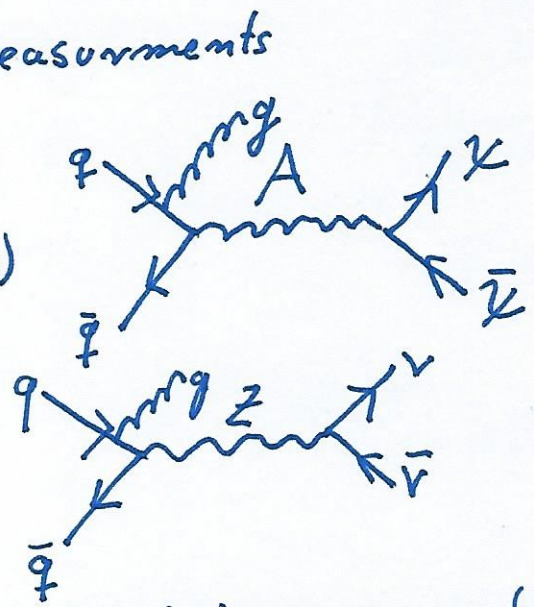


# Transitioning from a Dark Matter search to a precision measurement at the LHC

Lecture 1: - electro weak precision measurements  
- oblique parameters

Lecture 2: - LHC Dark matter searches  
- 2 Higgs Doublet Models (2HDMs)

Lecture 3: -  $\Gamma_Z$  (invisible) measurement



Standard Model of particle physics has 18 (+1) free parameters:

- 9 fermion masses (quarks + charged leptons)
- 3 + 1 angles (CKM + 1EP)
- $g_s$  strong coupling constant
- $(\Theta_{QCD})$

+ 4 electroweak parameters:

Electroweak symmetry breaking:  $SU_2(L) \times U(1)_Y \rightarrow U(1)_{em}$

-  $\mathcal{L}_{Higgs} = |D_\mu \phi|^2 - V(\phi)$  with  $V(\phi) = -\mu \phi^\dagger \phi + \lambda (\phi^\dagger \phi)^2$   
 $D_\mu = i \partial_\mu + g \frac{\tau^i}{2} W_\mu^i + g' \frac{Y}{2} B_\mu$   $\tau^i$ : Pauli matrices

$SU(2)$  scalar  $\phi = \begin{pmatrix} \phi^+ \\ \phi^0 \end{pmatrix} \xrightarrow{EWSB} \frac{1}{\sqrt{2}} \begin{pmatrix} 0 \\ v+h \end{pmatrix}$

$\Rightarrow$  defines:  $W_\mu^\pm = \frac{1}{\sqrt{2}} (W_\mu^1 \mp W_\mu^2)$

$Z_\mu = \frac{-g' B_\mu + g W_\mu^3}{\sqrt{g^2 + g'^2}} \equiv \sin \Theta_w B_\mu + \cos \Theta_w W_\mu^3$

$A_\mu = \frac{g B_\mu + g' W_\mu^3}{\sqrt{g^2 + g'^2}} \equiv \cos \Theta_w B_\mu + \sin \Theta_w W_\mu^3$

$\Rightarrow \sin \Theta_w = \frac{g'}{\sqrt{g^2 + g'^2}}$

gauge bosons acquire masses:

$m_W^2 = \frac{1}{4} g^2 v^2$ ;  $m_Z^2 = \frac{1}{4} (g^2 + g'^2) v^2$ ;  $m_A^2 = 0$



vacuum expectation value  $v$  combined with  $G_F$  (determined from  $\mu \rightarrow e \nu_e \bar{\nu}_\mu$ ):

$$\frac{G_F}{\sqrt{2}} = \frac{g^2}{8m_W^2} = \frac{1}{2v^2} \Rightarrow v = 246 \text{ GeV}$$

$$m_h = 2v^2 \lambda$$

$\Rightarrow$  4 fundamental parameters of EW sector: - couplings  $g, g'$   
-  $v, m_h$

**EW part of SM is overconstrained  $\Rightarrow$  predictive  $\Rightarrow$  testable!**

Set with smallest exp uncertainties:

$$m_Z = 91.1876 \pm 0.0021 \text{ GeV} \quad (\text{Z-line shape measurement @ LEP})$$

$$G_F = 1.1663787(6) \times 10^{-5} \text{ GeV}^{-2} \quad (\text{muon lifetime @ Muon PSI})$$

$$\alpha = 1/137.035999139(31) \quad (\text{pe measurements combined by CODATA})$$

$$m_h = 125.10 \pm 0.14 \text{ GeV} \quad (\text{LHC combination PDG})$$

examples of additional EW observables:  $m_W, \sin^2 \theta_W$

many results still from: - LEP ( $e^+e^-$  209 GeV)

- SLC ( $e^+e^-$  91 GeV)

- Tevatron ( $p^+p^-$  1.8 TeV)

$m_W$  measurement:

Motivation: SM predicts relations between  $m_W, m_h, m_e$

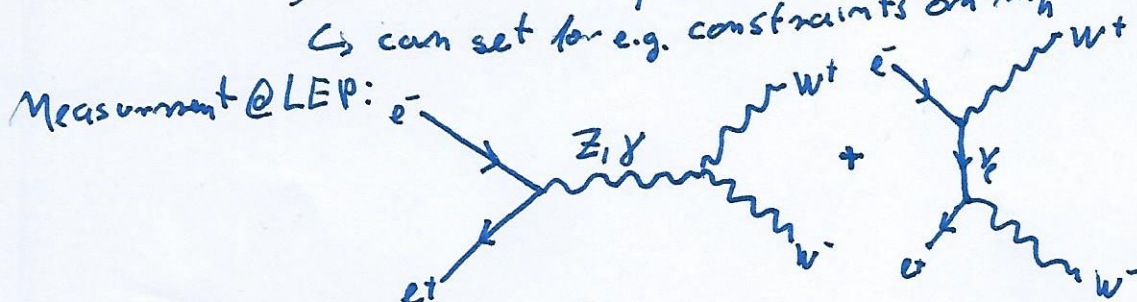
$$m_W = \left( \frac{\pi \alpha}{\sqrt{2} G_F} \right)^{1/2} \frac{1}{\sin \theta_W \sqrt{1 - 4s^2}}$$

$\uparrow$  radiative corrections  $\sim 3\%$   
depend on  $m_e^2, \log m_h, H$

$$m_W (\text{tree-level}) = 79.829 \text{ GeV}$$

$$m_W (\text{exp}) = 80.379 \pm 0.012 \text{ GeV}$$

$\Rightarrow$  consideration of radiative corrections necessary  
 $\hookrightarrow$  can set for e.g. constraints on  $m_h$



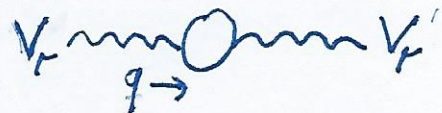


# Oblique Parameters (Peskin, Takeuchi)

- Extensions to the Higgs sector can be constrained by EW observables
- oblique parameters are reparametrisations of radiative correction variables

$$\Delta\alpha, \Delta\kappa, \Delta\varrho = \frac{m_W^2}{m_Z^2 \cos^2 \theta_W} - 1$$

## 1 loop self-energy corrections to $\gamma, W^\pm, Z$ :



$$i [\Pi_{VV'}(q^2) g^{\mu\nu} - \Delta_{VV'}(q^2) q^\mu q^\nu] \quad \text{2-point correlation functions}$$

↑ vanishes for light quark limit

$$\Rightarrow m_V^2 = m_V^2 + \Pi_{VV}(q^2 = m_V^2)$$

$$\gamma \text{ is massless} \rightarrow \Pi_{\gamma\gamma}(0) = 0 = \Pi_{\gamma Z}(0)$$

$$\frac{\Delta S}{4 \sin^2 \theta_W \cos^2 \theta_W} = \frac{\Pi_{ZZ}(m_Z^2) - \Pi_{ZZ}(0)}{m_Z^2} - \frac{c_W^2 - s_W^2}{c_W s_W} \frac{\Pi_{\gamma Z}(m_Z^2)}{m_Z^2} - \frac{\Pi_{\gamma\gamma}(m_Z^2)}{m_Z^2}$$

$$\Delta T = \frac{\Pi_{WW}(0)}{m_W^2} - \frac{\Pi_{ZZ}(0)}{m_Z^2} = \Delta\varrho$$

$$\frac{\Delta U}{4 s_W^2} = \frac{\Pi_{WW}(m_Z^2) - \Pi_{WW}(0)}{m_W^2} - c_W^2 \frac{\Pi_{ZZ}(m_Z^2) - \Pi_{ZZ}(0)}{m_Z^2} - 2 c_W s_W \frac{\Pi_{\gamma Z}(m_Z^2)}{m_Z^2} - s_W^2 \frac{\Pi_{\gamma\gamma}(m_Z^2)}{m_Z^2}$$

T is sensitive to weak isospin violation & proportional to  $\Delta\varrho$

S, T, U constructed in such a way that they vanish in SM

Global EW fit yields:  $S = -0.01 \pm 0.10$  (from  $M_Z$ )

$$T = 0.03 \pm 0.12$$
 (from  $\Gamma_Z^0$ )

$$U = 0.02 \pm 0.11$$
 (from  $M_Z$ )

e.g. extension of the SM Higgs sector by scalar S with  $\gamma_S = 0$  to  $Z$ :  $S \rightarrow S$ :

$$V = -f^2 \phi^\dagger \phi - m^2 S^2 + \lambda (\phi^\dagger \phi)^2 + \lambda_{\phi S} S^2 \phi^\dagger \phi - \frac{\lambda_S}{4} S^4 \quad \text{with } \langle S \rangle = s$$

$$\begin{pmatrix} h \\ H \end{pmatrix} = \begin{pmatrix} \cos \alpha & \sin \alpha \\ -\sin \alpha & \cos \alpha \end{pmatrix} \begin{pmatrix} \sqrt{2} \phi^0 - v \\ S - s \end{pmatrix} \quad \text{with masses } m_h, m_H \ll m_W, m_Z$$

SM Higgs couplings suppressed by  $\cos \alpha$  via mixing

$$\Delta S = \frac{1}{12\pi} \sin^2 \alpha \log \left( \frac{m_H^2}{m_h^2} \right)$$

$$\Delta T = -\frac{3}{16\pi \cos^2 \theta_W} \sin^2 \alpha \log \left( \frac{m_H^2}{m_h^2} \right) \Rightarrow m_H \text{ restricts } \sin \alpha!$$