

Cosmodynamics

**Quintessence and solution of
cosmological constant
problem should be related !**

Cosmon and fundamental mass scales

- Assume all mass parameters are proportional to scalar field χ (GUTs, superstrings,...)
- $M_p \sim \chi$, $m_{\text{proton}} \sim \chi$, $\Lambda_{\text{QCD}} \sim \chi$, $M_W \sim \chi$, ...
- χ may evolve with time
- m_n/M : (almost) constant - observation!

Only ratios of mass scales are observable

Dirac's hypothesis

The very small dimensionless numbers in physics are due to huge age of the universe .

(A)
may be true for
ratio of dark energy
over M^4

(B) nucleon mass $m_n / \bar{M}_p \approx 10^{-18}$

$$\left| \frac{d}{dt} \ln \left(\frac{m_n}{\bar{M}_p} \right) \right| < 3 \cdot 10^{-15} \text{ yr}^{-1}$$

from Oklo + GUT

less stringent direct bounds

\Rightarrow today essentially settled !

(perhaps more substantial evolution in very early cosmology)

Dilatation symmetry

- Lagrange density:

$$L = \sqrt{g} \left(-\frac{1}{2} \chi^2 R + \frac{1}{2} (\delta - 6) \partial^\mu \chi \partial_\mu \chi + V(\chi) + h \chi \bar{\psi} \psi \right)$$

- Dilatation symmetry for

$$V = \lambda \chi^4, \lambda = \text{const.}, \delta = \text{const.}, h = \text{const.}$$

- Conformal symmetry for $\delta=0$

Dilatation anomaly

- Quantum fluctuations responsible for dilatation anomaly
- Running couplings:

$$\partial\lambda/\partial\ln\chi = -A\lambda, \quad \partial\delta/\partial\ln\chi = E\delta^2$$

- $V \sim \chi^{4-A}$, $M_p(\chi) \sim \chi$
- $V/M_p^4 \sim \chi^{-A}$: decreases for increasing χ !!
- $E > 0$: crossover quintessence

Cosmology

Cosmology : χ increases with time !

(due to coupling of χ to curvature scalar)

“ late time cosmology explores the
ultraviolet ”

for large χ the ratio V/M^4 decreases to zero



Effective cosmological constant vanishes
asymptotically for large t !

Weyl scaling

Weyl scaling : $g_{\mu\nu} \rightarrow (M/\chi)^2 g_{\mu\nu}$,
 $\varphi/M = \ln (\chi^4/V(\chi))$

$$L = \sqrt{g} \left(-\frac{1}{2} M^2 R + \frac{1}{2} k^2(\phi) \partial^\mu \phi \partial_\mu \phi \right. \\ \left. + V(\phi) + m(\phi) \bar{\psi} \psi \right)$$

Exponential potential : $V = M^4 \exp(-\varphi/M)$

No additional constant !

Crossover Quintessence

$$\partial\delta/\partial\ln\chi = E\delta^2 \quad (\text{like QCD gauge coupling})$$

critical χ where δ grows large

critical φ where k grows large

$$k^2(\varphi) = \delta(\chi)/4$$

$$k^2(\varphi) = "1/(2E(\varphi_c - \varphi)/M)"$$

if $\varphi_c \approx 276/M$ (tuning!)

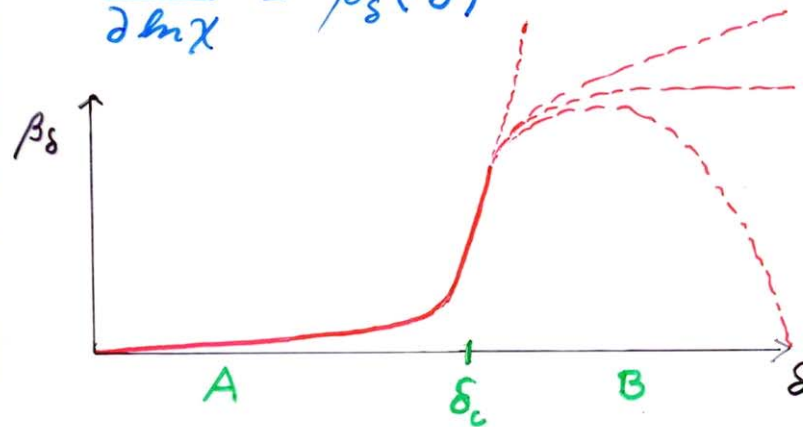
Relative increase of dark energy in present
cosmological epoch

Running dimensionless coupling $\delta(\chi/m)$

local conformal symmetry for $\delta=0, m=0$

for $\chi \gg m$:

$$\frac{\partial \delta}{\partial \ln \chi} = \beta_\delta(\delta)$$



A: vicinity of conformal fixed point at $\delta=0$

B: cosmon + metric interactions are
pure derivative interactions

δ_c : characteristic coupling for crossover

example :

$$\beta_g = E \delta^2 \quad (\delta_c \approx 1/E)$$

$$\delta = \frac{1}{E \ln(\chi_c/\chi)}$$

characteristic crossover scale χ_c

$$\text{define } \delta_i = \delta(\chi = m)$$

$$\Rightarrow \frac{\chi_c}{m} = \exp \frac{1}{E \delta_i}$$

(„Landau pole“)

$$\frac{\chi_c}{m} \gg 1$$

$$\text{for } \delta_i \approx \frac{1}{138 E} :$$

\Rightarrow realistic cosmology

scenario , where

a) „cosmological constant“ vanishes for $t \rightarrow \infty$

$$V \sim \exp - \frac{\Phi}{M_p} , \text{ not } \exp - \frac{\Phi}{M_p} + C$$

b) before present epoch :

δ varies slowly

$$\Omega_h \approx \begin{cases} \delta & \text{radiation dom.} \\ \frac{3}{4} \delta & \text{matter dom.} \end{cases}$$

Ω_h small, but not tiny in the past !

c) „today“ linked to small number
in fundamental theory

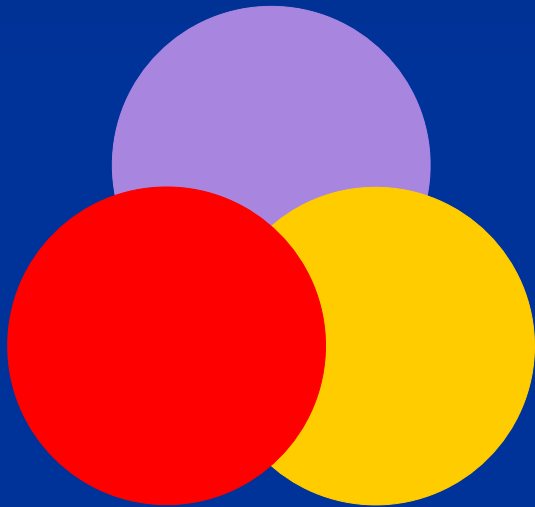
$$E\delta_i \approx 1 / 138$$

One could hope for explanation of
„cosmic coincidences“ that
relate H_0/M_p to fundamental parameters

$$\frac{H_0^2}{M_p^2} \approx \frac{1}{3} \exp - \frac{1}{\delta_i E}$$

“Fundamental” Interactions

Strong, electromagnetic, weak
interactions



gravitation

cosmodynamics

On astronomical
length scales:

graviton

+

cosmon



Cosmon

- *Tiny mass*
- $m_c \sim H$
- *New long - range interaction*

cosmon mass

for standard kinetic term

$$\blacksquare m_c^2 = V''$$

for standard exponential potential , $k = \text{const.}$

$$\begin{aligned}\blacksquare m_c^2 &= V'' / k^2 = V / (k^2 M^2) \\ &= 3 \Omega_h (1 - w_h) H^2 / (2 k^2)\end{aligned}$$

Are fundamental “constants” time dependent ?

Fine structure constant α (electric charge)

Ratio nucleon mass to Planck mass

“Fifth Force”

- Mediated by scalar field

R.Peccei,J.Sola,C.Wetterich,Phys.Lett.B195,183(1987)

- Coupling strength: weaker than gravity
(nonrenormalizable interactions $\sim M^{-2}$)

- Composition dependence

→ violation of equivalence principle

- Quintessence: connected to time variation of fundamental couplings

C.Wetterich , Nucl.Phys.B302,645(1988)

Quintessence and Time dependence of “fundamental constants”

- Fine structure constant depends on value of
cosmon field : $\alpha(\varphi)$

*(similar in standard model: couplings depend on
value of Higgs scalar field)*

- Time evolution of φ 
Time evolution of α

Jordan,...

Field dependent gauge coupling (gauge invariance maintained)

$$\mathcal{L}_g = \frac{1}{4} Z_F(\chi) F^{\mu\nu} F_{\mu\nu}$$

renormalized gauge coupling

$$g^2(\chi) = \bar{g}^2 Z_F^{-1}(\chi)$$

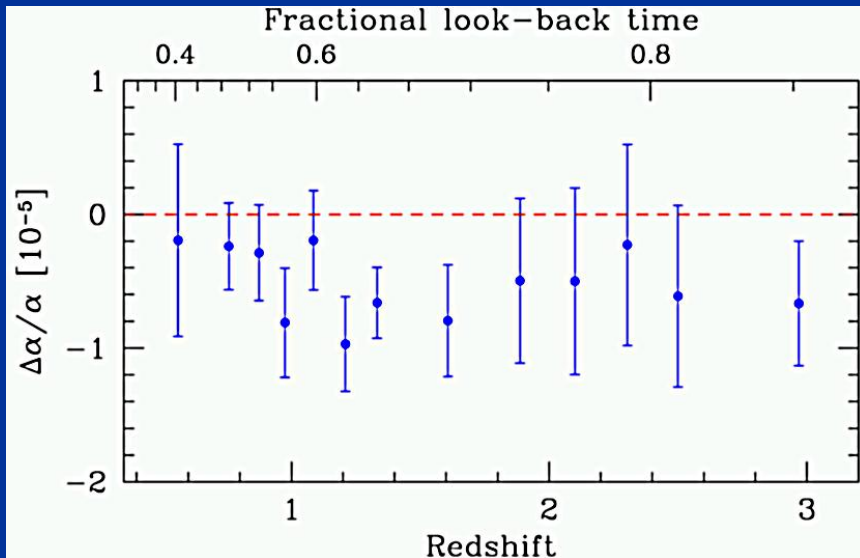
for GUT : C.Hill ; Q.Shafi , CW

GUT : running of electromagnetic and strong gauge coupling related

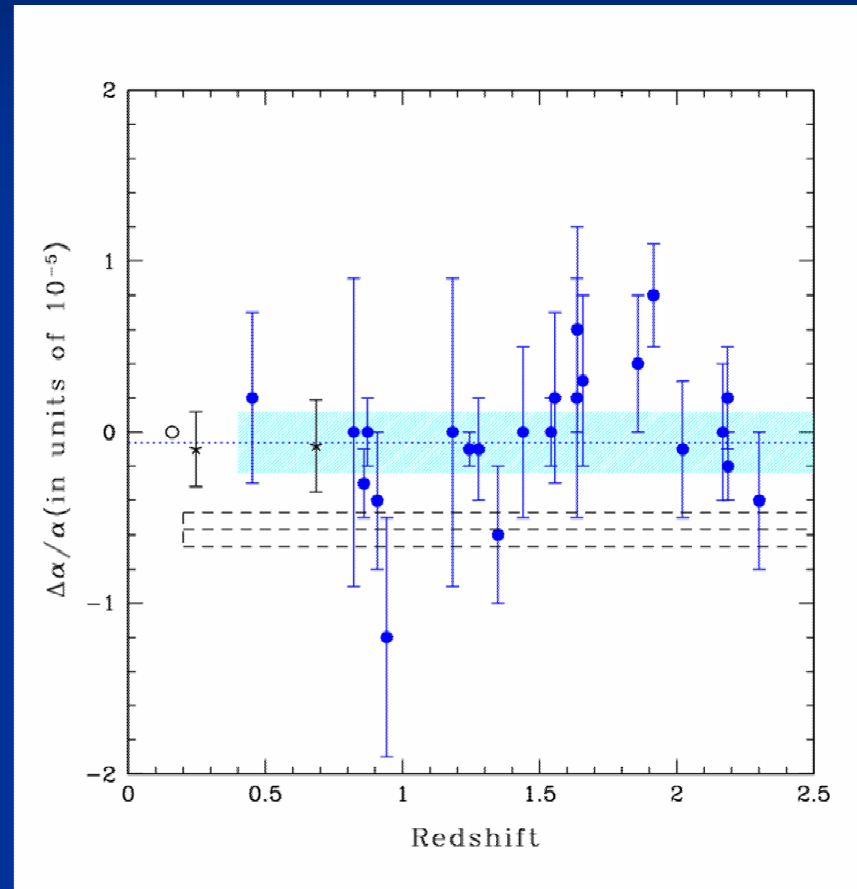
$$\alpha_{em}^{-1}(m_e) = \frac{32\pi Z_F(\chi)}{3\bar{g}^2} + \frac{5}{3\pi} \ln \frac{M_{GUT}}{M_W} \\ + \frac{10}{3\pi} \ln \frac{M_W}{m_n} + \frac{2}{\pi} \ln \frac{m_n}{m_e}$$
$$\frac{\Delta(m_n/M)}{(m_n/M)} = 39.1 \frac{\Delta\alpha_{em}}{\alpha_{em}}$$

strong effect from variation of nucleon mass
for time dependent couplings !

Variation of fine structure constant as function of redshift



Webb et al



Srianand et al

Variation of fine structure constant

Three independent data sets from Keck/HIRES

$$\Delta\alpha/\alpha = -0.54 (12) \cdot 10^{-5}$$

Murphy, Webb, Flammbaum, june 2003

VLT

$$\Delta\alpha/\alpha = -0.06 (6) \cdot 10^{-5}$$

Srianand, Chand, Petitjean, Aracil, feb.2004

$z \approx 2$

Crossover quintessence and time variation of fundamental “constants”

Upper bounds for relative variation of the fine structure constant

- Oklo natural reactor $\Delta\alpha/\alpha < 10^{-7}$ $z=0.13$
- Meteorites (Re-decay) $\Delta\alpha/\alpha < 3 \cdot 10^{-7}$ $z=0.45$
- **Crossover Quintessence leads to small variation of couplings at low z as compared to $z \approx 2$!**

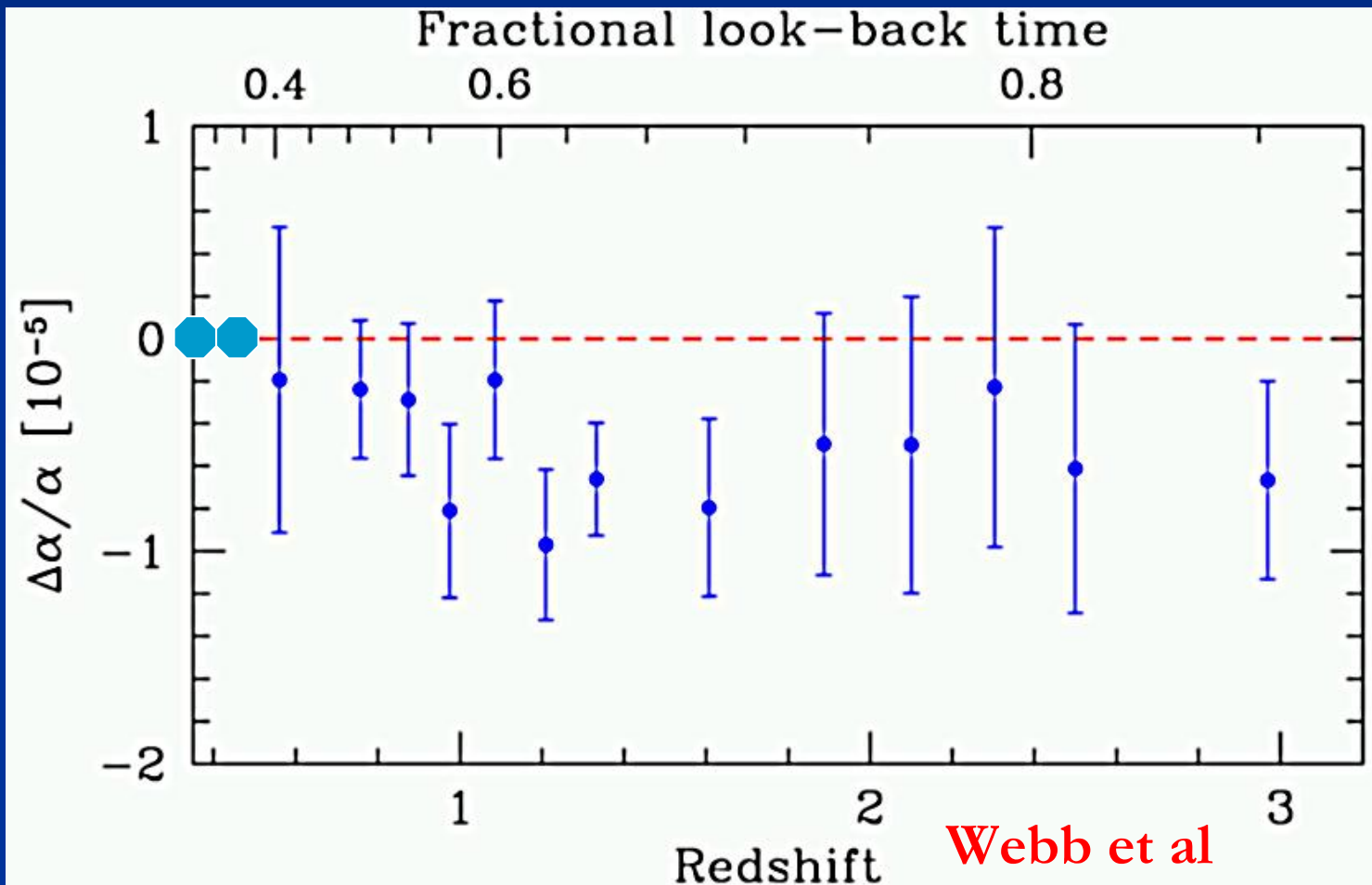
Time evolution of fundamental couplings
traces time evolution of quintessence

today w_h close to -1 :

- Small kinetic energy
- Slow change of ϕ
- Slow change of α

Very small $\Delta\alpha/\alpha$ for low z !

Variation of fine structure constant as function of redshift



Cosmon and time variation of couplings

- small coupling of cosmon to matter due to fixed points behavior

Dimensional transmutation

α_x : unified gauge coupling at unification scale $M_x \sim \chi$

running $\alpha_x(\chi)$:

$$\frac{\partial \alpha_x}{\partial \ln \chi} = b_2 \alpha_x - b_4 \alpha_x^2$$
$$b_2, b_4 > 0$$

IR - scale where α_x grows large : m

UV - fixed point reached for $\chi \gg m$

$$\alpha_{x*} = \frac{b_2}{b_4} \approx \frac{1}{40}$$

close to fixed point :
small time evolution of couplings
coupling to matter weaker
than gravitational strength

dependence of fixed point on δ could induce observable effect

$$\frac{\partial \alpha_x}{\partial \ln \chi} = 0.2 \alpha_x - 8 \alpha_x^2 - b_6 \alpha_x^3 \left(6 - \frac{\delta}{1+6\delta} \right)$$

adjusting b_6 to reproduce results by Webb et al:

$\Delta \alpha_{em} / \alpha_{em}$ for various z

$$z: \begin{array}{|c|c|c|c|} \hline 0.13 & 0.45 & 2 & 10^{10} \\ \hline \frac{\Delta \alpha}{\alpha} & -1.7 \cdot 10^{-7} & -8.2 \cdot 10^{-7} & -7 \cdot 10^{-6} \\ \hline \end{array} \quad \begin{array}{|c|} \hline -7.8 \cdot 10^{-5} \\ \hline \end{array}$$

differential acceleration : $\eta = 4.4 \cdot 10^{-14}$

smaller for Srianand et al !

Realistic example, crossover quintessence

$$E=5 \quad \left(\frac{\partial S}{\partial \ln \chi} = 5\delta^2 \right)$$

$$h = 0.66, \quad \Omega_{\chi}^{(0)} = 0.7$$

$$w_{\chi}^{(0)} = -0.93$$

$$t^{(0)} = 13.7 \cdot 10^9 \text{ yr}$$

$$\Omega_{\chi}^{(1)} = 0.02$$

$$\ell_3 = 796, \quad \sigma_8 / \sigma_8^{(1)} = 0.7$$

Time variation of coupling constants
is tiny –

would be of very high significance !

Possible signal for Quintessence

Πάντα ρει

Cosmodynamics

Cosmon mediates new long-range interaction

Range : size of the Universe – horizon

Strength : weaker than gravity

| | |
|--------|-----------------|
| photon | electrodynamics |
|--------|-----------------|

| | |
|----------|---------|
| graviton | gravity |
|----------|---------|

| | |
|--------|---------------|
| cosmon | cosmodynamics |
|--------|---------------|

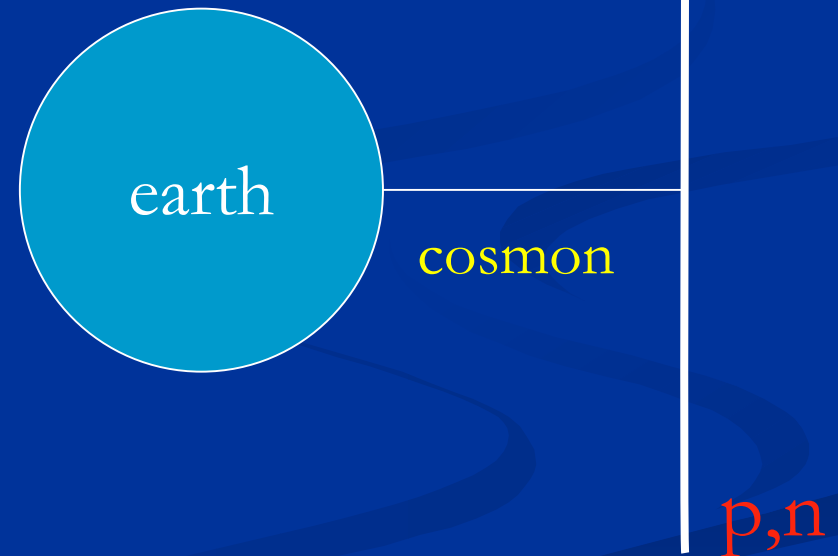
Small correction to Newton's law

Violation of equivalence principle

Different couplings of
cosmon to proton and
neutron

Differential acceleration

Violation of equivalence
principle



Differential acceleration η

For unified theories (GUT) :

$$\eta = -1.75 \cdot 10^{-2} \Delta R_z \left(\frac{\partial \ln \alpha}{\partial z} \right)^2 \frac{1 + \tilde{Q}}{\Omega_h (1 + w_h)}$$

$$\Delta R_z = \frac{\Delta Z}{Z + N} \approx 0.1$$

$$\eta = \Delta a / 2a$$

Q : time dependence of other parameters

Link between time variation of α

and violation of equivalence principle

typically : $\eta = 10^{-14}$

if time variation of α near Oklo upper bound

to be tested by MICROSCOPE

$$(1) \quad \alpha_x(\varphi) \rightarrow \Lambda_{QCD}(\varphi) \rightarrow m_n(\varphi)$$

nucleon mass depends on value
of the cosmon field
(and therefore on time)

(2) expand around cosmological
value $\varphi_0(t)$:

$$\varphi(\vec{x}, t) = \varphi_0(t) + \delta\varphi(\vec{x}, t)$$

$$m_n = m_n(\varphi_0) + \frac{\partial m_n}{\partial \varphi} \Big|_{\varphi_0} \delta\varphi$$

\Rightarrow cosmon - nucleon vertex $\sim \bar{n} n \delta\varphi$



\Rightarrow earth is source for surrounding
local cosmon field $\delta\varphi(|\vec{r}|)$

(3) Test body carries effective
„cosmon charge“

$$Q_c = k^{-1} \frac{\partial m_t}{\partial \varphi}$$

to be compared with „gravitational charge“

$$Q_g = \frac{m_t}{\sqrt{2} M_p}$$

⇒ Correction to Newtonian potential

$$V_N = - \frac{G_N M m_t}{r} (1 + \alpha_t)$$

$$\alpha_t = \frac{2 M_p^2}{k^2} \frac{\partial \ln M}{\partial \varphi} \frac{\partial \ln m_t}{\partial \varphi}$$

(4) Protons and neutrons have different
cosmon charges, $\frac{\partial m_p}{\partial \varphi} \neq \frac{\partial m_n}{\partial \varphi}$

This leads to differential acceleration !

observation

$$|\eta| < 3 \cdot 10^{-13}$$

Baessler et al.

Atomic clocks and OKLO

* Atomic clocks:

$$\frac{\dot{\alpha}_{em}}{\alpha_{em}} = -5.4 \cdot 10^{-10} \frac{\Delta \alpha_{em}}{\alpha_{em}} (z=0.13) \text{ yr}^{-1}$$

$$\text{observation } \frac{\dot{\alpha}_{em}}{\alpha_{em}} = (4.2 \pm 6.9) \cdot 10^{-15} \text{ yr}^{-1}$$

Sortais et al.

assumes that both effects are dominated
by change of fine structure constant

small change of couplings in space

- ★ Fine structure constant depends on location in space (satellites!)

$$r = 2R : \frac{S_{\text{dem}}}{\alpha_{\text{em}}} = 3 \cdot 10^{-19} \text{ } h^{-2}$$

Summary

- $\Omega_h = 0.7$
- Q/Λ : dynamical and static dark energy
will be distinguishable
- Q : time varying fundamental coupling “constants”

violation of equivalence principle

????????????????????

Why becomes Quintessence dominant in the present cosmological epoch ?

Are dark energy and dark matter related ?

Can Quintessence be explained in a fundamental unified theory ?

Cosmon dark matter

Can the cosmon account
for dark matter ?

Scalar field (cosmon)
can vary in space

$$\varphi(\vec{x}, t) = \varphi_0(t) + \chi(\vec{x}, t)$$

quintessence ,
homogeneous dark
energy

cosmon
fluctuations,
cosmon dark
energy

$$\varphi_0(t) = \frac{1}{V} \int d^3x \varphi(\vec{x}, t)$$

cosmological expectation value

- * similar to gravity
- * different for gauge bosons, fermions

energy density in cosmon fluctuations ρ_c

$$\rho_c = \frac{1}{2} \int \frac{d^3k}{(2\pi)^3} \left\{ |\dot{\chi}_k|^2 + \left(\frac{k^2}{a^2} + V''(\varphi_0) \right) |\chi_k|^2 + \text{higher order terms} \right\}$$

quintessence ρ_q

$$\rho_q = \frac{1}{2} \dot{\varphi}_0^2 + V(\varphi_0) = T + V$$

Different equation of state
for ρ_c, ρ_g ?

$$w = p / \rho$$

Well possible !

e.g. • ρ_c dominated by modes inside
the horizon, $\frac{k^2}{a^2} \gg H^2$

• neglect higher order terms

$$a) \quad \frac{k^2}{a^2} \gg V'' \Rightarrow \frac{p_c}{\rho_c} = \frac{1}{3}, \text{ radiation}$$

$$b) \quad \frac{k^2}{a^2} \ll V'' \Rightarrow \frac{p_c}{\rho_c} = 0, \text{ matter}$$

but

$$\frac{p_g}{\rho_g} = \frac{T - V}{T + V}, \text{ can be negativ !}$$

◆ most quintessence models :

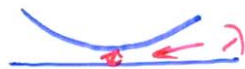
$$V'' \approx H^2$$

$$\Rightarrow \frac{p_c}{\rho_c} = \frac{1}{3} \quad \text{or}$$

nonlinear terms play a role !

◆ one can construct models
with $V'' \gg H^2$ (Matos et al)

\Rightarrow cosmon dark matter



$$(H \approx 10^{-33} \text{ eV})$$

Can nonlinear effects induce an
effective dynamical mass term ?

Why has quintessence become
important now ?

$$\frac{d\Omega_h}{d \ln a} = -3w_h \Omega_h (1 - \Omega_h)$$

($a \gg a_{eq}$)

a) properties of cosmon
potential or kinetic term

(i) Late quintessence

$$w_h \approx \text{const} , \quad (w_h < -\frac{1}{3})$$

Ω_h negligible in early cosmology

needs tiny parameter like cosmological

constant, $V(a_p)/M_p^4 = 10^{-\text{many}}$

(ii) early quintessence

Ω_h changes only modestly

w_h changes with time

special feature in cosmon potential

or kinetic term becomes important

"now", "tuning" at % level

b) quintessence reacts to
„special event“ in cosmology

* onset of Matter dominance

k -essence

Amendarias-Picon, Mukhanov, Steinhardt

needs higher derivative kinetic term

* appearance of non-linear structure

„backreaction effect“

scalar evolution equation

$$\langle \ddot{\phi} + 3H\dot{\phi} + V'(\phi) \rangle = 0$$

$$0 = \ddot{\phi}_0 + 3H\dot{\phi}_0 + V'(\phi_0) + V''(\phi_0)\langle\chi\rangle + \frac{1}{2}V''(\phi_0)\langle\chi^2\rangle$$

fluctuation effect

backreaction

(In principle, same for metric, but
small effect)

- * Needs large inhomogeneities in
cosmon field after structure has
formed
- * Local cosmon field participates in structure

Need for understanding of
fluctuations on nonlinear level

Non-equilibrium QFT



end