

Einstein-Cartan Theory as an Averaged Theory of Gravity

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The Averaging Problem

Teleparallel Gravity

Averaging Process

Einstein-Cartan Theory

Gravitation in Macroscopic Media

The Averaging Problem of General Relativity

- ▶ Standard Cosmology based on

$$G_{\mu\nu}(\langle g_{\mu\nu} \rangle) = 8\pi G \langle T_{\mu\nu} \rangle + \Lambda \langle g_{\mu\nu} \rangle$$

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⇒ Modifications can in principle act as a dark energy

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⇒ Such a manifold is called a **Riemann-Cartan spacetime** U_4 .

- In a Riemann-Cartan spacetime the difference between the affine connection and the Levi-Civita connection defines the **contortion tensor**

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⇒

$$(L_4, g) \xrightarrow{Q=0} U_4 \xrightarrow{T=0} V_4 \xrightarrow{R=0} R_4$$

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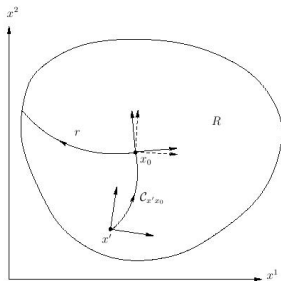
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$$\Rightarrow \frac{du_\lambda}{ds} - \Gamma_{\mu\lambda\nu} u^\mu u^\nu = K_{\lambda\mu\nu} u^\mu u^\nu$$

Averaging Process in General Relativity

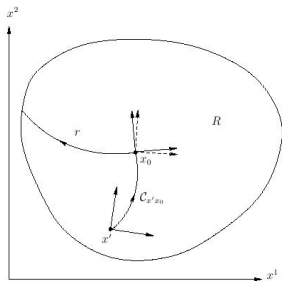


Parallel transport along geodesics $C_{xx'}$
realized by Wegner-Wilson line operator

$$V(x', x; C_{xx'}) = \mathcal{P} \exp \left[- \int_{C_{xx'}} dz^\mu \Gamma_\mu(z) \right]$$

where $\Gamma_\mu(x)$ are four matrices with
components $(\Gamma_\mu(x))^\lambda_\nu = \Gamma^\lambda_{\mu\nu}(x)$

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where $\Gamma_\mu(x)$ are four matrices with
components $(\Gamma_\mu(x))^\lambda_\nu = \Gamma^\lambda_{\mu\nu}(x)$

$$\langle T^\mu{}_\nu \rangle(x) = \frac{1}{V_\Sigma} \int_\Sigma V^{\mu}{}_{\mu'}(x, x'; C_{xx'}) \hat{V}_\nu{}^{\nu'}(x, x'; C_{xx'}) T^{\mu'}{}_{\nu'}(x') \sqrt{-g(x')} d^4 x'$$

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Hehl, von der Heyde, Kerlick, and Nester:

- ▶ Lagrangian invariant under Poincaré gauge transformations

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- ▶ They are related according to

$$\mu^{\nu\lambda\mu} = \frac{1}{2} (\tau^{\nu\lambda\mu} - \tau^{\lambda\mu\nu} + \tau^{\mu\nu\lambda})$$

and

$$\Sigma^{\mu\nu} = \sigma^{\mu\nu} + (\nabla_\lambda - T^\rho{}_{\lambda\rho}) \mu^{\mu\nu\lambda}$$

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$$\begin{aligned} & \mathring{R}^{\mu\nu} - \frac{1}{2} \mathring{R} g^{\mu\nu} \\ &= 8\pi G \sigma^{\mu\nu} + \frac{(8\pi G)^2}{4} (-2\tau_{\lambda}{}^{\mu\rho} \tau_{\rho}{}^{\nu\lambda} + 2\tau^{\lambda\mu}{}_{\lambda} \tau^{\rho\nu}{}_{\rho} - 2\tau^{\lambda\mu\rho} \tau_{\lambda}{}^{\nu}{}_{\rho} \\ & \quad + \tau^{\mu\rho\lambda} \tau_{\rho}{}^{\nu}{}_{\lambda} + \frac{1}{2} g^{\mu\nu} (2\tau_{\lambda\rho}{}^{\sigma} \tau_{\sigma}{}^{\rho\lambda} - 2\tau^{\lambda}{}_{\rho\lambda} \tau^{\sigma\rho}{}_{\sigma} + \tau^{\lambda\sigma\rho} \tau_{\lambda\sigma\rho})) \end{aligned}$$

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with the polarization tensor

$$P_{\mu\nu\rho\sigma} = \frac{1}{2} \left(-Q'_{\rho\sigma\epsilon[\mu}{}^{,\epsilon} - \frac{1}{3}\eta_{\rho[\mu} Q'^{\gamma}{}_{\sigma\gamma\epsilon}{}^{,\epsilon]} \right)_{,\nu]}$$

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