

NEUTRINOS & EUCLID

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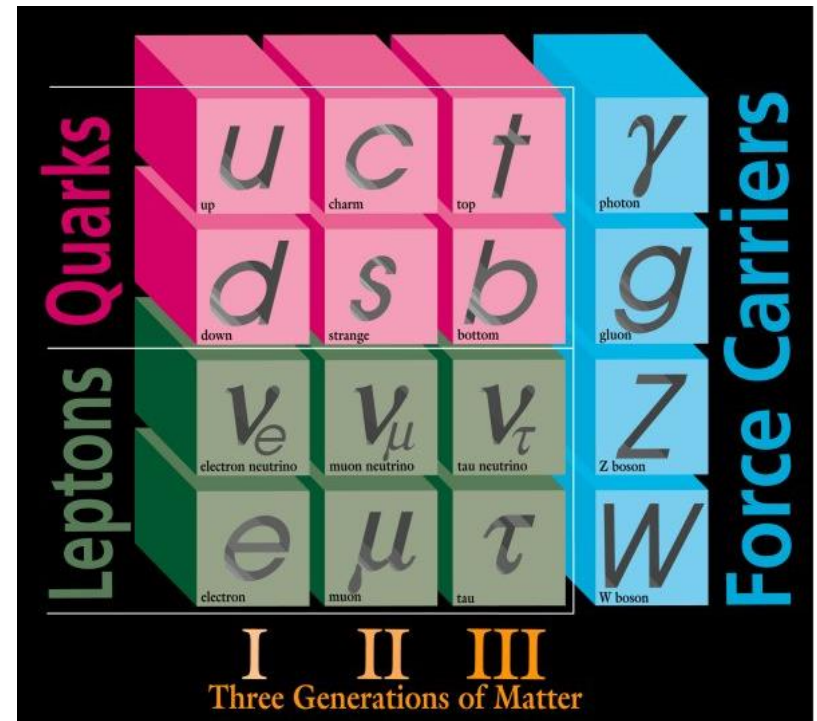
- ❖ Neutrino properties
- ❖ Matter power spectrum
- ❖ Relic neutrinos
- ❖ Neutrino mass hierarchy
- ❖ Majorana or Dirac ?

NEUTRINOS IN THE STANDARD MODEL

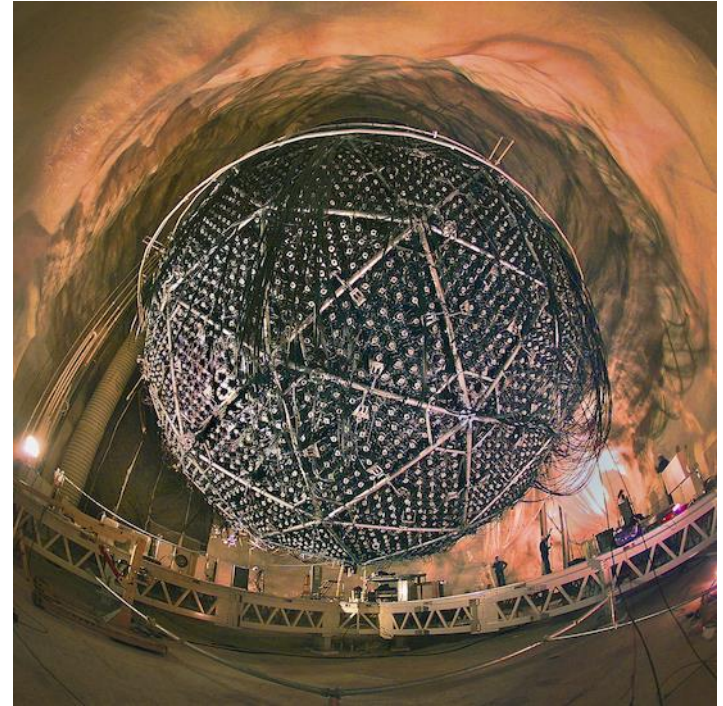
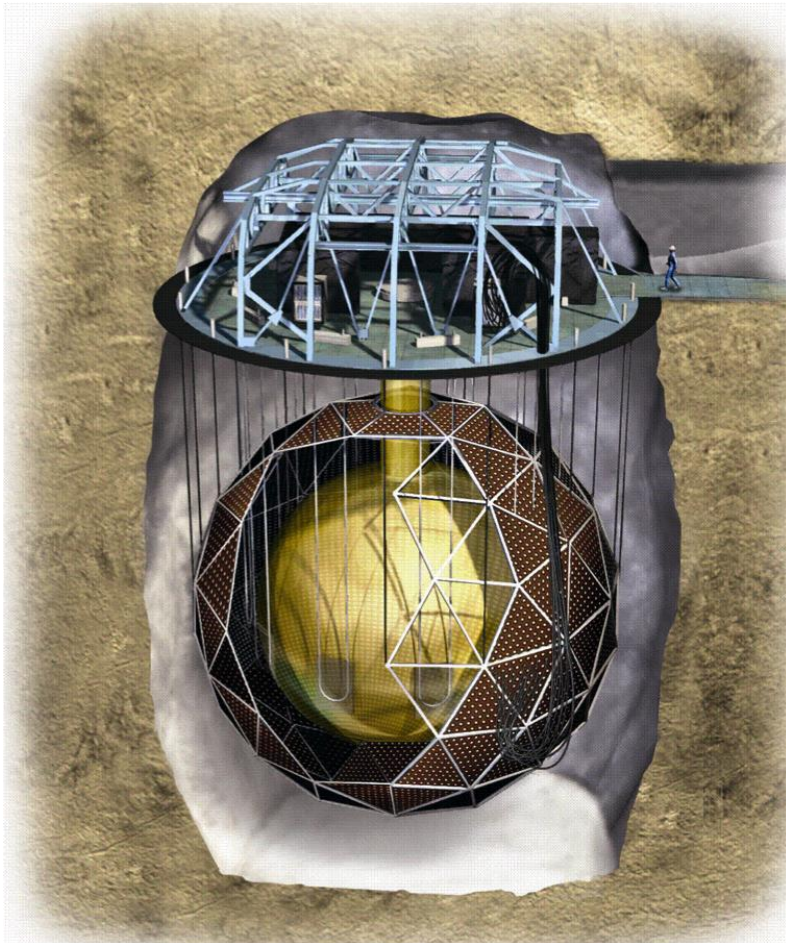
- ❖ Three neutrino generations
- ❖ Purely weak-interacting
- ❖ Only coupling to left-handed ν and right-handed $\bar{\nu}$

$$j_{\mu} = \bar{\psi}\gamma_{\mu}(1 - \gamma_5)\psi$$

- ❖ Massless $m_{\nu,i} = 0$



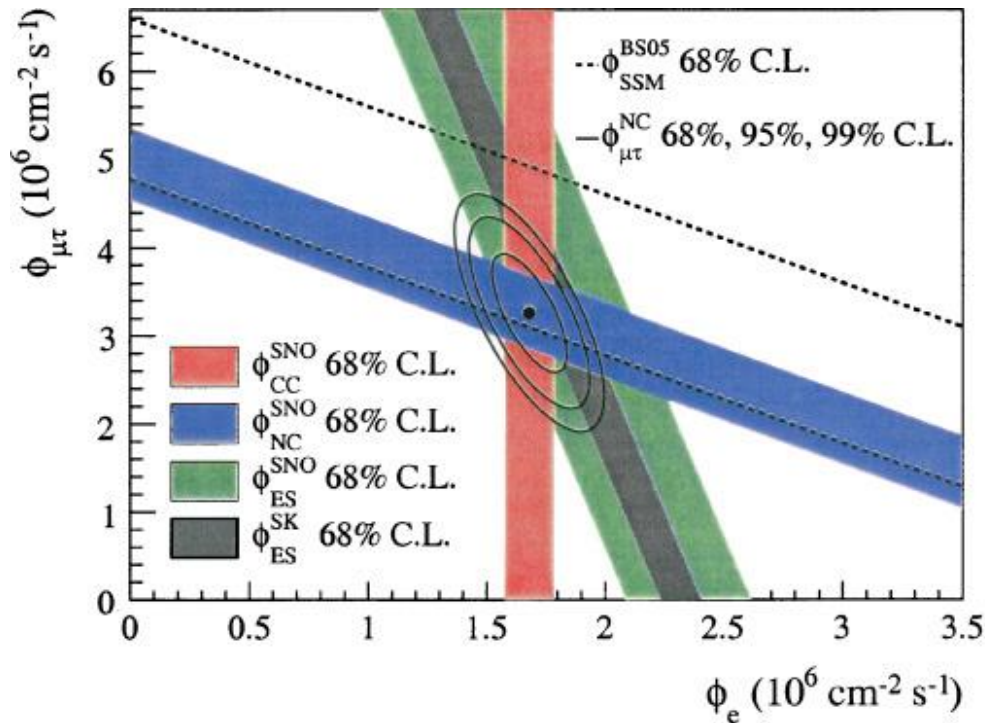
THE SNO EXPERIMENT



- ❖ 1000 tons heavy water detector
- ❖ Measures solar $\phi(\nu_e)$, $\phi(\nu_{\mu\tau})$, $\phi(\nu_{tot})$
- ❖ Sun only produces ν_e
- ❖ *Expected flux:*

$$\phi(\nu_e) = 5.1 \times 10^6 \text{ cm}^{-1} \text{ s}^{-1}$$

NEUTRINO FLAVOUR OSCILLATIONS



Results:

$$\phi(\nu_{\text{tot}}) = 5.1 \times 10^6 \text{ cm}^{-1} \text{ s}^{-1}$$

$$\phi(\nu_e) = 1.76 \times 10^6 \text{ cm}^{-1} \text{ s}^{-1}$$

$$\phi(\nu_{\mu\tau}) = 3.41 \times 10^6 \text{ cm}^{-1} \text{ s}^{-1}$$

→ Neutrinos oscillate !

$$P(\nu_e \rightarrow \nu_\mu) \propto \left(\sin \frac{c\Delta m^2 L}{E} \right)^2$$

E: Energy, L: Distance, $\Delta m = m_\mu - m_e$

→ Neutrinos have mass !

HOW TO ADD NEUTRINO MASS?

Dirac mass term: $\mathcal{L}_D = -m_D(\bar{\nu}_R\nu_L + \bar{\nu}_L\nu_R)$

Majorana mass term: $\mathcal{L}_M = -\frac{1}{2}M(\bar{\nu}_R^C\nu_R + \bar{\nu}_R\nu_R^C)$

} \rightarrow Existence of ν_R

} \rightarrow Neutrino its own anti-particle

Combine:

$$\mathcal{L}_{DM} = -\frac{1}{2}(\bar{\nu}_L \quad \bar{\nu}_R^C) \begin{pmatrix} 0 & m_D \\ m_D & M \end{pmatrix} \begin{pmatrix} \nu_L^C \\ \nu_R \end{pmatrix}$$

Assume $m_D \ll M$:

\rightarrow Light neutrino state: $|m_\nu| \approx \frac{m_D^2}{M}$

\rightarrow Heavy neutrino state: $m_N \approx M$

Is the neutrino majorana?

MATTER POWER SPECTRUM

Density perturbation field, with matter density ρ

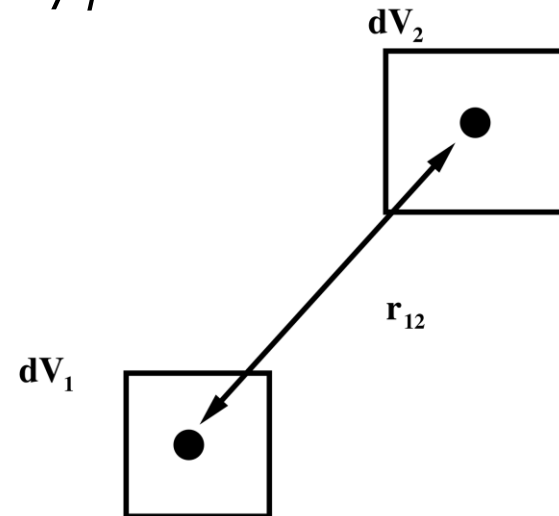
$$\delta(\mathbf{x}) \equiv \frac{\rho(\mathbf{x}) - \langle \rho \rangle}{\langle \rho \rangle}$$

Correlation function

$$\xi(\mathbf{r}) \equiv \langle \delta(\mathbf{x})\delta(\mathbf{x} + \mathbf{r}) \rangle$$

Matter power spectrum $P(k)$

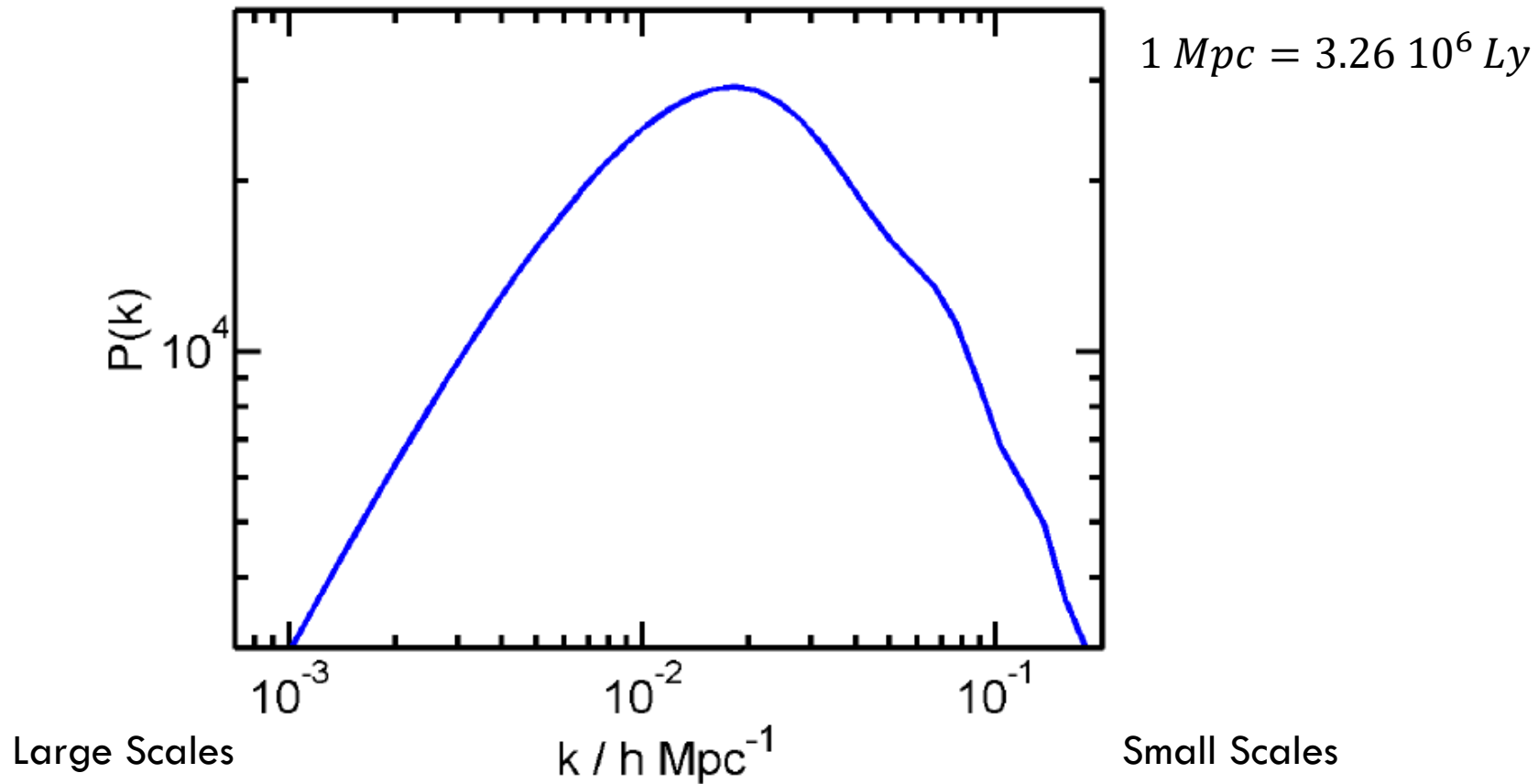
$$\xi(r) = \int \frac{d^3k}{(2\pi)^3} e^{i\mathbf{k}\cdot\mathbf{r}} P(k)$$



Probability to find a neighbour

$$dW = \rho_0^2 [1 + \xi(r)] dV_1 dV_2$$

THE MATTER POWER SPECTRUM



→ Can be determined from EUCLID's weak lensing measurement

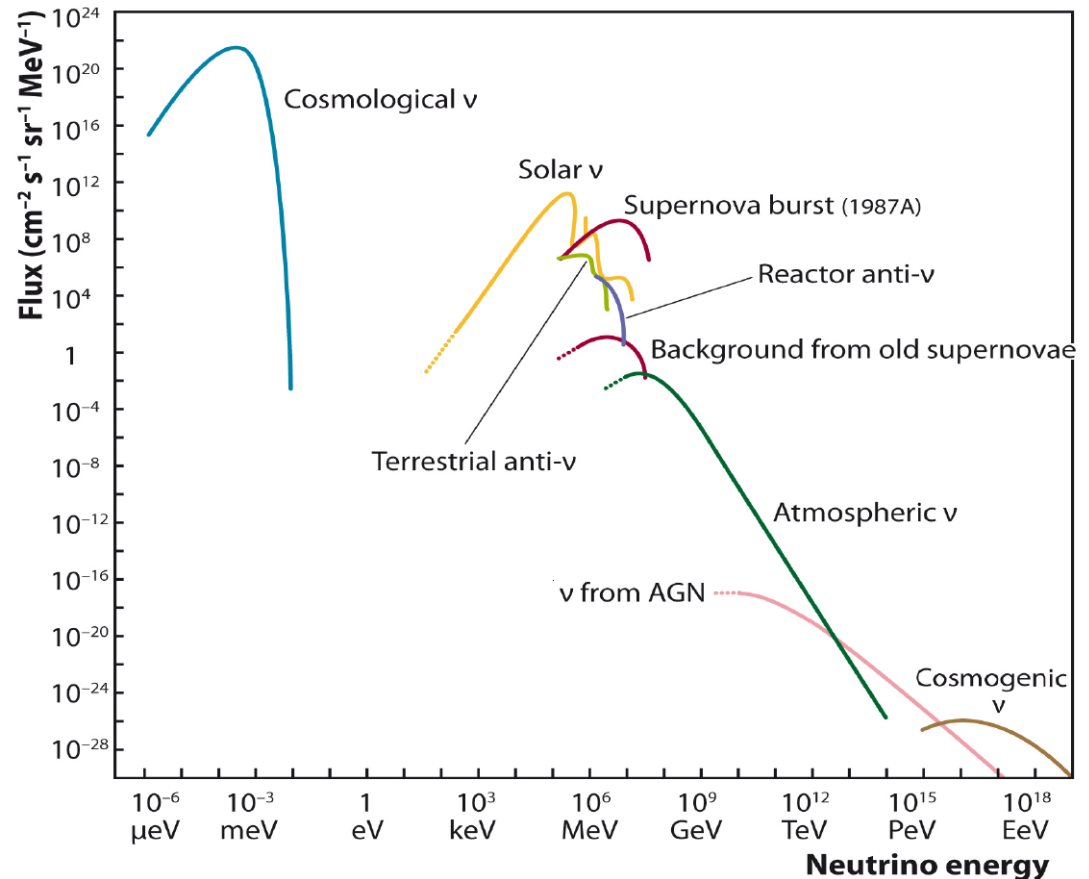
RELIC NEUTRINOS

- ❖ Standard hot big bang produces neutrinos

$$\sim 100 N_\nu \text{cm}^{-3}$$

- ❖ Not yet (directly) detected!

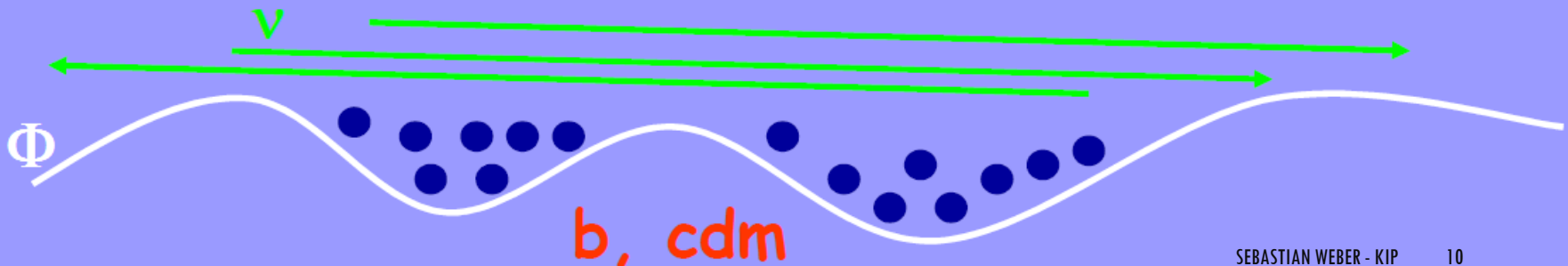
- ❖ EUCLID could detect influence on structure formation



RELIC NEUTRINOS & STRUCTURE FORMATION

- ❖ Neutrinos decouple very early ($\sim 0.2 s$) from primordial plasma
- ❖ Initially $T \approx 1 MeV \rightarrow$ Neutrinos ultra-relativistic
- ❖ Still relativistic in the matter dominated universe
- ❖ Small-scale density fluctuations damped by free-streaming

Neutrino Free Streaming



RELIC NEUTRINOS – FREE STREAMING LENGTH

- ❖ Free-streaming length λ_{FS} determines damping of density fluct.

→ Neutrinos cannot be confined within λ_{FS}

- ❖ Maximum of λ_{FS} at non-relativistic transition

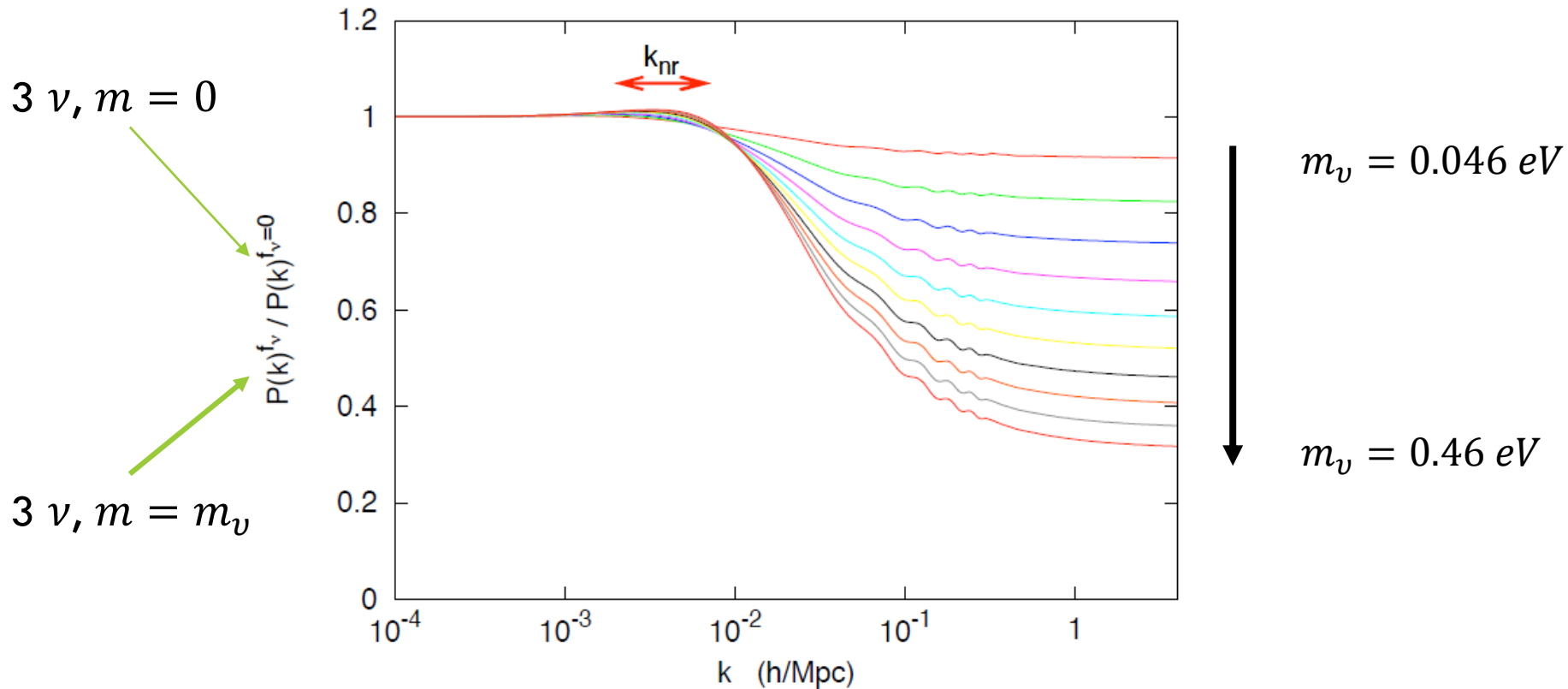
→ Minimum free-streaming wavenumber

$$k_{nr} \propto \left(\frac{m_\nu}{1 \text{ eV}} \right)^{1/2} h \text{ Mpc}^{-1}$$

- ❖ Damping of $P(k)$ for modes with $k > k_{nr}$ due to Neutrino mass

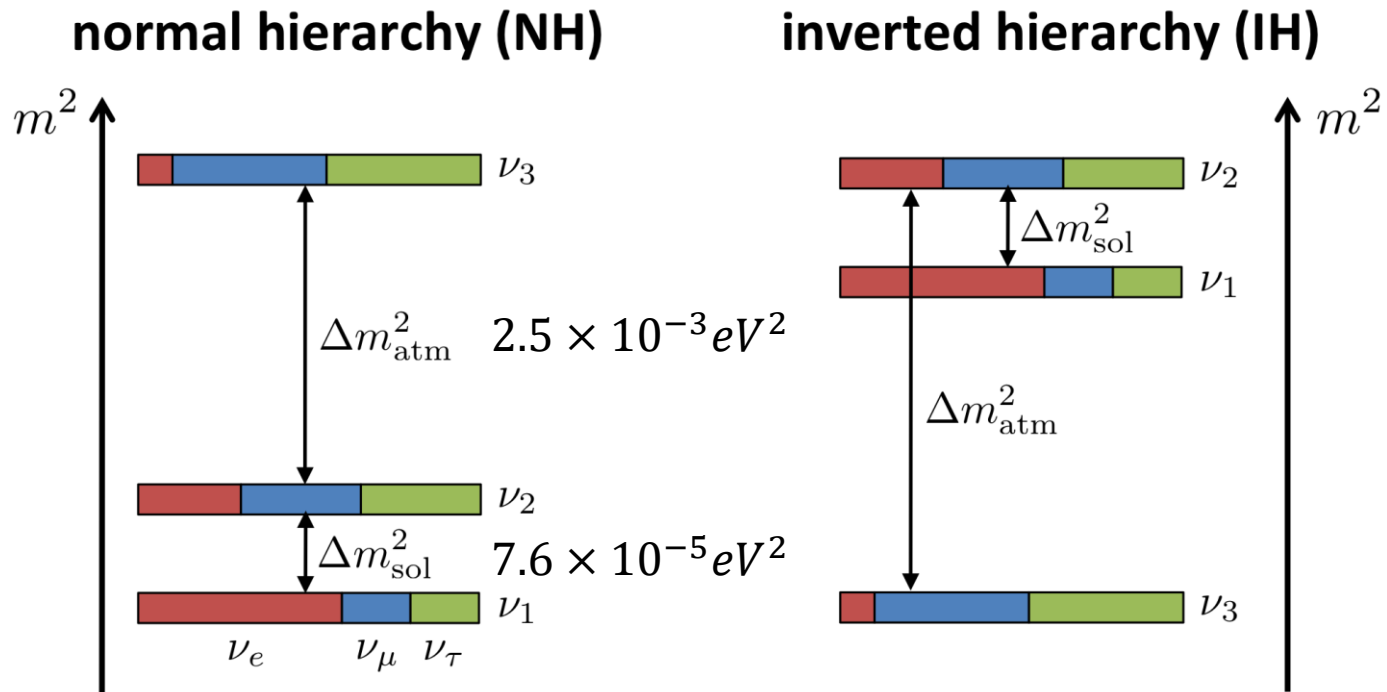
→ Detection of relic Neutrinos within reach of EUCLID

RELIC NEUTRINOS – POWER SPECTRUM



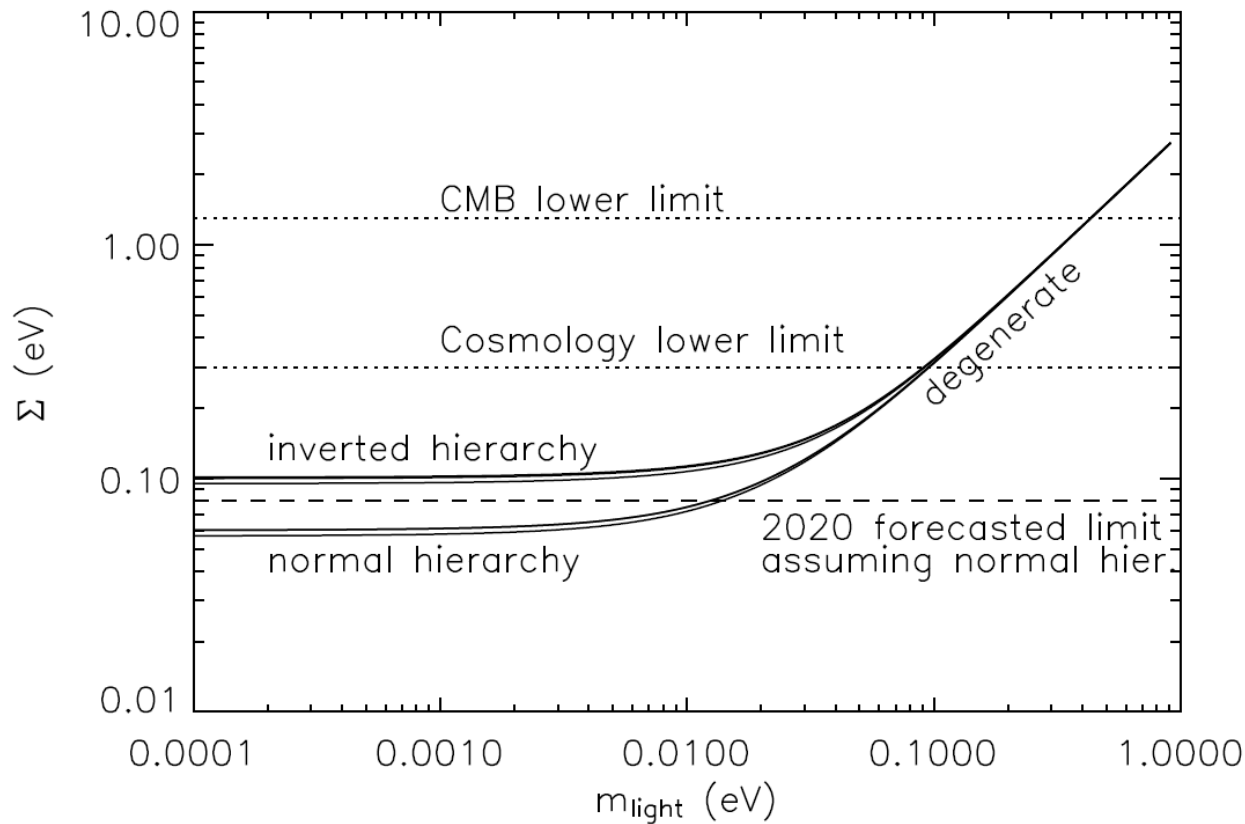
→ Higher neutrino mass leads to stronger damping! → EUCLID: $\Sigma < 0.05 \text{ eV}$

NEUTRINO MASS HIERARCHY



Third possibility: Splitting small against masses \rightarrow Degenerate Case

NEUTRINO MASS HIERARCHY



→ For total mass $\Sigma < 0.1 \text{ eV}$, EUCLID can constrain hierarchy

NEUTRINO MASS HIERARCHY

- ❖ Neutrinos of different mass have different free streaming length

→ Detailed shape of $P(k)$ depends on individual m_ν

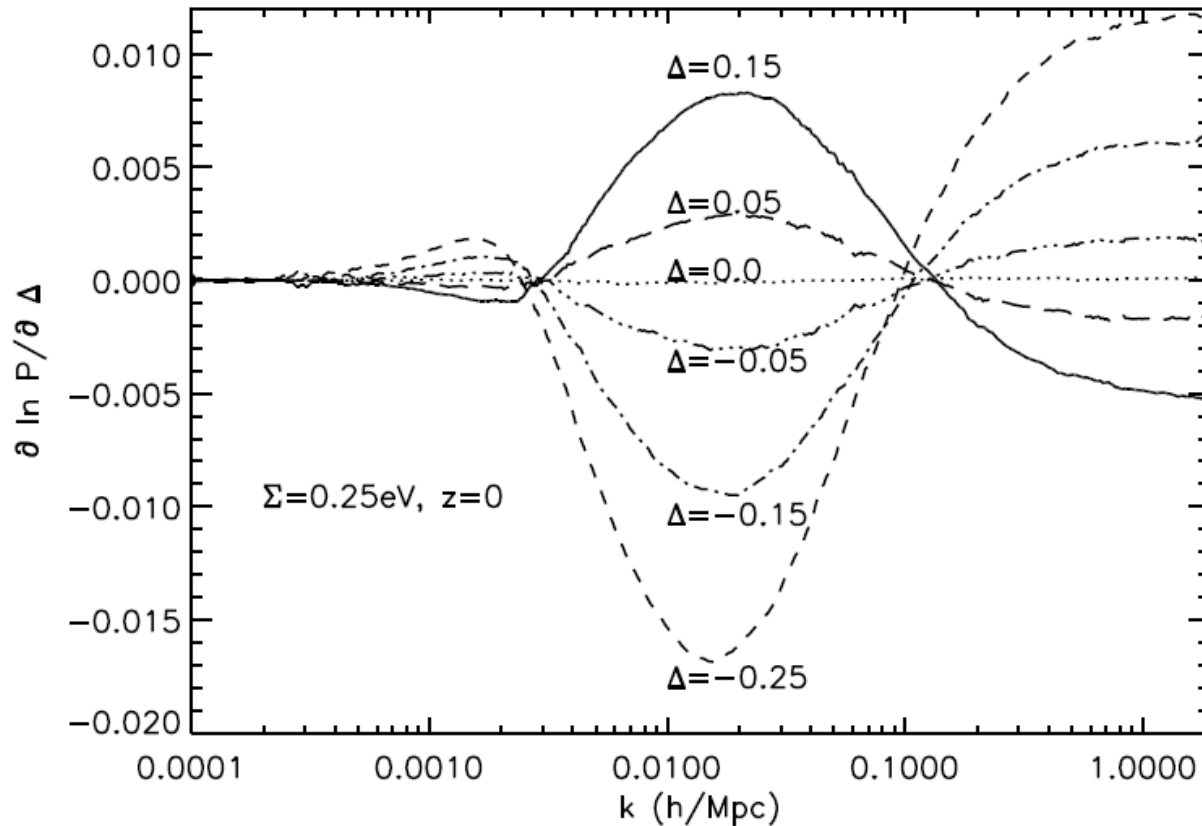
- ❖ Smaller splitting negligible, consider light mass (m) and heavy mass (M)

$$\text{NH:} \quad \Sigma = 2m + M \quad \Delta = (M - m)/\Sigma > 0$$

$$\text{IH:} \quad \Sigma = m + 2M \quad \Delta = (m - M)/\Sigma < 0$$

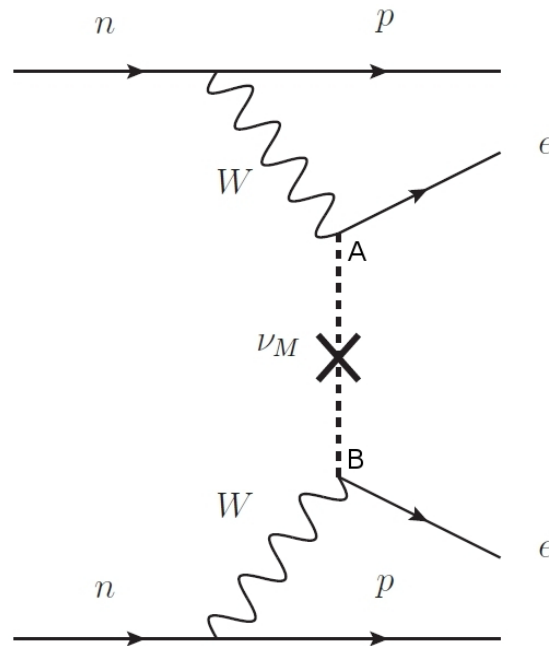
- ❖ Exploit dependence of $P(k)$ on Δ , sign of Δ gives hierarchy

NEUTRINO MASS HIERARCHY



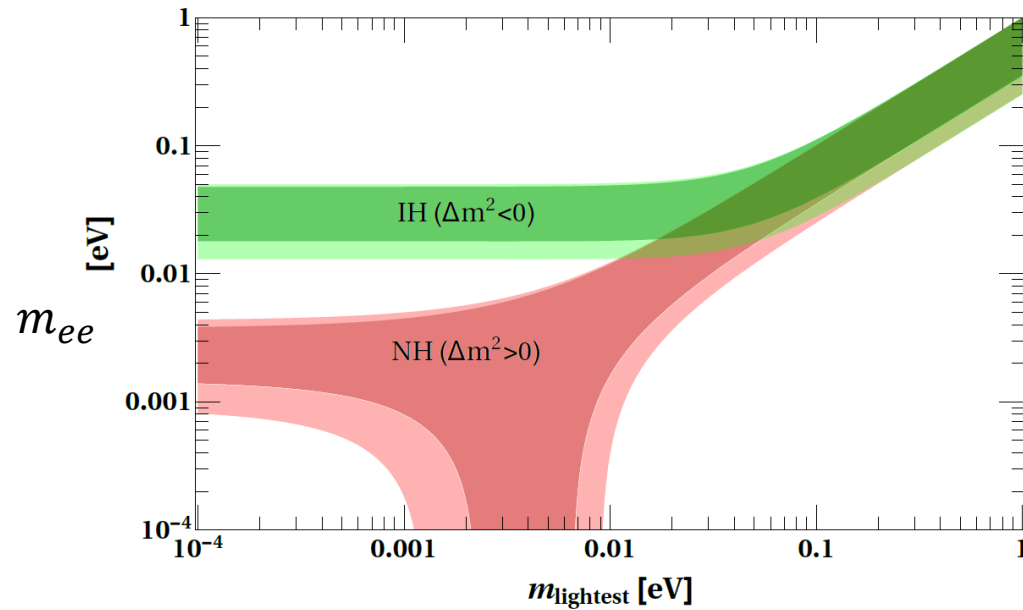
EUCLID might be able to measure this!

MAJORANA VS DIRAC



- ❖ With Majorana neutrinos, neutrinosless double-beta decay possible
 - ❖ Search: Gerda, Majorana, ...
- ❖ Rate is proportional to effective mass squared m_{ee}^2
- ❖ Next generation experiments probe $m_{ee} \approx 10 \text{ meV}$

MAJORANA VS DIRAC



If no $0\nu\beta\beta$ observed

EUCLID determines IH

Neutrino is a dirac particle!

CONCLUSION

- ❖ EUCLID can perform very precise measurement of matter power spectrum
- ❖ EUCLID could detect relic neutrinos & determine the total neutrino mass
- ❖ EUCLID has a chance to determine the neutrino mass hierarchy
- ❖ EUCLID can contribute to the question if neutrinos are Majorana or Dirac