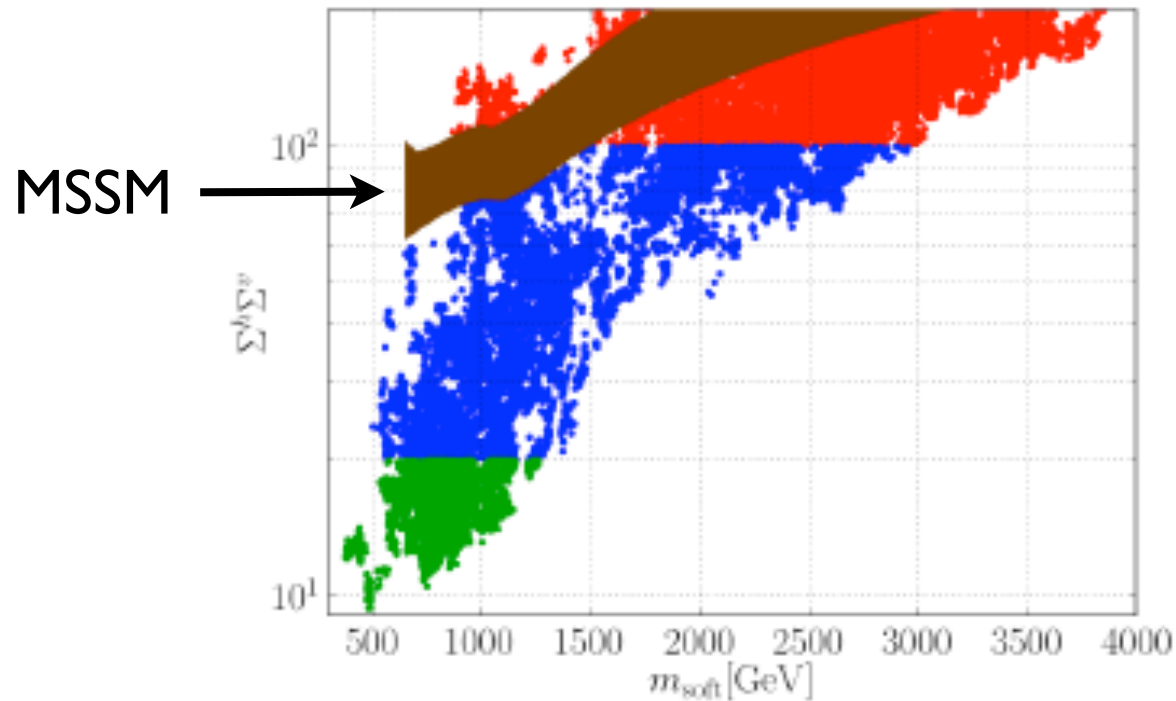
 Large λ hurts

Result: Optimal value $\lambda \sim 1$

Combined tuning $\Sigma^v \Sigma^h$

where $\Sigma^v \equiv \max_i \left| \frac{d \log v^2}{d \log \xi_i(\Lambda_{\text{mess}})} \right|$ $\Sigma^h \equiv \max_{\xi_i} \left| \frac{d \log m_h^2}{d \log \xi_i} \right|$

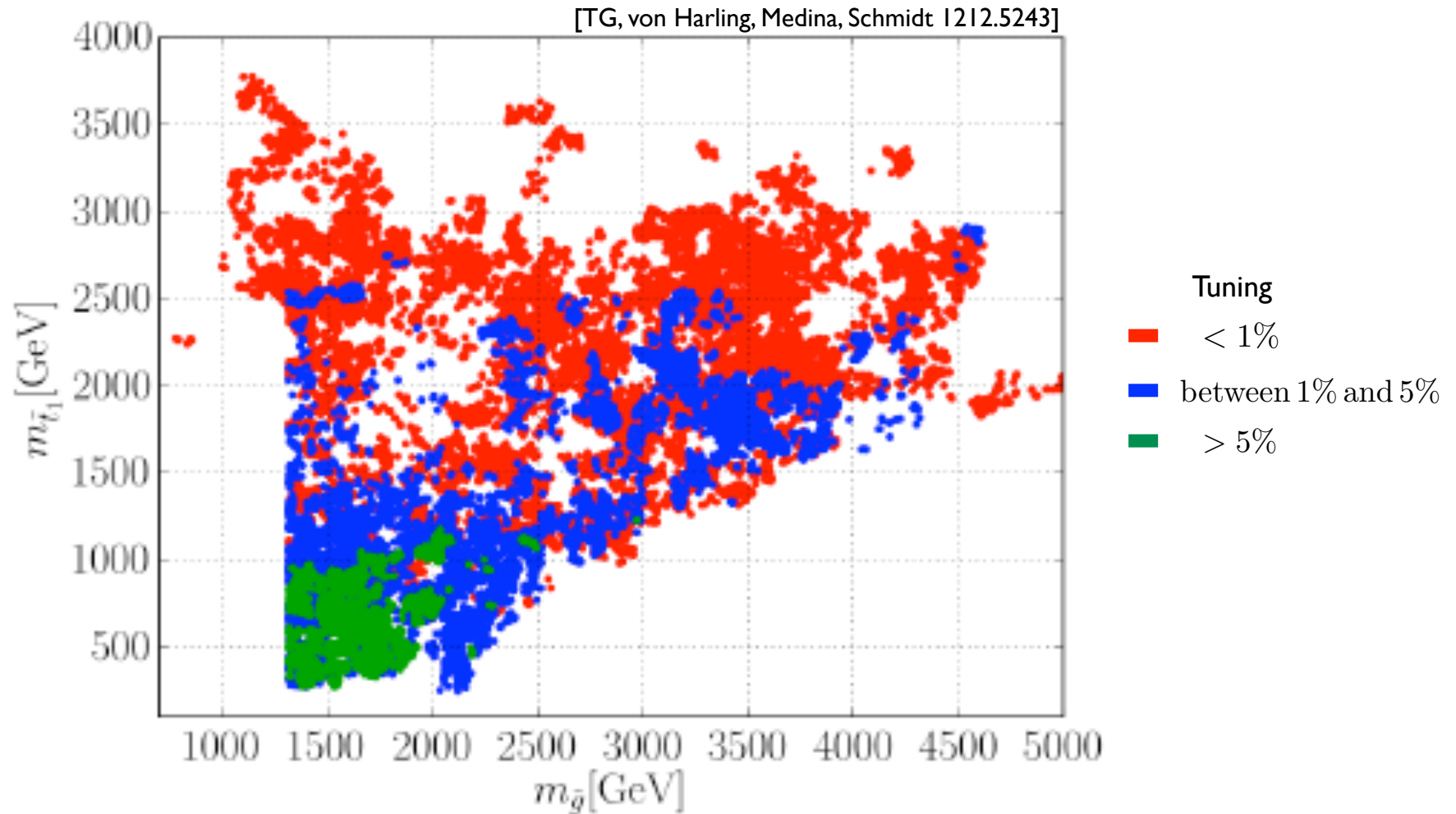
[TG, von Harling, Medina, Schmidt 1212.5243]



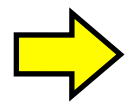
LHC Run I

➔ *natural* SUSY ($\sim 20\%$) has become $< 10\%$ tuning

Glino-stop masses



For tuning $\sim 5 - 10\%$



$$m_{\tilde{g}} \lesssim 2 \text{ TeV}$$

$$m_{\tilde{t}_1} \lesssim 1.2 \text{ TeV}$$

Caveat: Bottom-up approach is naive sampling of parameter space

 Tuning is probably worse!

Important question:

What about explicit supersymmetry breaking models?

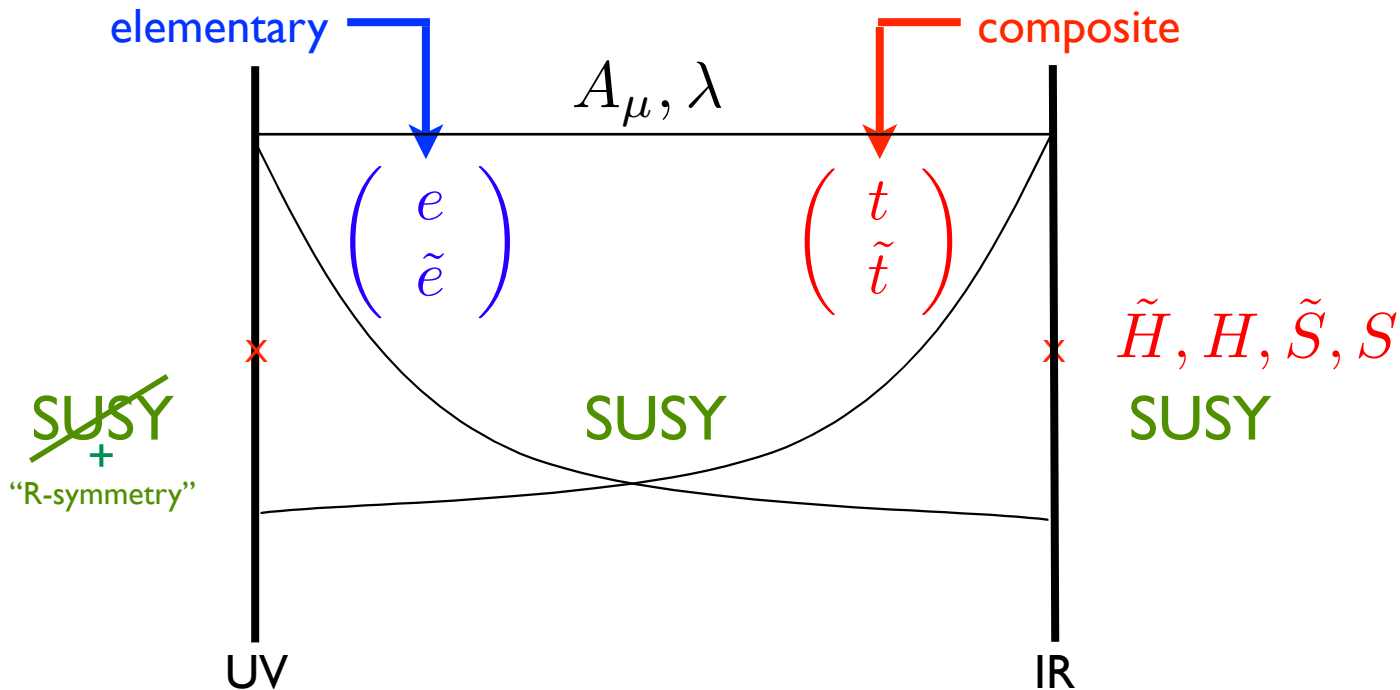
- Low messenger scale
- Split family spectrum
- Dark matter
- Gauge coupling unification

A 5D model: “Accidental SUSY”

SUSY broken at UV scale!

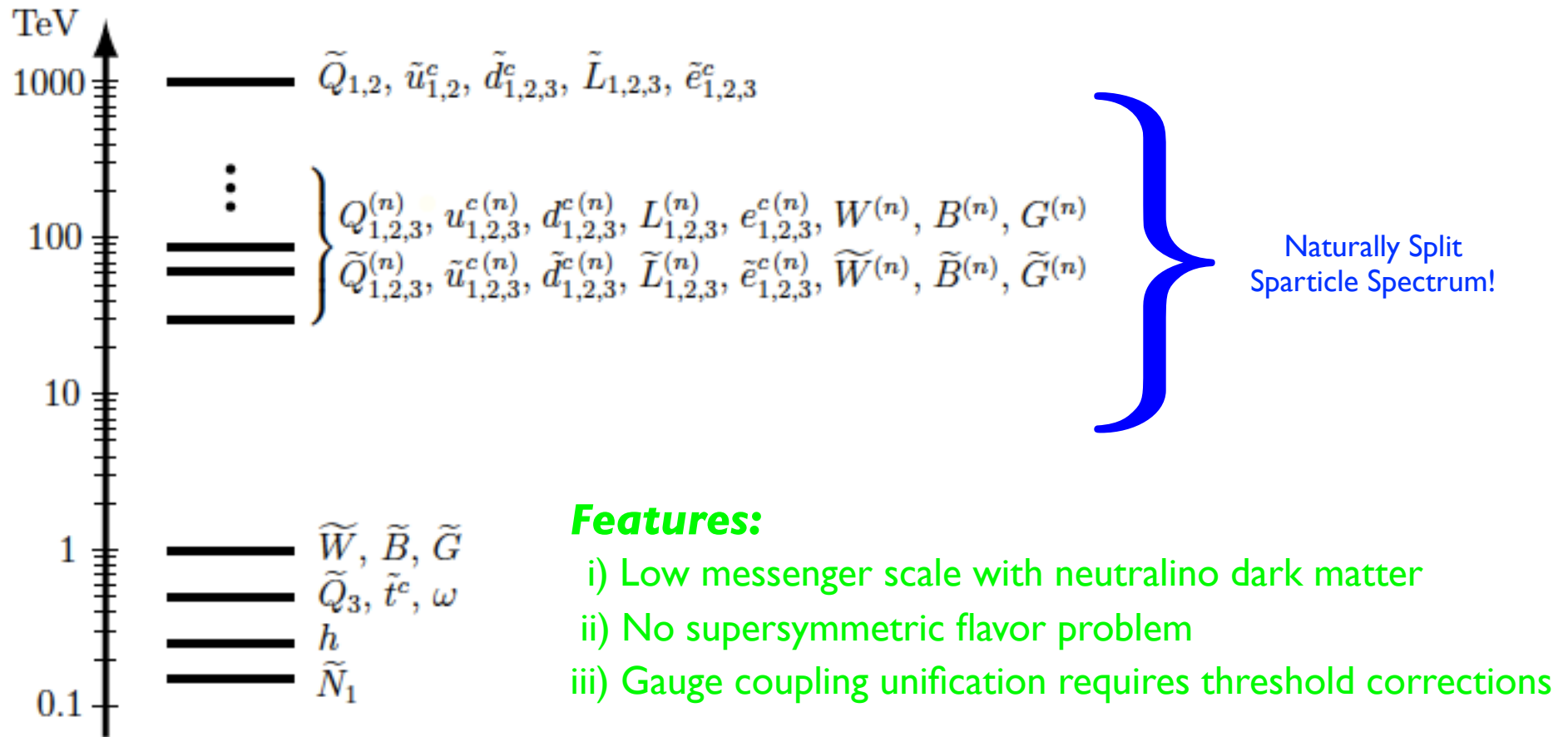
[TG, Pomarol '03; Sundrum '09]

[TG, von Harling, Setzer arXiv:1104.3171]



Fermion mass spectrum determines **sfermion** spectrum!

Particle spectrum:



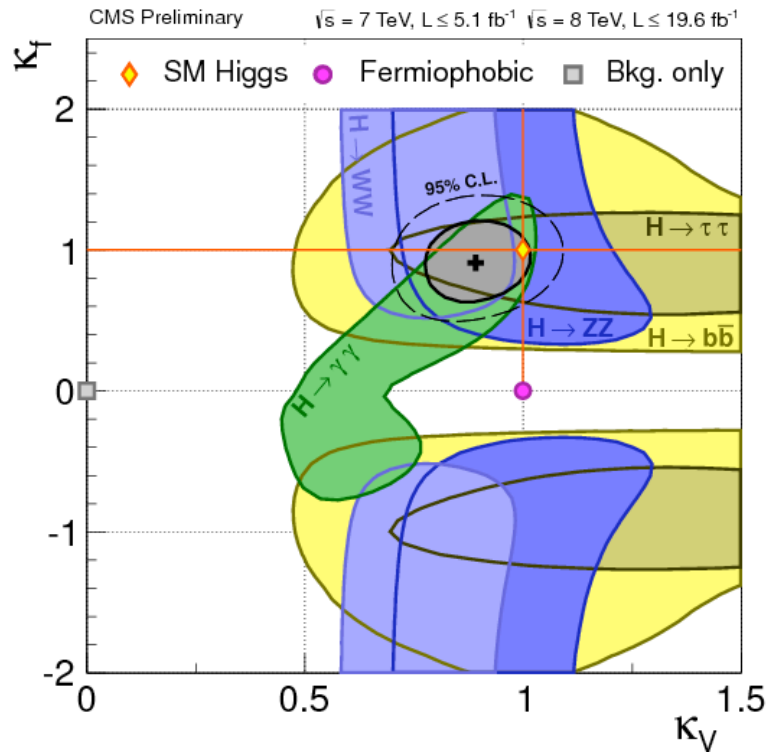
Other possibilities: Dirac gauginos, RPV, Susy Twin Higgs,....

Natural SUSY model building worth exploring!

What about Higgs couplings?

[TG, von Harling, Medina, Schmidt | 401.8291]

[see also Farina, Perelstein, Shakya | 310.0459]



Define

$$r_u \equiv \frac{\text{Higgs coupling to up-type fermions}}{\text{SM Higgs coupling to up-type fermions}} \quad \text{similarly } r_d, r_V$$

where for NMSSM

$$r_u \simeq 1 - \frac{m_{hH}^4}{2m_H^4} - \frac{m_{hs}^4}{2m_s^4} + \cot \beta \left(\frac{m_{hH}^2}{m_H^2} - \frac{m_{hs}^2 m_{Hs}^2}{m_H^2 m_s^2} + \frac{m_h^2 m_{hH}^2}{m_H^4} \right)$$

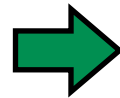
$$r_d \simeq 1 - \frac{m_{hH}^4}{2m_H^4} - \frac{m_{hs}^4}{2m_s^4} - \tan \beta \left(\frac{m_{hH}^2}{m_H^2} - \frac{m_{hs}^2 m_{Hs}^2}{m_H^2 m_s^2} + \frac{m_h^2 m_{hH}^2}{m_H^4} \right)$$

$$r_V \simeq 1 - \frac{m_{hH}^4}{2m_H^4} - \frac{m_{hs}^4}{2m_s^4}$$

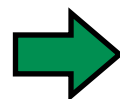
with CP even mass-squared matrix

$$\mathcal{M}^2 = \begin{pmatrix} \overset{\text{SM Higgs}}{m_h^2} & m_{hH}^2 & m_{hs}^2 \\ m_{hH}^2 & m_H^2 & m_{Hs}^2 \\ m_{hs}^2 & m_{Hs}^2 & m_s^2 \end{pmatrix}$$

SM-like couplings



$$m_{hs}, m_{hH} \ll m_h \ll m_H, m_s$$



new source of tuning

NMSSM

$$W \supset \lambda S H_u H_d + \kappa S^3$$

$$V = (m_{H_u}^2 + \lambda^2 |S|^2) |H_u^0|^2 + (m_{H_d}^2 + \lambda^2 |S|^2) |H_d^0|^2 + \lambda^2 |H_u^0 H_d^0|^2 + m_S^2 |S|^2 + \kappa^2 |S|^4 + \left[\frac{a_\kappa}{3} S^3 - (a_\lambda S + \lambda \kappa S^2) H_u^0 H_d^0 + \text{h.c.} \right] + \tilde{g}^2 (|H_u^0|^2 - |H_d^0|^2)^2$$

with $m_h^2 = 2\tilde{g}^2 v^2 \cos^2 2\beta + \frac{1}{2} \lambda^2 v^2 \sin^2 2\beta$

$$m_H^2 = \csc 2\beta \left(\sqrt{2} v_s a_\lambda + \kappa \lambda v_s^2 - \frac{v^2}{2} \sin^3 2\beta (\lambda^2 - 4\tilde{g}^2) \right)$$

$$m_s^2 = \frac{a_\kappa v_s}{\sqrt{2}} + 2\kappa^2 v_s^2 + \frac{a_\lambda v^2 \sin 2\beta}{\sqrt{8} v_s}$$

$$m_{h_s}^2 = v \left(\lambda^2 v_s - \sin \beta \cos \beta (\sqrt{2} a_\lambda + 2\kappa \lambda v_s) \right)$$

$$m_{hH}^2 = \frac{v^2}{4} \sin 4\beta (4\tilde{g}^2 - \lambda^2)$$

$$m_{H_s}^2 = \frac{v}{2} \cos 2\beta (\sqrt{2} a_\lambda + 2\kappa \lambda v_s)$$

Tuning: $\Sigma \gtrsim \frac{1}{4} (4\tilde{g}^2 - \lambda^2) f(\lambda, \kappa, \tan \beta, \tilde{g}) \cdot \begin{cases} \frac{\cot \beta \sin 4\beta}{|r_u - 1|} \\ \frac{\tan \beta \sin 4\beta}{|r_d - 1|} \end{cases}$

$$\Sigma \gtrsim \frac{1}{4} (4\tilde{g}^2 - \lambda^2) f(\lambda, \kappa, \tan \beta, \tilde{g}) \frac{\sin 4\beta}{\sqrt{2(1 - r_V)}}$$

Fine-tuning
grows as
 $r_{u,d,V} \rightarrow 1$

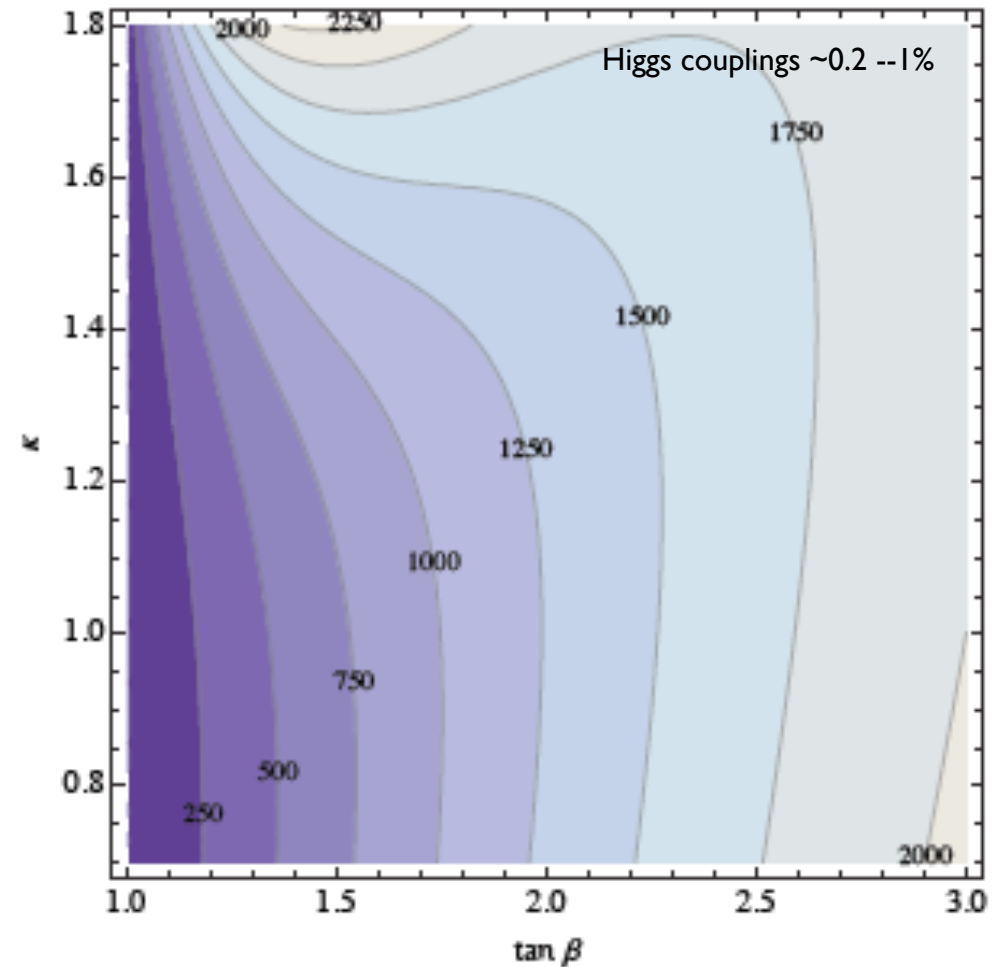
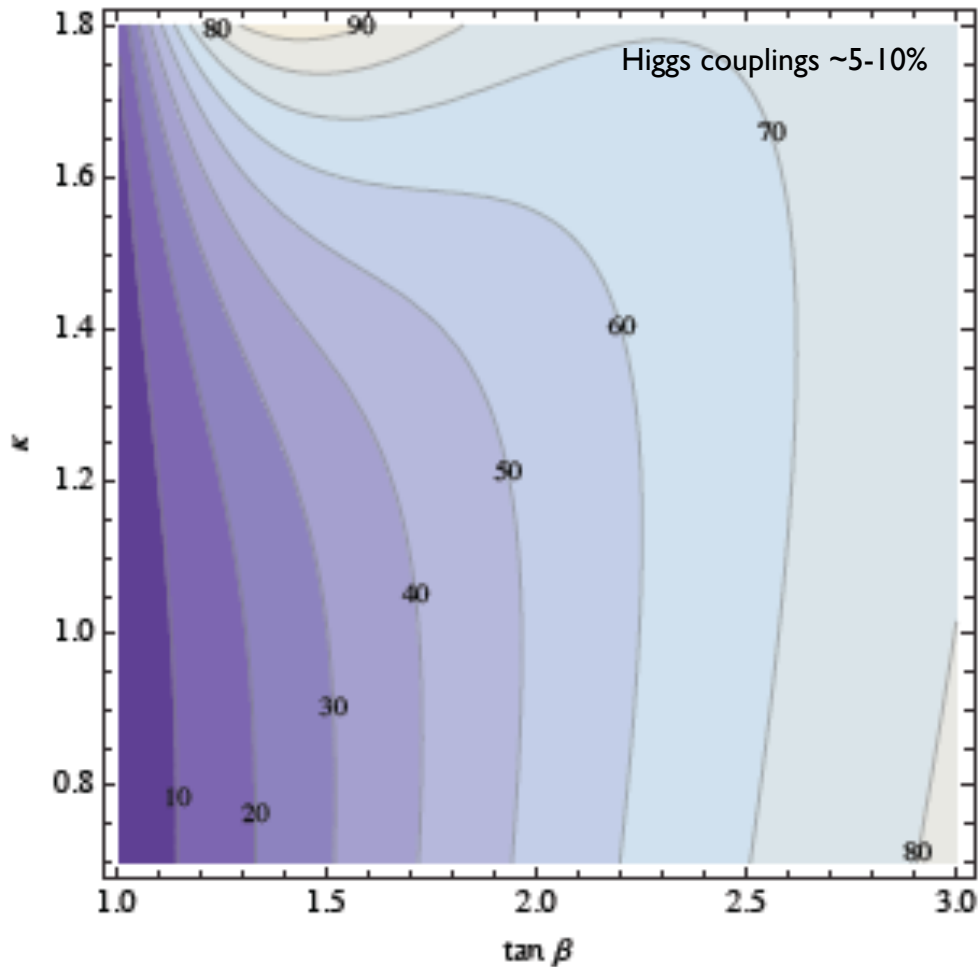
Note: EWPT $\rightarrow \tan \beta \lesssim 4$ So tuning cannot be offset with large $\tan \beta$

Contours of fine tuning measure Σ

[TG, von Harling, Medina, Schmidt 1401.8291]

LHC 14 TeV 300 fb⁻¹

ILC 1TeV 2500 fb⁻¹



Conclusion

- Naturalness not yet ruled out
 - Composite Higgs: tuning $\lesssim 10\%$
 - Natural SUSY: tuning $\lesssim 5\%$
- SO(6)/SO(5) composite Higgs model has a simple UV description
- Natural SUSY models require some elaborate structure
- A resonance or superpartner could still show up at Run II! If not, SM-like Higgs couplings will further test naturalness at ILC