

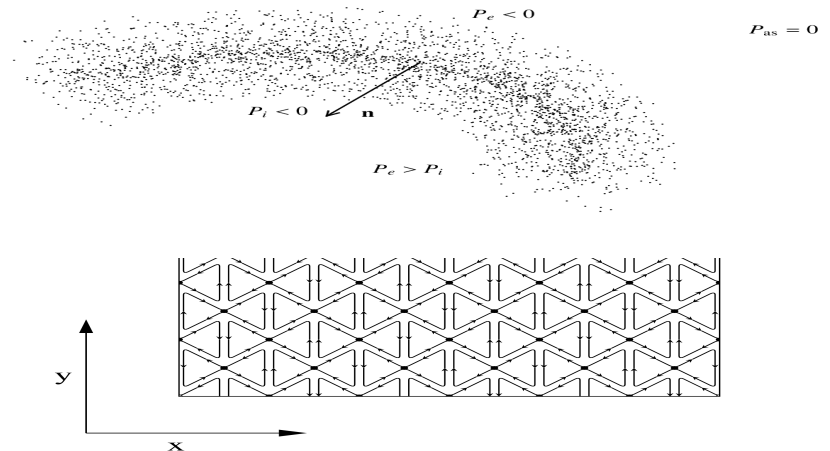
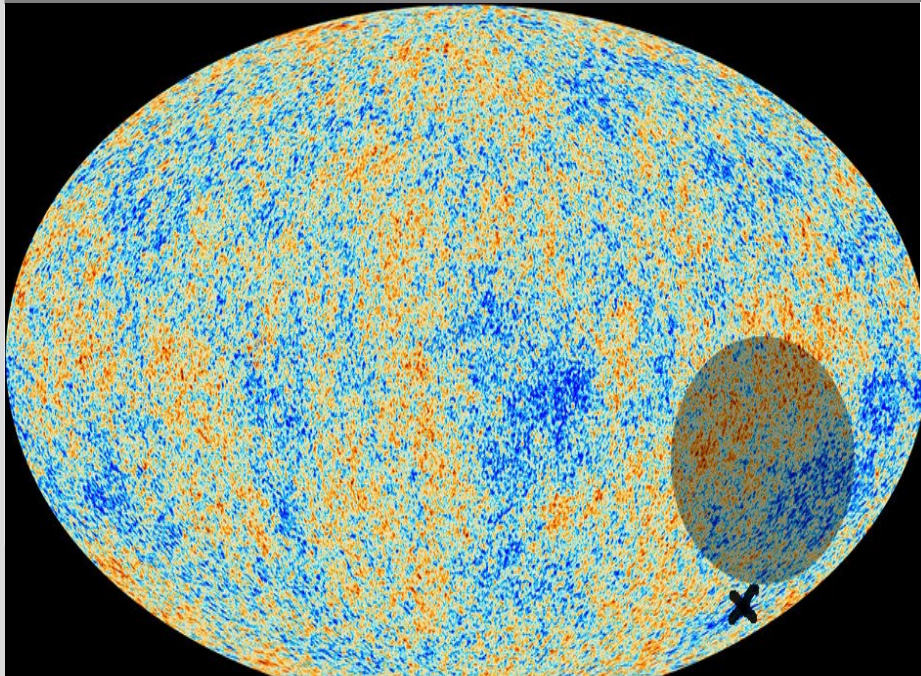


SU(2) Yang-Mills thermodynamics

Seminar, condensed matter theory, 20 November 2014

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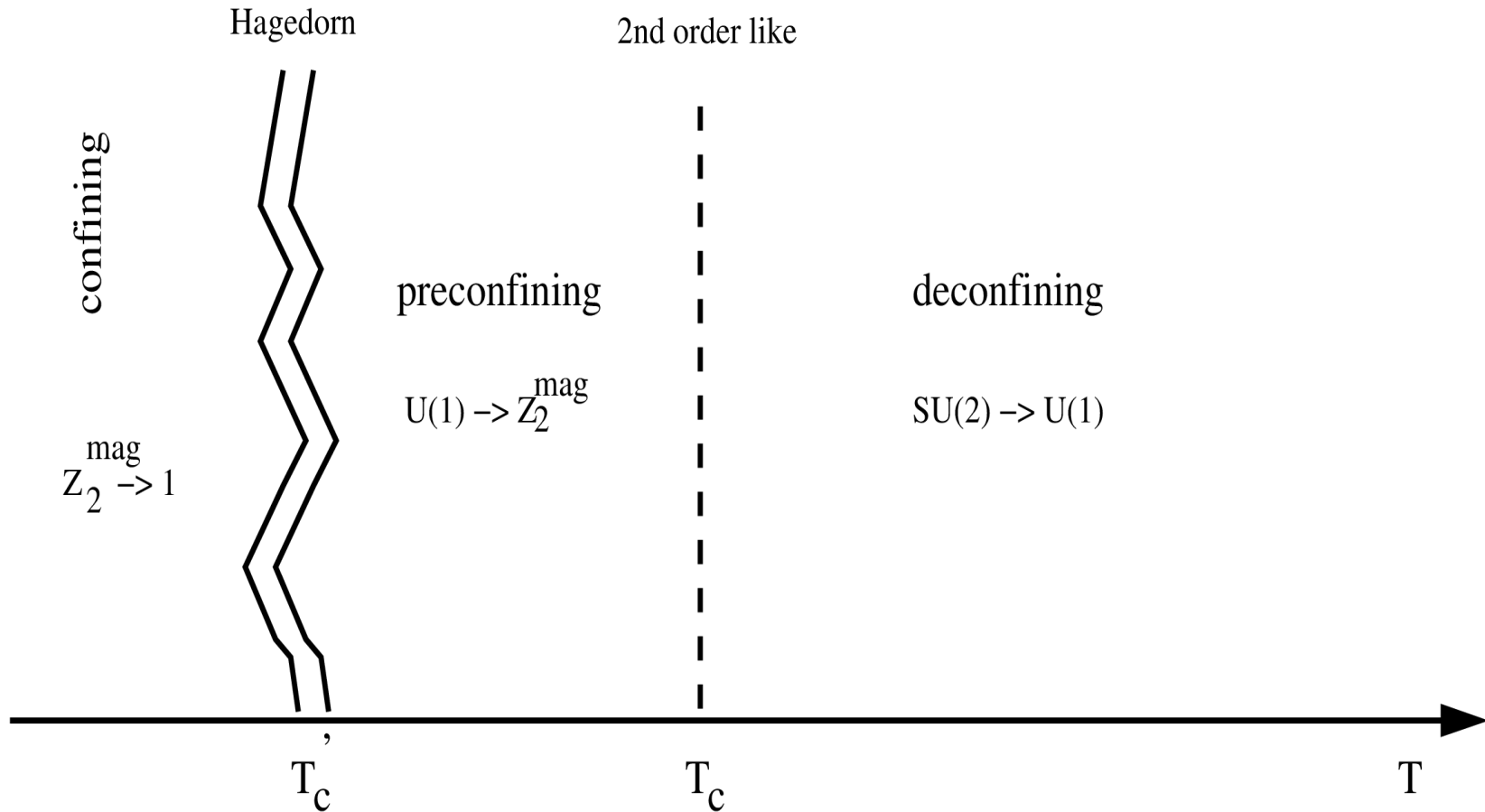
outline

- **motivation:** nonperturbative, analytical approach to YMTD
- **preview on phase structure**
- **essentials, thermal ground state:**
 - coarse-graining over nonpropagating
(anti-)calorons of winding number unity, effective action
- **adjoint Higgs mechanism:**
 - massive vector modes and kinematic constraints (1),
coupling, deconf.-preconf. phase boundary, (anti-)caloron action,
- **radiative corrections:**
 - kinematic constraints (2), polarisation tensor of massless mode,
longitudinal and transverse thermal dispersion, action of a(n)
(anti)caloron, e-m dual interpretation
- **SU(2) postulate for photon propagation:**
 - Yang-Mills scale or critical temperature
(radio-frequency CMB observations)
- **CMB large-angle anomalies (WMAP, Planck):**
 - possible explanation via SU(2) dispersion law,
onset of dynamical breaking of statistical isotropy at redshift unity,
SU(2) vector modes and cosmic neutrinos

- **Hagedorn transition: confining phase**
- **nonthermal behaviour:**
 - Borel resummation of partition function
- **evolving center vortex loops (CVLs)**
 - single CVLs, mass gap
 - one-fold selfintersecting CVLs, emergent order

- Andrei Linde (1980):
„Infrared Problem in the Thermodynamics of the Yang-Mills Gas“
 - soft magnetic sector screened weakly in perturbation theory (infrared instability)
- - not even naive „convergence“ of series since kinetic and interaction energies comparable in this sector
- issue of finite-volume constraints in lattice gauge theory
 - correlations mediated by soft magnetic sector insufficiently probed by available lattice sizes
 - nonperturbative, lattice β function

preview on phase structure



nonperturbative Yang-Mills thermodynamics: SU(2)

[Herbst et Hofmann (2004), Hofmann (2005-2007), Giacosa et Hofmann (2006), Schwarz et al. (2007), Ludescher et Hofmann (2008), Falquez et al. (2010- 2011), Hofmann (2012)]

thermal ground state at high temperature:

- Euclidean action:

$$S = \frac{\text{tr}}{2} \int_0^\beta d\tau \int d^3x F_{\mu\nu} F_{\mu\nu} ,$$

($\beta \equiv 1/T$)

where $F_{\mu\nu} \equiv \partial_\mu A_\nu - \partial_\nu A_\mu - ig[A_\mu, A_\nu]$

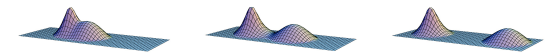
- (anti)selfdual gauge fields:

[Schafer et Shuryak (1996)]

$$F_{\mu\nu}[A] = \pm \tilde{F}_{\mu\nu}[A] \Rightarrow \theta_{\mu\nu}[A] \equiv 0 .$$

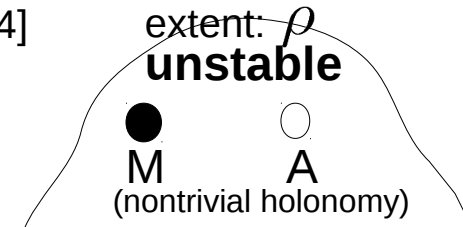
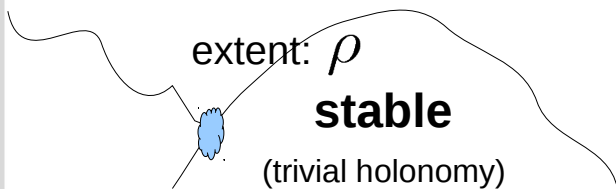
field configs. stabilized by winding: $S_3 \rightarrow SU(2) = S_3$

- in particular: (anti)calorons of winding number unity



[Harrington et Shepard (1977)]

[Nahm (1981-84), Lee et Lu (1998), Kraan et v. Baal (1998), Diakonov et al. 2004]



spatial coarse-graining over (anti-)calorons: inert, adjoint scalar field ϕ

$$\{\hat{\phi}^a\} \equiv \sum_{\pm} \text{tr} \int d^3x \int d\rho t^a F_{\mu\nu}(\tau, \vec{0}) \left\{ (\tau, \vec{0}), (\tau, \vec{x}) \right\} F_{\mu\nu}(\tau, \vec{x}) \left\{ (\tau, \vec{x}), (\tau, \vec{0}) \right\}$$

- unique, definition of (dimensionless) **family of phases**, where

$$\left\{ (\tau, \vec{0}), (\tau, \vec{x}) \right\} \equiv \mathcal{P} \exp \left[i \int_{(\tau, \vec{0})}^{(\tau, \vec{x})} dz_{\mu} A_{\mu}(z) \right] \quad \text{and}$$

$$\left\{ (\tau, \vec{x}), (\tau, \vec{0}) \right\} \equiv \left\{ (\tau, \vec{0}), (\tau, \vec{x}) \right\}^{\dagger}$$

- magnetic-magnetic correlations contribute only

- uniquely determined, annihilating operator:

$$D = \partial_{\tau}^2 + \left(\frac{2\pi}{\beta} \right)^2$$

- $\{\hat{\phi}^a\}$ sharply dominated by cut-off for ρ integration, later!

spatial coarse-graining over (anti-)calorons: inert, adjoint scalar field

- no explicit β dependence in ϕ field dynamics (caloron action!)
- absorb β dependence of D into potential V

(BPS and EL yield: $\frac{\partial V(|\phi|^2)}{\partial |\phi|^2} = -\frac{V(|\phi|^2)}{|\phi|^2} \Rightarrow$

$$V(|\phi|^2) = \frac{\Lambda^6}{|\phi|^2}$$

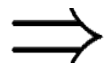
(Yang-Mills scale)

and

$$|\phi| = \sqrt{\frac{\Lambda^3 \beta}{2\pi}}$$

- BPS equation:

$$\partial_\tau \phi = \pm 2i \Lambda^3 t_3 \phi^{-1} = \pm i V^{1/2}(\phi)$$



no **additive** ambiguity for V !

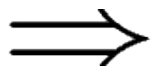
effective action (deconfining phase)

$$\mathcal{L}_{\text{eff}}[a_\mu] = \text{tr} \left(\frac{1}{2} G_{\mu\nu} G_{\mu\nu} + (D_\mu \phi)^2 + \frac{\Lambda^6}{\phi^2} \right)$$

- (i) perturbative renormalizability
- (ii) ϕ 's inertness – no higher dim. operators to mediate 4-momentum transfer between ϕ and a_μ
- (iii) gauge invariance)

- effective YM equation $D_\mu G_{\mu\nu} = ie[\phi, D_\nu \phi]$ has ground-state solution:

$$a_\mu^{\text{gs}} = \mp \delta_{\mu 4} \frac{2\pi}{e\beta} t_3 \quad (D_\nu \phi \equiv G_{\mu\nu} \equiv 0)$$



$$P_{gs} = -\rho_{gs} = -4\pi\Lambda^3 T.$$



interacting small-holonomy
(anti)calorons
(collapsing monopole-
antimonopole pairs)

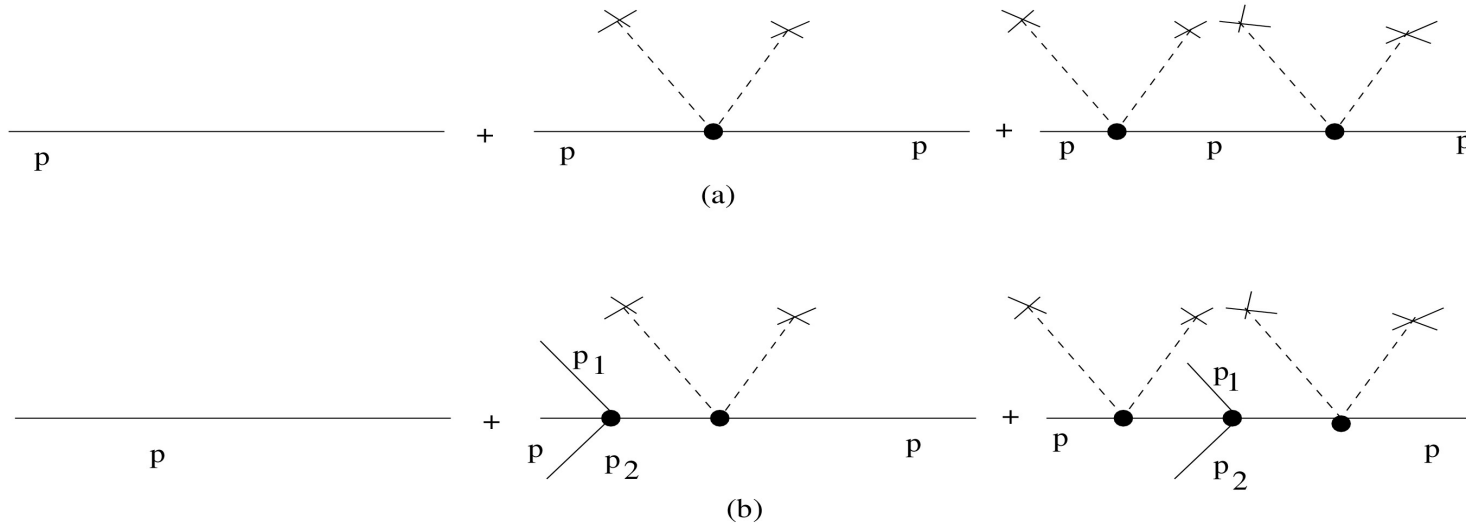
adjoint Higgs (deconfining phase)

(SU(2) → U(1))

- from effective action:

$$m_a^2 = -2e^2 \text{tr} [\phi, t_a][\phi, t_a] \xrightarrow{\text{unitary gauge}} m_1^2 = m_2^2 = 4e^2 \frac{\Lambda^3}{2\pi T}, m_3 = 0$$

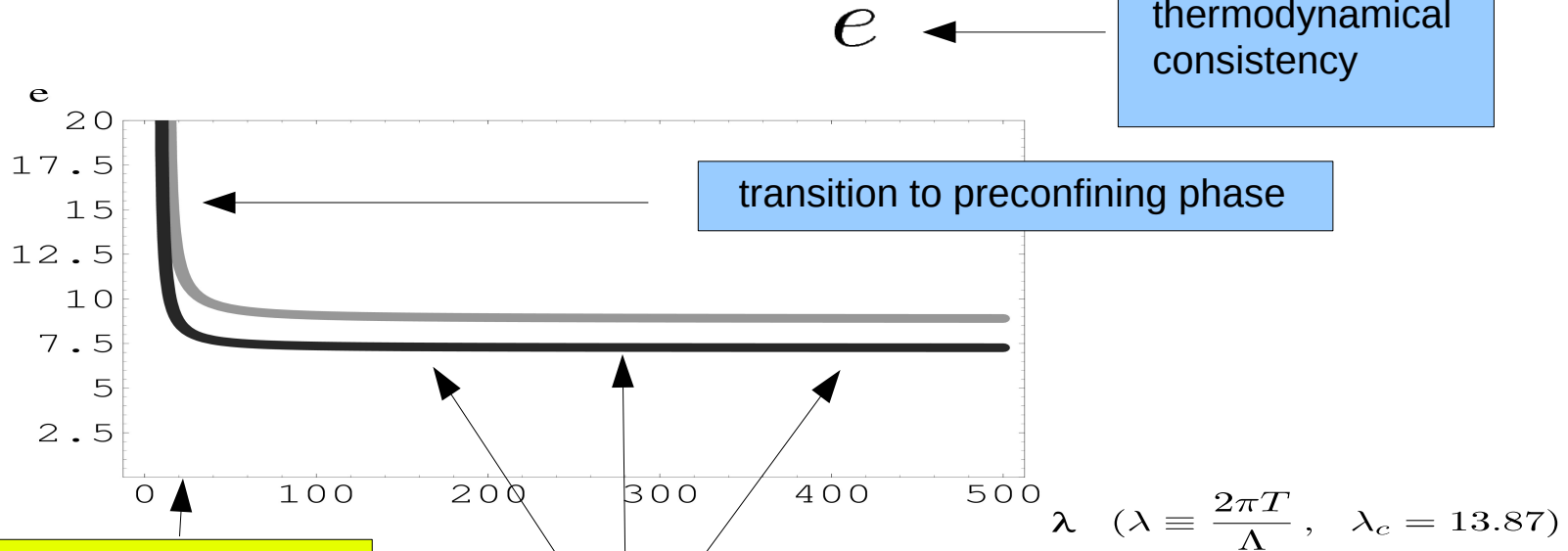
- no momentum transfer to ϕ , but this infinitely often
(Dyson series for mass generation):



- no off-shell propagation of massive modes
(otherwise: momentum transfer to ϕ !)

effective gauge coupling

- evolution of effective gauge coupling:



$$\lambda_c = \frac{2\pi T_c}{\Lambda} = 13.87$$

$$e = \sqrt{8\pi}$$

[Dolan et Jackiw (1974)]

coarse-graining dominated by $\rho \sim |\phi|^{-1}$

- restore \hbar

$$e = \frac{\sqrt{8\pi}}{\sqrt{\hbar}}$$

$$S_{C/A} = \hbar.$$

[Brodsky et al. (2011); Kaviani et Hofmann 2012, Hofmann (2012,2013)]

in Euclidean, fundamental (classical) theory value of unit of action not set, only after comparison with effective theory, where \hbar is determined to occur in certain powers at pointlike vertices do we know that

$$S_{C/A} = \hbar$$

electric-magnetically dual interpretation:

- if SU(2) something to do with photons (later!) then **electric-magnetically dual** interpretation required:
in units $c = \epsilon_0 = \mu_0 = k_B = 1$ fine-structure constant

$$\alpha = \frac{Q^2}{4\pi\hbar},$$

for α to be unitless:

$$Q \propto \frac{1}{e}.$$

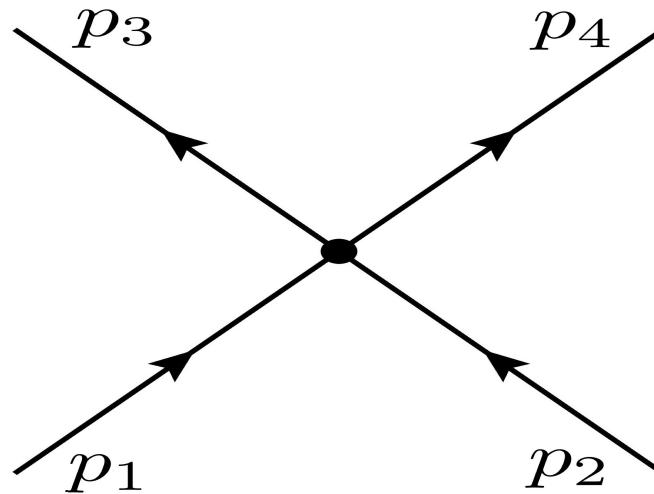
But: magnetic coupling
in SU(2)
(charge of a magnetic
monopole)

$$g = \frac{4\pi}{e}.$$

\Rightarrow **U(1) part of SU(2)** to be interpreted in an **electric-magnetically dual way**.
(e.g., magnetic monopole \longleftrightarrow electric monopole, etc.)

radiative corrections (deconfining phase)

- momentum transfer in effective 4-vertex (unitary-Coulomb gauge):



s-channel:

$$|(p_1 + p_2)^2| \leq |\phi|^2$$

t-channel:

$$|(p_1 - p_3)^2| \leq |\phi|^2$$

u-channel:

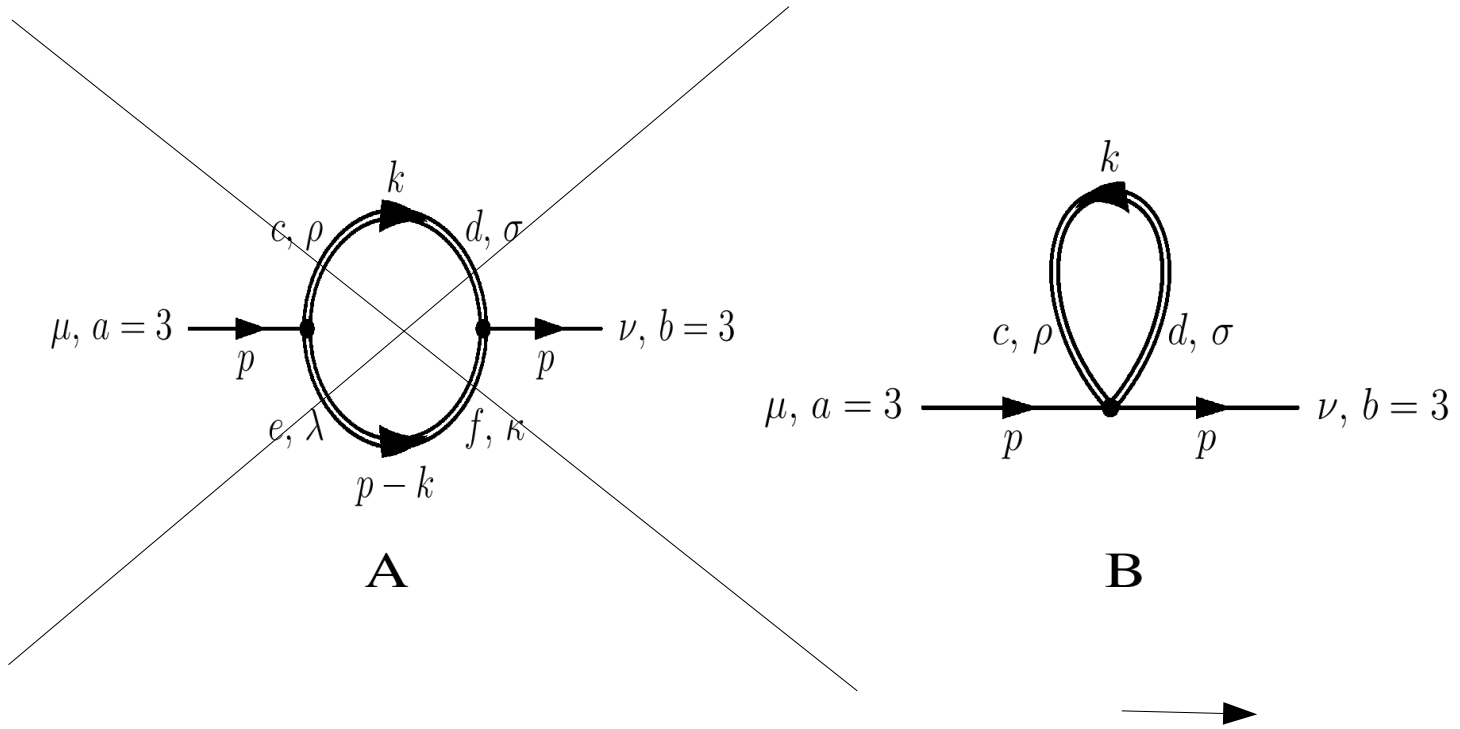
$$|(p_1 - p_4)^2| \leq |\phi|^2$$

- coherent average over all three Mandelstam channels \longrightarrow
thermodynamical quantities: 2-loop/1-loop ($<10^{-3}$), 3-loop/1-loop ($<10^{-7}$),
loop expansion into 1-PI diagrams probably terminates at finite order

[RH, 2006]

radiative corrections (deconfining phase)

- polarization tensor of massless mode:

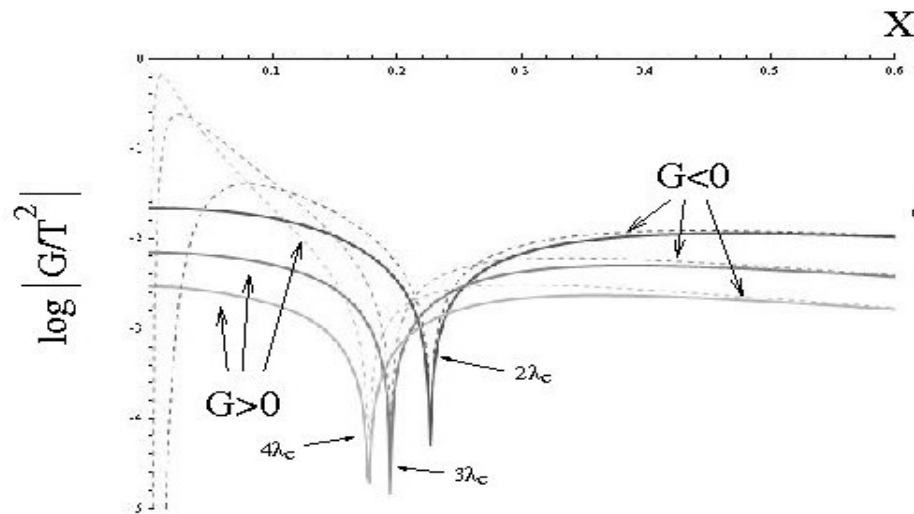


(excluded by kinematic constraints)

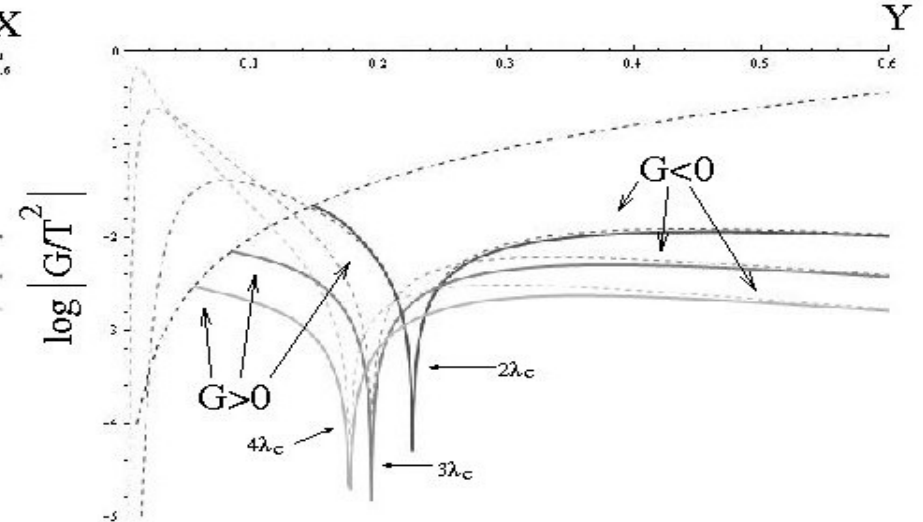
screening functions
 G, F (transverse, longitud.)
 as solutions of respective gap equations

radiative corrections (deconfining phase)

- transverse photons, screening function G :
[Schwarz et al. (2007), Ludescher & RH (2008)]



$$\left(X \equiv \frac{|\vec{p}|}{T} \right)$$



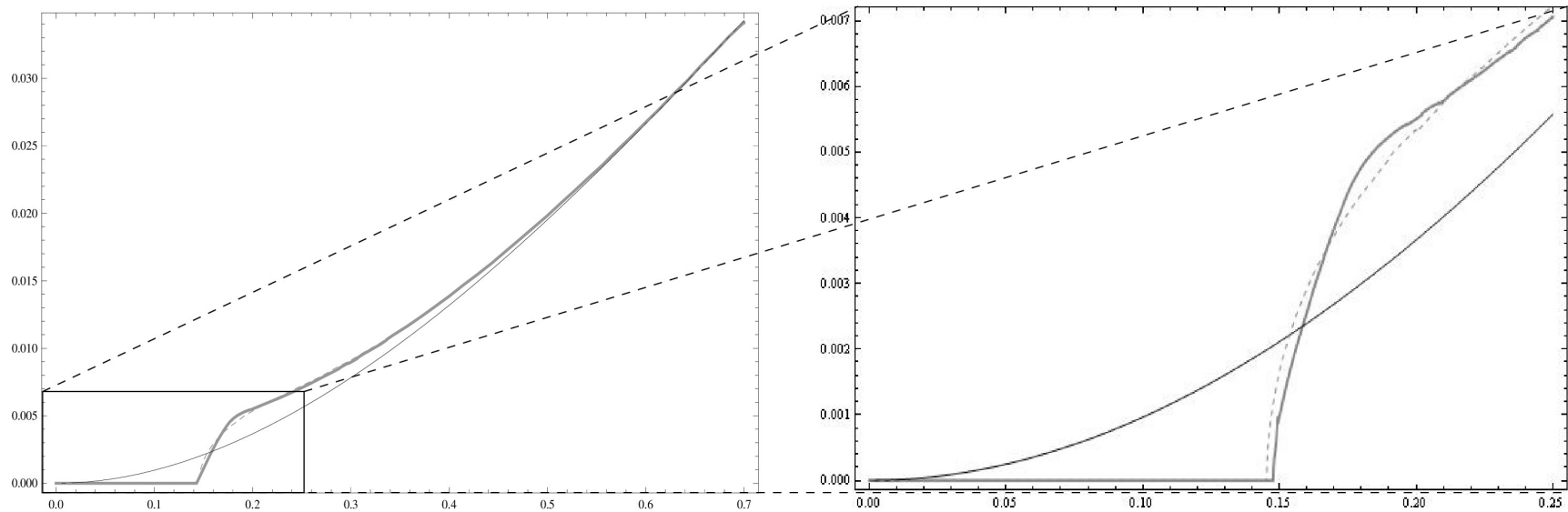
$$\left(Y \equiv \frac{\omega}{T} \right)$$

$$Y^2 = X^2 + \frac{G}{T^2}$$

(spectral) radiative corrections (deconfining phase)

- spectral distribution of energy density, massless mode – transverse propagation at $T = 2T_c$ in Rayleigh-Jeans regime

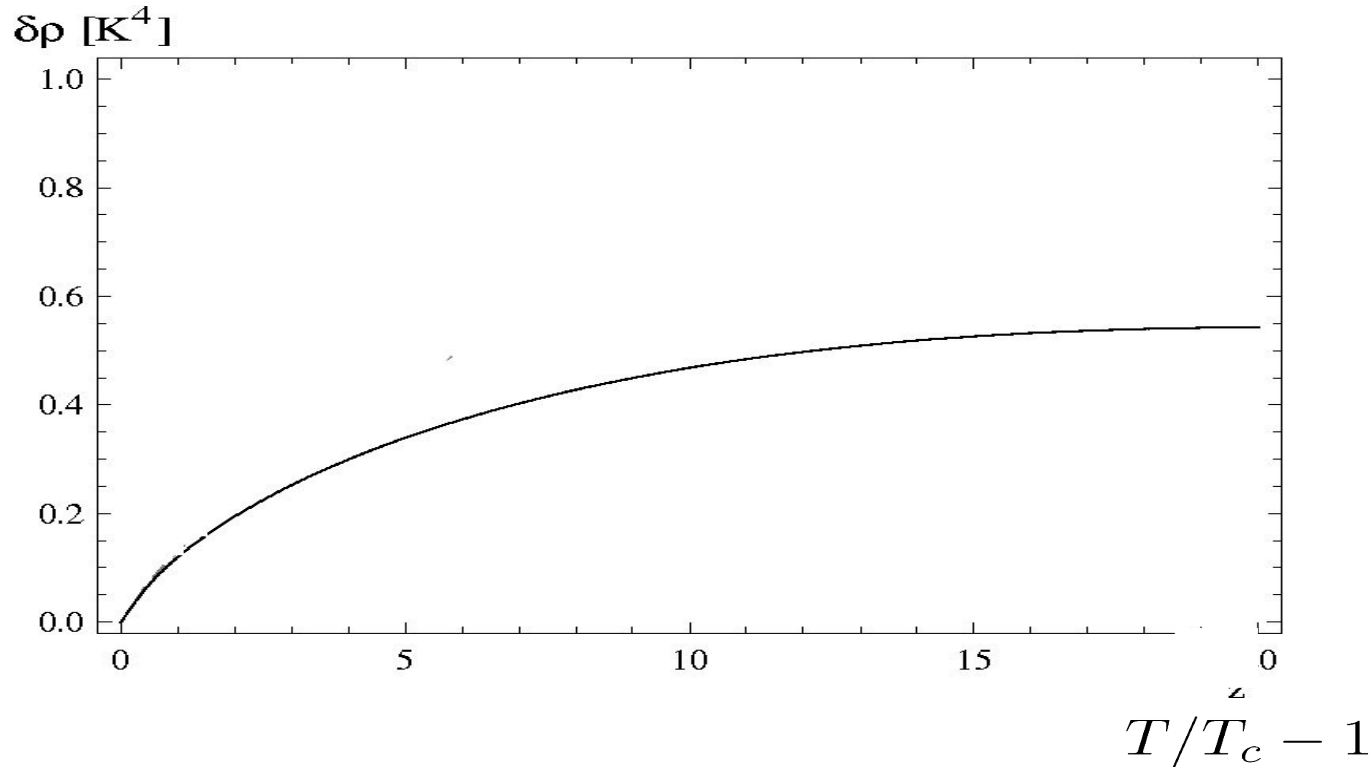
I/T^3



Y

(integrated) radiative corrections (deconfining phase)

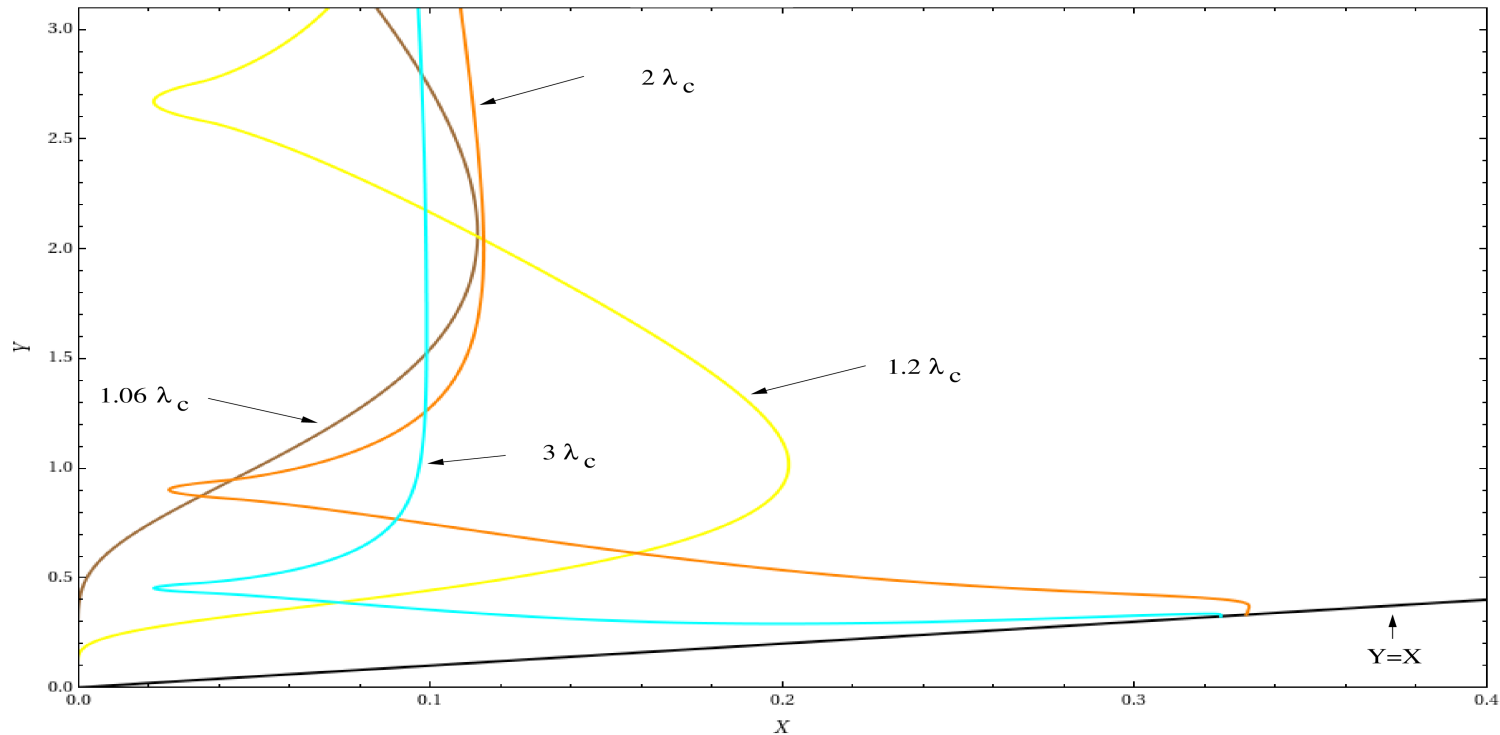
- difference between energy density of SU(2) and U(1),
massless mode – transverse polarisations



(**positive** slope \longleftrightarrow bias for **negative** temperature fluctuations, later!)

(spectral) radiative corrections (deconfining phase)

- low-momentum-support dispersion law, massless mode - longitudinal propagation



(charge-density waves: real-world magnetic modes,
intergalactic magnetic fields [Falquez et al (2011)])

SU(2) postulate for photon propagation

- What is T_c ?
- strong increase of CMB line temperature below $\nu = 3$ GHz

$$T(\nu) = T_0 + T_R \left(\frac{\nu}{\nu_0} \right)^\beta$$

[Fixsen et al. (2009),
Haslam et al. (1981),
Reich et Reich (1986),
Roger et al. (1999),
Maeda et al. (1999)]

where: $T_0 = 2.725$ K; $\nu_0 = 1$ GHz;
 $\beta = -2.62 \pm 0.04$; $T_R = (1.19 \pm 0.14)$ K.

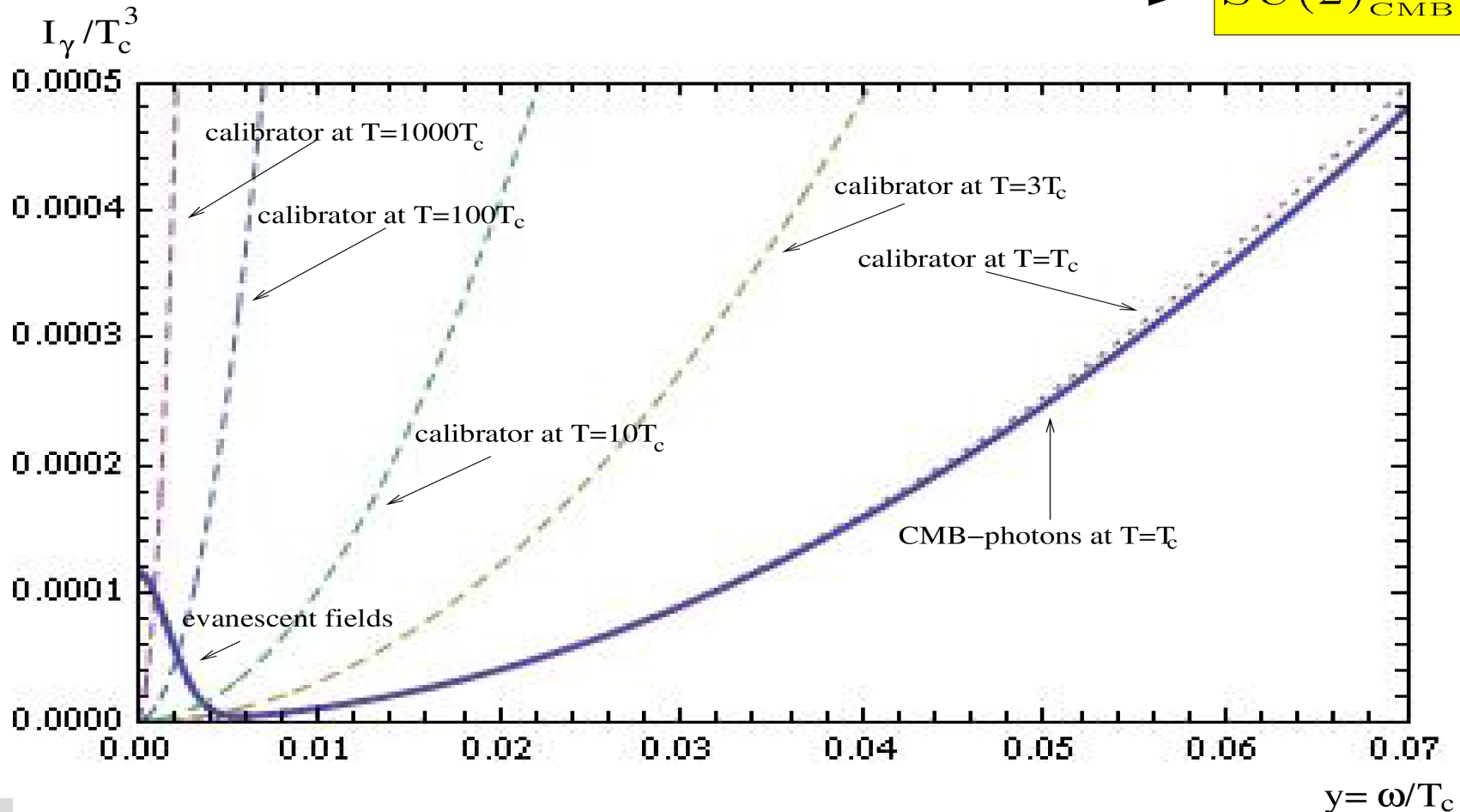
(radio-frequency surveys of CMB yield line temperatures as:

source	ν [GHz]	T [K]
Roger	0.022	21200 ± 5125
Maeda	0.045	4355 ± 520
Haslam	0.408	16.24 ± 3.4
Reich	1.42	3.213 ± 0.53
Arcade2	3.20	2.792 ± 0.010
Arcade2	3.41	2.771 ± 0.009 .)

evanescent low-frequency modes

- bump from evanescent modes ($\omega < m_\gamma$),
 m_γ photon Meissner mass (condensation of electric monopoles)
- T_c very close to present CMB temperature T_0 (onset of dec.-prec. PT)
 [Hofmann (2009)]

$SU(2)_{CMB}$



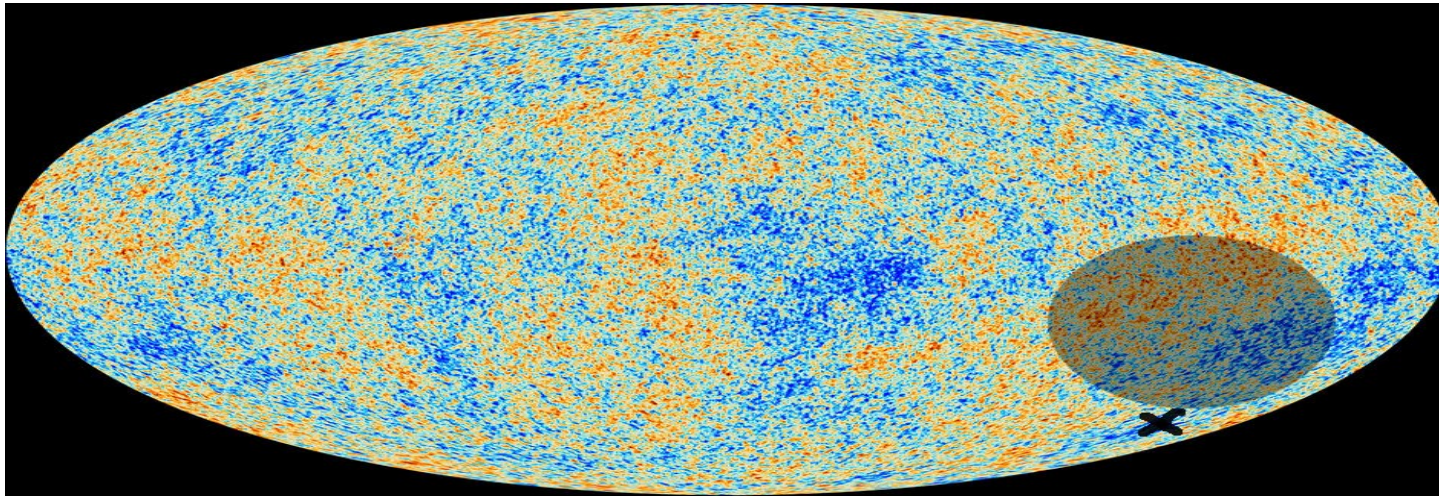
Yang-Mills scale of $SU(2)_{\text{CMB}}$:

$$T_c = \frac{13.87}{2\pi} \Lambda_{\text{CMB}} = 2.725 \text{ Kelvin} \sim 2 \times 10^{-4} \text{ eV}$$

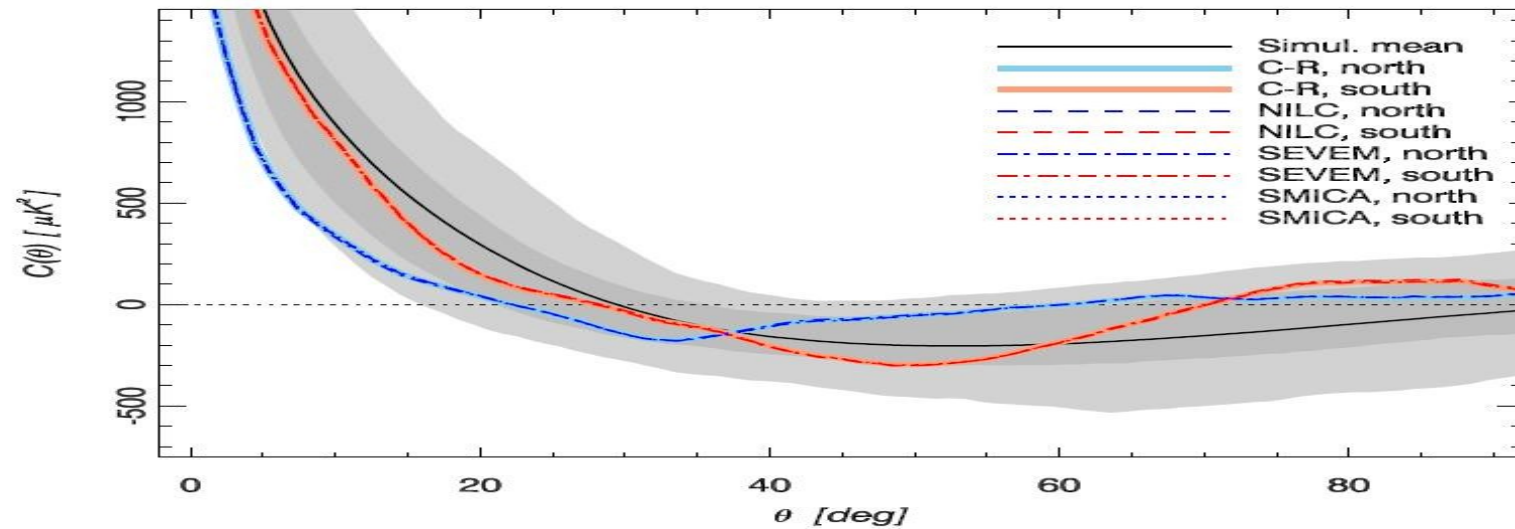
some CMB large-angle anomalies: WMAP and Planck

- dipolar power asymmetry (extends from $l = 2, \dots, 600$ in blocks of $\Delta l = 100$)
[Hansen et al. (2009), Ade et al. (2013), etc.]
- low variance on ecliptic North, associated with $l=2,3$
[Monteserin et al. (2008), Cruz et al., (2011), Ade et al. (2013), etc.]
- anomalous alignment of $l=2,3$ (3° - 9°)
[Tegmark et al. (2003), de Oliveira-Costa et al. (2004), Ade et al. (2013), etc.
(estimator of axis: maximum of angular momentum dispersion),
Copi et al. (2004), Schwarz et al. (2004), Bielewicz et al. (2005,2009), Copi et al. (2010), etc.
(multipole vector decomposition)]
- cold spot ($-73\mu\text{K}@4^\circ$; $-20\mu\text{K}@10^\circ$; $l,b=207.8^\circ,-56.3^\circ$)
[Viola et al. (2004), Cruz et al. 2005, Rudnick et al. (2007), Ade et al. (2013), etc.]
- hemispherical asymmetry
(for $l=2$ - 40 max. larger power on hemisphere $l,b=237^\circ,-20^\circ$)
[Eriksen et al. (2004), Hansen et al. (2004), Park (2004), Ade et al. (2013), etc.]
- mirror parity violation (plane of max. antisymmetry: $l,b=262^\circ,-14^\circ$)
[Finelli et al.(2012); Ben-David et al. (2012), etc.]
- **suppression of $\langle TT \rangle(\theta) \equiv C(\theta)$ for $\theta \geq 60^\circ$ on ecliptic North**
[Spergel et al. (2003), Copi et al. (2004,2007,2009,2010), Ade et al. (2013), etc.]

cold spot



TT suppression on ecliptic North



successful phenomenological attempt at explanation: multiplicative, dipolar modulation model

[Gordon et al. (2005), Eriksen et al. (2007), Hoftuft et al. (2009), Ade et al. (2013)]

$$\vec{d}(\vec{n}) = (1 + A\vec{p} \cdot \vec{n})\vec{s}_{\text{iso}} + \vec{n}$$

dipole amplitude

dipole direction

instrumental noise

isotropic CMB sky

maximum likelihood at: $A \sim 0.07$; $l_p \sim 220^\circ$; $b_p \sim -21^\circ$

- robust against change of foreground treatment and experiment
(WMAP, Planck)

- comparison with CMB cold spot: $l_{cs} \sim 207.8^\circ$; $b_{cs} \sim -56.3^\circ$

$$\Rightarrow \angle \vec{p}, \vec{e}_{cs} \sim 36^\circ$$

dynamical breaking of statistical isotropy:

- integrated blackbody anomaly due to $SU(2)_{\text{CMB}}$:

◆ $\delta\rho(T) \equiv \rho_{SU(2)_{\text{CMB}}} - \rho_{U(1)}$

◆ $T = \bar{T}(t) + \delta T(t, \vec{x})$

(Silk cutoff)

◆ $SU(2)_{\text{CMB}}$ bias factor $F(\bar{T}, \delta T)$ for δT in phys. voxel volume $\Delta V \sim \frac{(2\pi a_s)^3}{k_s^3}$

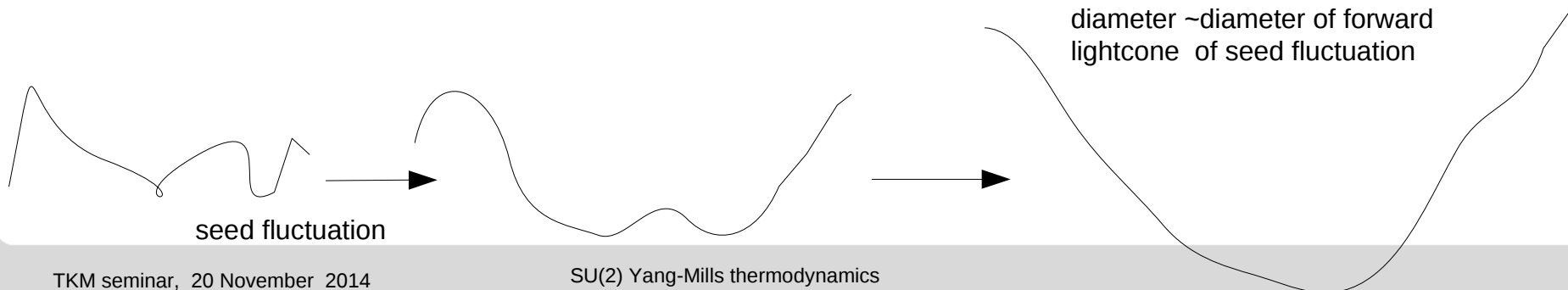
$F(\bar{T}, \delta T) = \frac{P_{SU(2)}}{P_{U(1)}}$

where

$P = \frac{\exp(-\rho\Delta V/\bar{T})}{\int_{T_0}^{\infty} dT \exp(-\rho\Delta V/T)}$

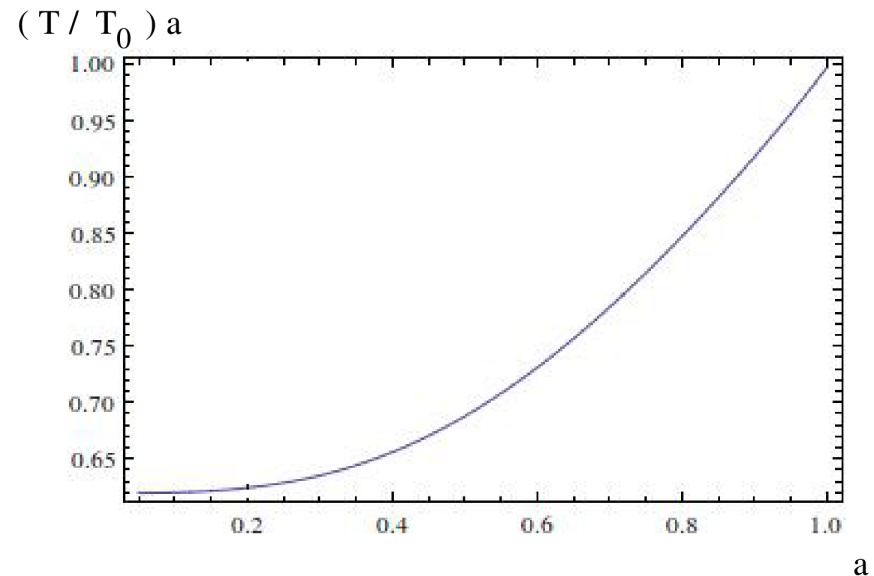
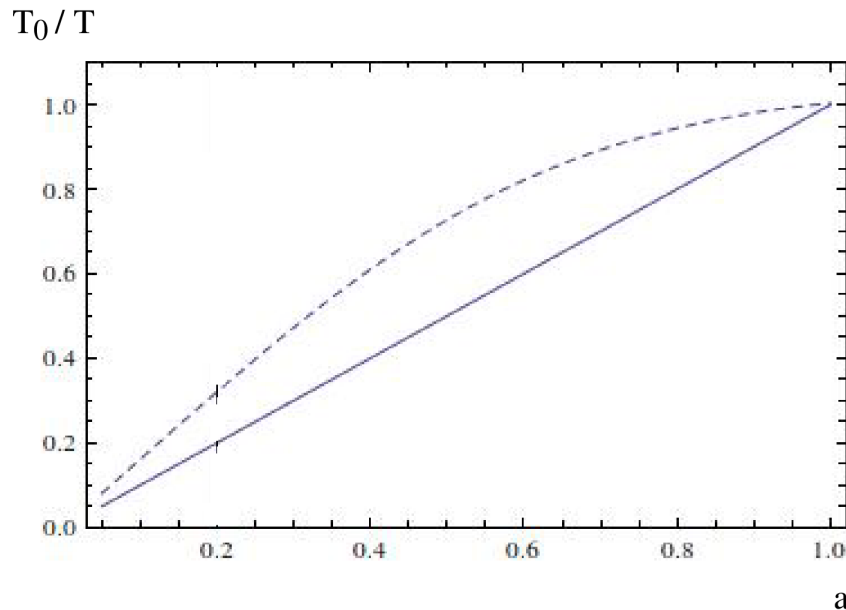
Since slope of $\delta\rho$ positive \Rightarrow negative δT favoured!

(dynamical breaking of statistical isotropy, RH Nature Phys. 2013)



Violation of conformal scaling $T \propto a^{-1}$

(a scale factor in FRLW cosmology, $a = 1$ today) [RH, 2014]



$$\Rightarrow T = 0.62 a^{-1} \times T_0, \quad \left(a \leq \frac{1}{10}\right)$$

\Rightarrow discrepancy between redshifts for instantaneous reionisation extracted by observation of Gunn-Peterson trough in quasar spectra and from CMB TT angular spectrum (Planck) repealed

(only tree-level properties of effective SU(2), thermal ground state + massive modes and photon, are required!)

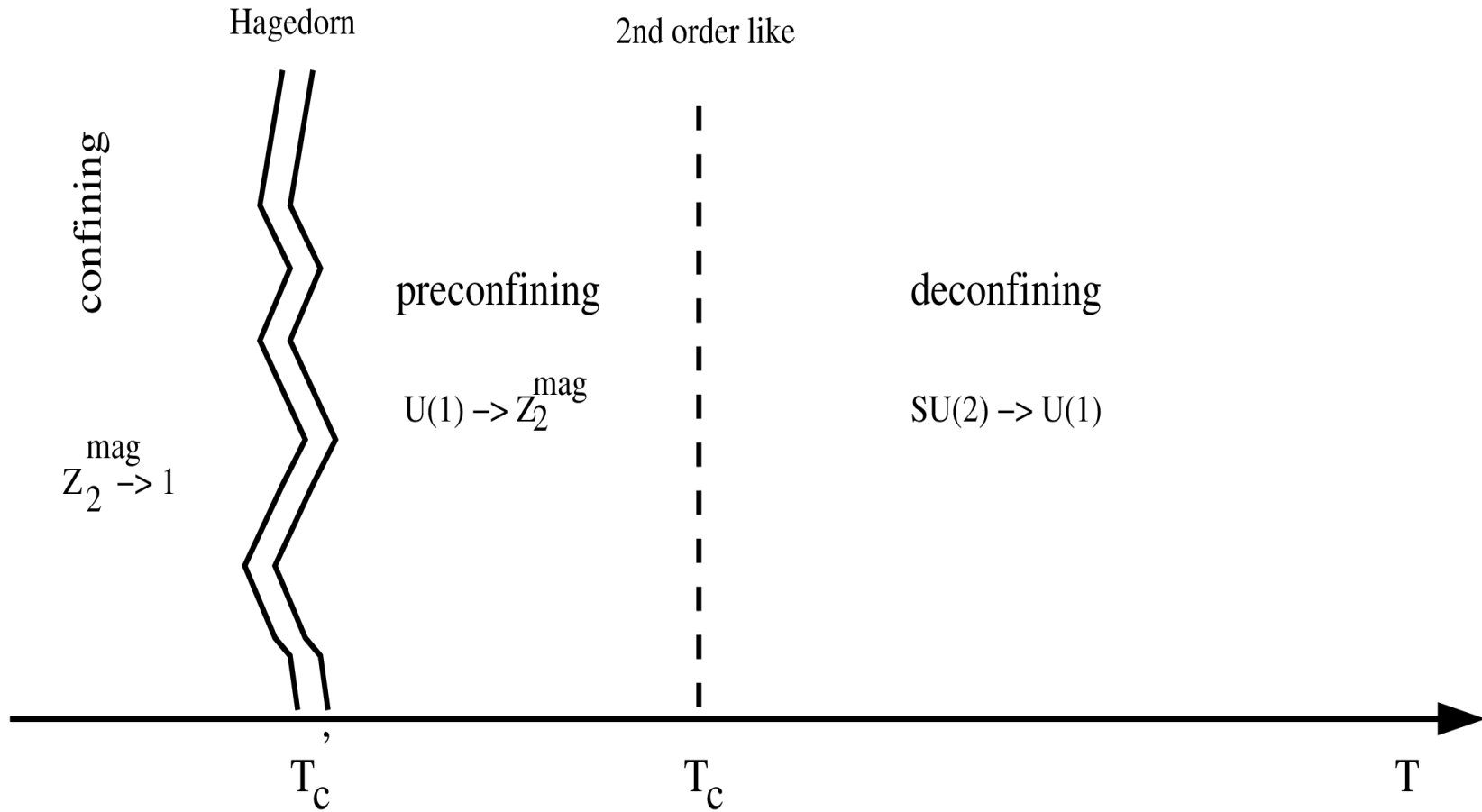
cosmic neutrinos:

⇒ to have $N_{\text{eff}} = 3.36$ (Planck) today, cosmic neutrinos could be **thermal quasi-particles** whose temperature **coincides** with the CMB temperature and whose universal **mass is induced** by interaction with the CMB,

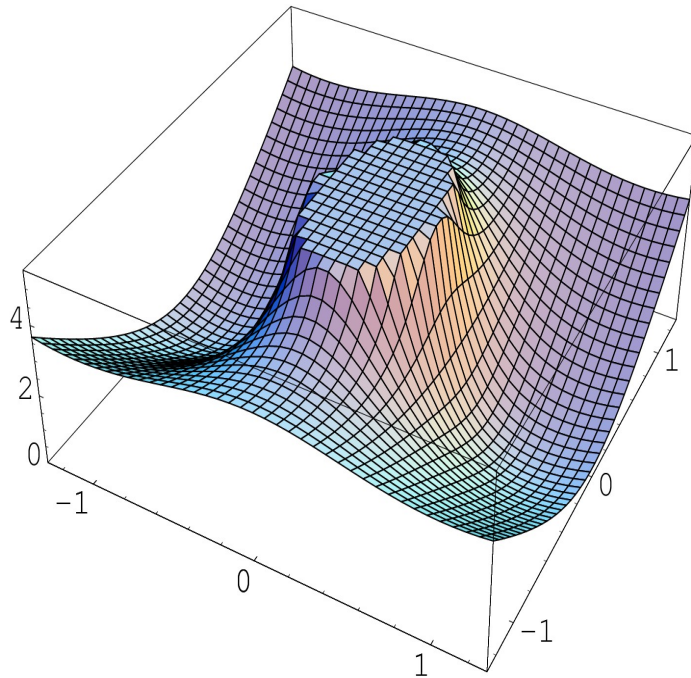
$$m_\nu = \xi T \quad (*)$$

with $\xi = 3.973$ [RH, 2014]

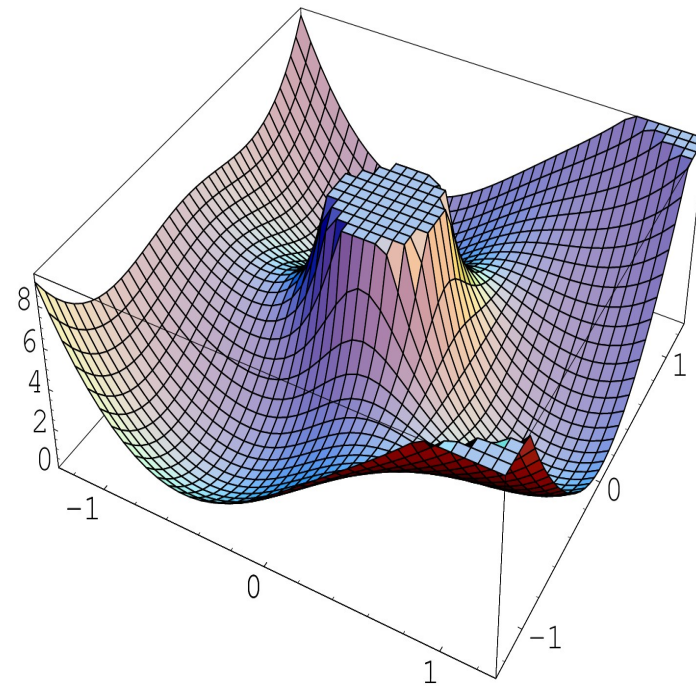
⇒ (*) is expected if neutrinos are single center-vortex loops in the confining phase of (other) SU(2) Yang-Mills theories



Hagedorn transition



SU(2)



SU(3)

- at T'_c no propagating gauge modes, decay of ground state described by local magnetic center jumps ('t Hooft loop relaxes to minimum of potential by creating units of center flux) [Scheffler, Stamatescu, RH 2007; RH 2007]

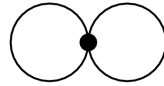
spectrum and pressure, confining phase

n=0:
(N₀=1)



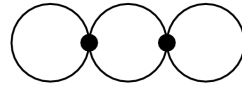
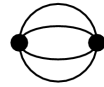
(neutrinos)

n=1:
(N₁=1)



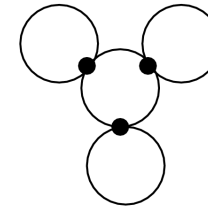
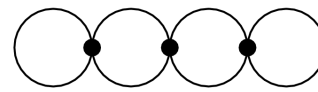
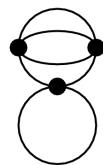
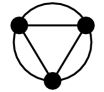
(massive leptons)

n=2:
(N₂=2)



(unstable in presence of photons)

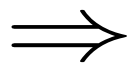
n=3:
(N₃=4)



- multiplicity = number of bubble diagrams in $\lambda\phi^4$ theory
- Borel resummation of thermodynamical pressure [Bender & Wu, 1969, 1971]

$$P_{\text{as}} \sim \frac{\Lambda^4}{2\pi^2} \hat{\beta}^{-4} \left(\frac{7\pi^4}{180} + \sqrt{2\pi} \hat{\beta}^{\frac{3}{2}} \sum_{l=0}^L a_l \sum_{n \geq 1} (32\lambda)^n n! n^{\frac{3}{2}+l} \right), \quad (\hat{\beta} \equiv \frac{\Lambda}{T}, \lambda \equiv \exp(-\hat{\beta})) .$$

analytic continuation and inverse Borel

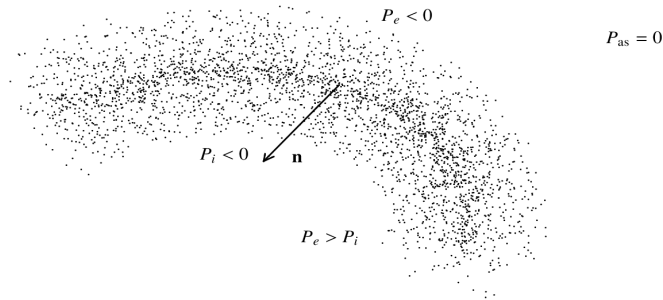


pressure is not thermodynamic but contains imaginary contamination which outgrows real part with increasing „temperature“

[RH 2007]

evolving $n=0$ and $n=1$ center vortex loops

(J. Moosmann and RH, 2008)



- shrinking of vortex loop by 1-dim analog of Ricci flow (curve shrinking equation):

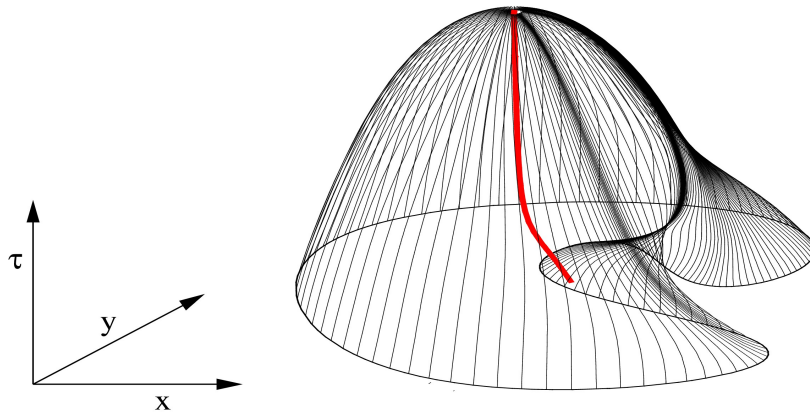
$$\partial_{\tau} \vec{x} = \frac{1}{\sigma} \partial_s^2 \vec{x},$$

where \mathcal{S} arc length, σ string tension, and \mathcal{T} (dimensionless) flow parameter

- upon re-scaling $\hat{x} \equiv \sqrt{\sigma} \vec{x}, \xi = \sqrt{\sigma} s \Rightarrow$

$$\partial_{\tau} \hat{x} = \partial_{\xi}^2 \hat{x}$$

evolving n=0 center vortex loops



n=0 CVL shrinks into
round point after finite
„time“

$$\tau = T$$

[Gage & Hamilton 1886, Grayson 1987]

- consider an ensemble of n=0 CVLs of same area A_0
- factorise effective „action“ into a conformal and nonconformal part
 $S = F_c \times F_{nc}$ in accord with 2D Euclidean point symmetry
 and demand conformal invariance for curve length $L \rightarrow \infty$, e.g.

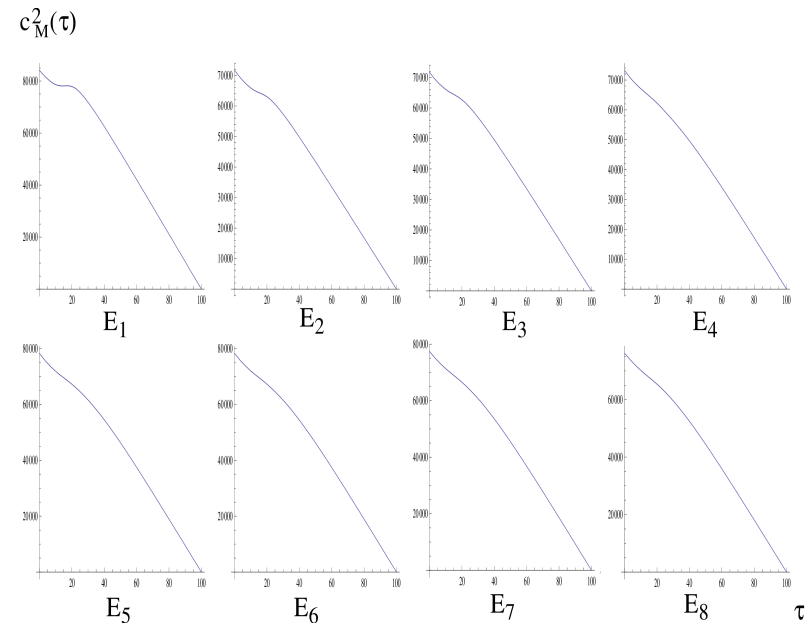
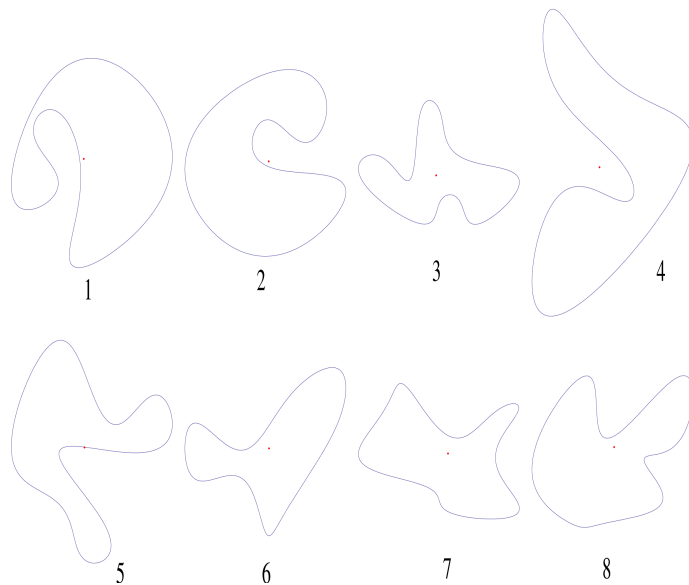
$$S = \frac{L(\tau)^2}{A(\tau)} \left(1 + \frac{c(\tau)}{L(\tau)} \right)$$

isoperimetric ratio

- determine $c(\tau)$ from $\frac{d}{d\tau} Z = 0$ under curve shrinking

evolving $n=0$ center vortex loops, cntd.

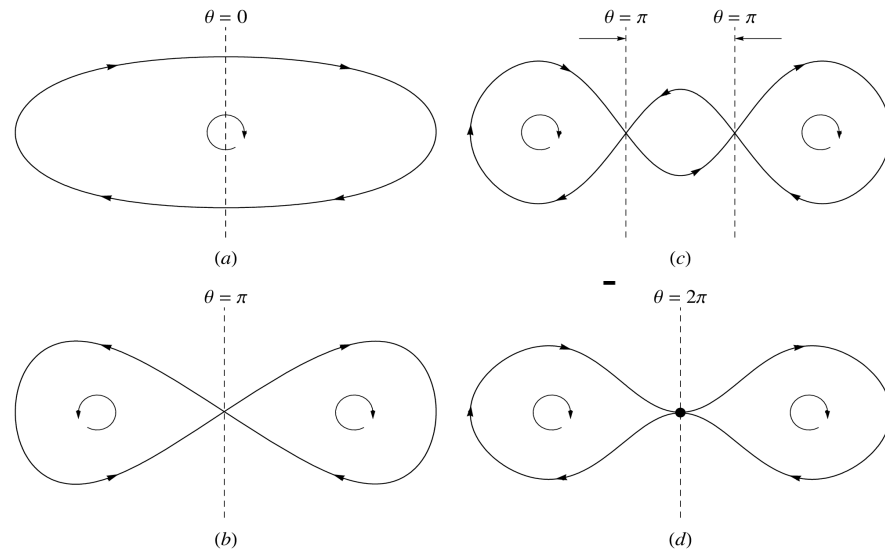
- $c(\tau)$ acts as order parameter for restoration of conformal symmetry when limit of round points is reached



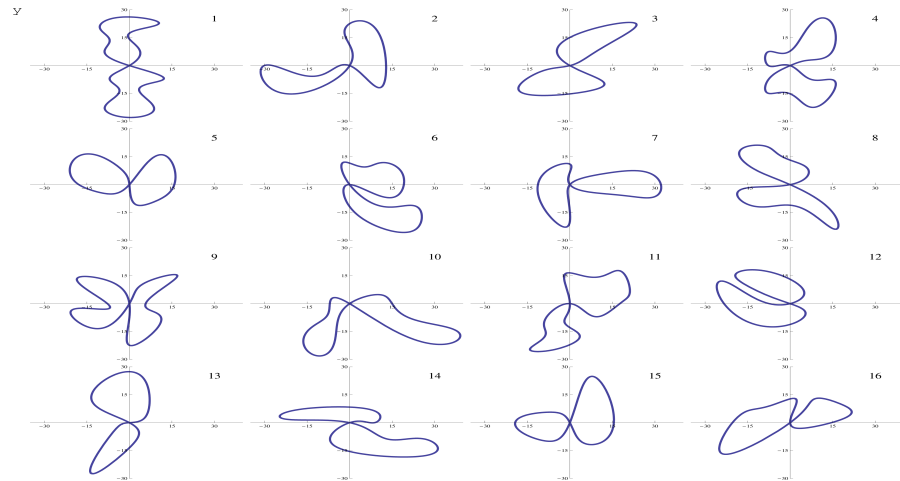
- even in sense of an ensemble average, $n=0$ CVLs disappear from spectrum at **finite resolution** (mass gap)

evolving n=1 center vortex loops

- topological transition from n=0 to n=1 by twisting and pinching

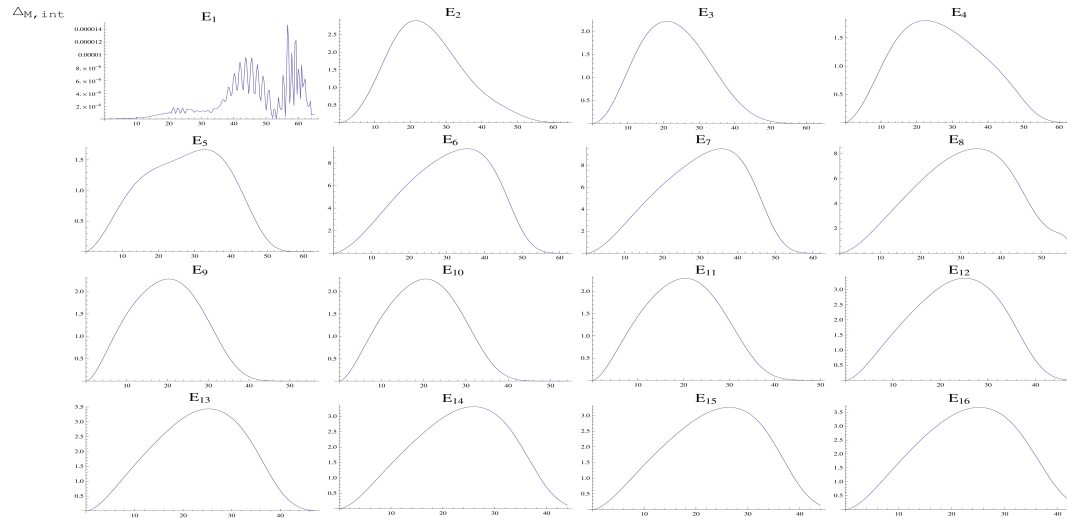


- ensemble of n=1 CVLs

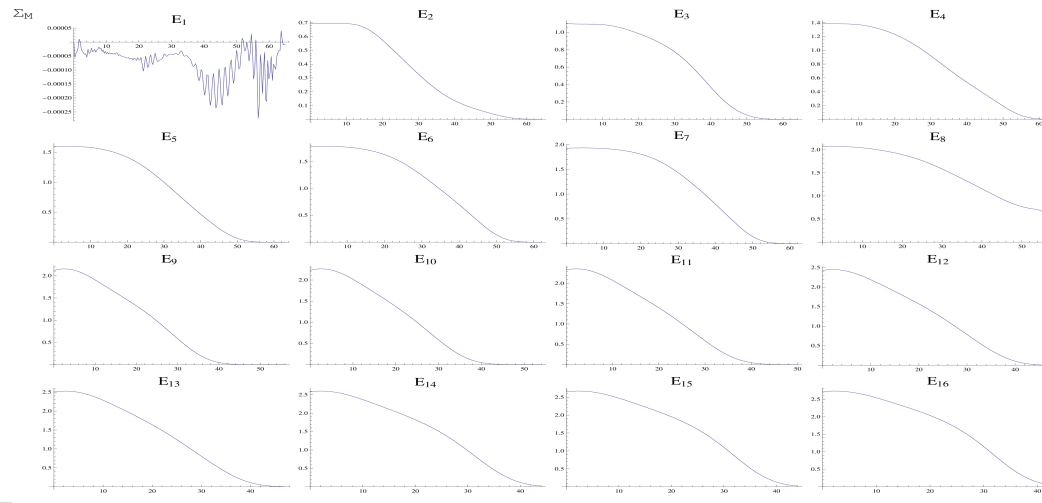


evolving $n=1$ center vortex loops, cntd.

- evolution of variance of position of intersection point

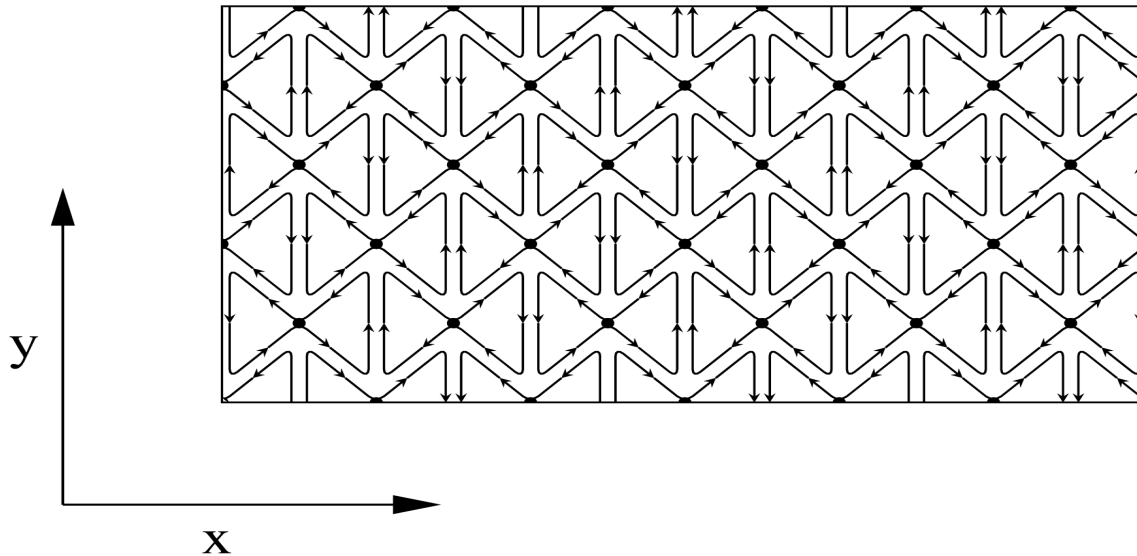


- evolution of entropy



evolving $n=1$ center vortex loops, cntd.

- possible configuration of planar electrons in high T_c superconducting state?

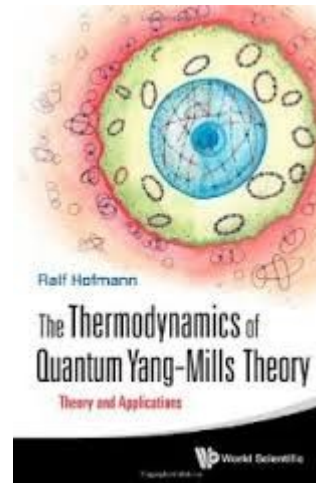


Summary

- SU(2) thermodynamics nonperturbatively:
caloron, thermal ground state, adjoint Higgs mechanism, caloron action= \hbar
- blackbody anomaly:
thermal photon dispersion, critical temperature for dec.-prec. PT from low-frequency spectral anomaly (Arcade2, terrestrial radio-frequency CMB observations), longitudinal (magnetic) charge density waves
- CMB large-angle anomalies:
Yang-Mills favours **negative temperature fluctuations**, semiquantitative model, cosmic neutrinos and relativistic vector modes
- confining phase: nonthermal Hagedorn transition, $n=0,1$ as stable solitons, curve shrinking as a renormalisation group flow, mass gap for $n=0$, collapse of ensemble onto one member $n=1$ (entropy vanishes at finite resolution)

Thank you.

more info in:



(World Scientific, 2011)

and:

- R. Hofmann, “ Relic photon temperature versus redshift and the cosmic neutrino background“, arxiv: 1407.1266 [physics, hep-th], subm. Ann. d. Physik
- R. Hofmann, “The fate of statistical isotropy”, Nature Physics **9**, 686-689 (2013).
- N. Krasowski and R. Hofmann, “Evidence for absence of one-loop photon-photon scattering in a thermal, deconfining SU(2) Yang-Mills plasma”, Annals of Physics **347**, 287-308 (2014); arxiv:1301.4716 [physics, hep-th]
- R. Hofmann and D. Kaviani, “The quantum of action and finiteness of Radiative corrections: Deconfining SU(2) Yang-Mills thermodynamics”, Quantum Matter **1**, 41-52 (2012)