Nuclear Forces and the Renormalization Group

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Strong interaction physics in the lab and cosmos

Matter at the extremes:

density $\rho \sim \! 10^{11} \ldots 10^{15} \text{ g/cm}^3$

neutron-rich to proton-rich Z/N ~0.05...0.6

temperatures T $\sim ...30$ MeV





Interaction challenges: $QCD \Rightarrow chiral EFT \Rightarrow RG evolved$ low-momentum interactions for all nuclei Many-body challenges

Astrophysics challenges

Outline

Resolution dependence of nuclear forces

Effective field theory and the renormalization group

Many-body developments

Λ / Resolution dependence of nuclear interactions

with high-energy probes: quarks+gluons at low energies: complex QCD vacuum

lowest energy excitations: pions, nearly massless, m_{π} =140 MeV





 Λ_{chiral} momenta Q ~ λ^{-1} ~ m_{π}

 $\Lambda_{\text{pionless}}$ Q << m_{\pi}=140 MeV

Λ / Resolution dependence of nuclear interactions

with high-energy probes: quarks+gluons

cf. scale/scheme dependence of parton distribution functions



Lattice QCD

Effective theory for NN, many-N interactions, operators depend on resolution scale Λ

 $H(\Lambda) = T + V_{NN}(\Lambda) + V_{3N}(\Lambda) + V_{4N}(\Lambda) + \dots$

 Λ_{chiral} momenta Q ~ λ^{-1} ~ m_{π}: chiral effective field theory nucleons interacting via pion exchanges + contact interactions typical Fermi momenta in nuclei ~ m_{π}



 $\Lambda_{\text{pionless}}$ Q << m_{\pi}=140 MeV: pion not resolved pionless effective field theory large scattering lengths + corrections



applicable to loosely-bound, dilute systems, reactions at astro energies

Lattice QCD and nuclear forces





Weinberg, van Kolck, Kaplan, Savage, Wise, Epelbaum, Meissner, Nogga, Machleidt,...



Weinberg, van Kolck, Kaplan, Savage, Wise, Epelbaum, Meissner, Nogga, Machleidt,...

Chiral EFT highlights



Navratil et al. (2007)

impressive agreement, highlights the importance of 3N interactions

Low-momentum interactions from the Renormalization Group evolve to lower resolution/cutoffs by integrating out high-momenta, can be carried out exactly for NN interactions Bogner, Kuo, AS (2003)



implemented by RG equations or unitary transformation

Low-momentum interactions from the Renormalization Group evolve to lower resolution/cutoffs by integrating out high-momenta, can be carried out exactly for NN interactions Bogner, Kuo, AS (2003)



method to vary resolution scale without loss of low-energy NN physics

Chiral EFT and RG



Weinberg eigenvalue diagnostic

study spectrum of $G_0(z)V |\Psi_{\nu}(z)\rangle = \eta_{\nu}(z) |\Psi_{\nu}(z)\rangle$ at fixed energy z governs convergence $T(z) |\Psi_{\nu}(z)\rangle = (1 + \eta_{\nu}(z) + \eta_{\nu}(z)^2 + ...) V |\Psi_{\nu}(z)\rangle$ can write as Schrödinger eqn $\left(H_0 + \frac{1}{\eta_{\nu}(z)}V\right) |\Psi_{\nu}(z)\rangle = z |\Psi_{\nu}(z)\rangle$

high momenta/large cutoffs lead to flipped-potential bound states of - λV for small λ /large $\eta \Rightarrow$ Born series always nonperturbative with cores

Repulsive core eigenvalues small for lower cutoffs





sharp or smooth cutoff $V_{low k}(\Lambda)$ from RG equations or unitary transformations smooth cutoff $V_{SRG}(\lambda)$ evolves towards band diagonal Bogner et al. (2007)

high momenta decouple

Similarity RG interactions

Unitary transformations to band-diagonal $V_{srg}(\lambda)$ from flow equations Glazek, Wilson (1993), Wegner (1994)

$$\frac{dH_s}{ds} = [\eta_s, H_s] = [[G_s, H_s], H_s]$$

with flow operator $G_s=T_{rel}$ and resolution $\lambda=s^{-1/4}$ Bogner et al. (2007)

$$\frac{dV_s(k,k')}{ds} = -(k^2 - k'^2)^2 V_s(k,k') + \frac{2}{\pi} \int_0^\infty q^2 dq \left(k^2 + k'^2 - 2q^2\right) V_s(k,q) V_s(q,k')$$

intermediate momenta $k > k_{max} \sim \lambda$ decouple for low energies





weakens off-diag coupling



Block diagonalization using SRG





low-momentum blocks very similar to $V_{low k}$

formal equivalence? connections to EFT?



Advantages of low-momentum interactions for nuclei

high momenta/large cutoffs lead to slow convergence for nuclei

evolution of chiral EFT interactions to low-momentum beneficial



weakens off-diagonal coupling in HO states

lower cutoffs need smaller basis Bogner et al. (2007)

$$10^3$$
 states for N_{max}=2 vs.
 10^7 states for N_{max}=10



direct convergence in structure calcs

Pushing the limits



Impact on binding energies

 $V_{low k}(\Lambda)$ defines class of NN interactions with cutoff-independent low-energy NN observables

cutoff variation estimates errors due to neglected parts in $H(\Lambda)$

cutoff dependence explains "Tjon line", 3N required by renormalization

large scattering lengths drive correlation Platter et al. (2005)

35

30

25

20 7.5 A=3

B, [MeV]

8

8.5

 B_{α} [MeV]



Towards evolving 3N interactions

SRG evolution for 1d systems with T_{rel} in Jacobi HO basis Jurgenson et al., in prep.



gearing up to evolve chiral 3N interactions

for now: use chiral EFT is complete basis

Low-momentum 3N interactions

from leading N²LO chiral EFT ~ $(Q/\Lambda)^3$ van Kolck (1994), Epelbaum et al. (2002)



$$= \frac{1}{2} E(\boldsymbol{\tau}_1 \cdot \boldsymbol{\tau}_3) + \text{perm.}$$



c₃,c₄ important for structure, large uncertainties at present

chiral EFT is complete basis \rightarrow 3N up to truncation errors D term could be fixed by tritium beta decay

Low-momentum 3N fits

fit D,E couplings to A=3,4 binding energies for range of cutoffs

linear dependences in fits to triton binding

3N interactions perturbative for $\Lambda \leq 2 \, \mathrm{fm}^{-1}$ Nogga, Bogner, AS (2004)

 $\left< V_{_{\rm JN}} \right> / \left< V_{_{\rm low}\,k} \right>$

0.01

 $\Lambda = 1.6 \text{ fm}^{-1}$

 $\Lambda = 1.9 \text{ fm}^{-1}$

 $\Lambda = 2.1 \text{ fm}^{-1}$

⁴He

 ^{3}H

 $\Lambda = 2.3 \text{ fm}^{-1}$

nonperturbative at larger cutoffs cf. chiral EFT $\Lambda \approx 3 \text{ fm}^{-1}$

3N exp. values natural $\sim (Q/\Lambda)^3 V_{NN} \sim 0.1 V_{NN}$



 $k_{F} [fm^{-1}]$

1.2

1.0

Subleading chiral EFT 3N interactions

parameter-free $N^3LO \sim (Q/\Lambda)^4$ Status from Epelbaum @ TRIUMF 3N workshop (2007)

rich operator structure (includes spin-orbit interactions)

I-loop diagrams with all vertices from $\mathcal{L}^{(0)}_{eff}$

 2π – exchange

The calculated corrections simply shift the LECs c_i as follows:

$$\delta c_1 = \frac{g_A^2 M_\pi}{64\pi F_\pi^2} \sim 0.13 \text{ GeV}^{-1} \qquad \delta c_3 = \frac{3g_A^4 M_\pi}{16\pi F_\pi^2} \sim 2.5 \text{ GeV}^{-1} \qquad \delta c_4 = -\frac{g_A^4 M_\pi}{16\pi F_\pi^2} \sim -0.85 \text{ GeV}^{-1}$$

 2π - 1π – exchange

$$\oint = \oint + \frac{1}{2} + \frac{1}{$$

ring diagrams

contact- 1π – exchange

$$\bigvee_{i=1}^{n} = \bigvee_{i=1}^{n} + \bigvee_{i=1}^{n} + \bigvee_{i=1}^{n-1} + \bigvee_{i=1}^{n-1} + \bigvee_{i=1}^{n-1} + \cdots$$

contact- 2π – exchange

$$\bigvee_{i=1}^{n} = \bigvee_{i=1}^{n-1} + \bigvee_{i=1}^{n-1} + \bigvee_{i=1}^{n-1} + \bigvee_{i=1}^{n-1} + \cdots$$

Theoretical uncertainties

Cutoff variation estimates errors due to neglected parts in $H(\Lambda)$



from A. Nogga

atomic EDMs

Advantages of low-momentum interactions for nuclei conventional G matrix approach does not solve off-diagonal coupling, renders Bethe-Brueckner-Goldstone expansion necessarily nonperturb.



Possibility of perturbative nuclear matter with NN and 3N start from chiral EFT to given order, soften with RG nuclear matter converged at \approx 2nd order, motivated by Weinberg eigenvalue analysis



Possibility of perturbative nuclear matter with NN and 3N start from chiral EFT to given order, soften with RG nuclear matter converged at \approx 2nd order, motivated by Weinberg eigenvalue analysis



Possibility of perturbative nuclear matter with NN and 3N

start from chiral EFT to given order, soften with RG

nuclear matter converged at \approx 2nd order, motivated by Weinberg eigenvalue analysis

reduced cutoff dependence at low densities, 3N drives saturation Bogner, AS, Furnstahl, Nogga (2005) + improvements, in prep.



overy through Advanced Computin

provides guidance to UNEDF http://unedf.org

Neutron matter from NN and 3N



uncertainties from c_i overwhelm errors due to cutoff variation, mainly c_3 for neutron matter

combine with knowledge of basic nuclear properties

important for dense matter in astrophysics

neutron star mergers →gravitational waves





Towards a universal nuclear energy density functional creation of heavy elements in r(-apid neutron capture) process

requires understanding highly neutron-rich systems need to improve extrapolations in masses





Bender et al. (2003)

masses and ground state properties from density functional theory

based on densities not wave functions from Kohn's 1998 Nobel Prize lecture

I begin with a provocative statement. In general the many-electron wavefunction $\Psi(r_1,...,r_N)$ for a system of N electrons is not a legitimate scientific concept, when $N \ge N_0$, where $N_0 \approx 10^3$.

I will use two criteria for defining "legitimacy": a) That Ψ can be calculated with sufficient accuracy and b) can be recorded with sufficient accuracy.

Nuclear masses and pairing

first microscopic pairing functional from low-momentum interactions Lesinski, Duguet, arXiv:0711.4386 and in prep.



Density functional RG for nuclei



$$S_{oldsymbol{\lambda}}[\psi^{\dagger},\psi] = \int \psi^{\dagger} \Big[\partial_t - rac{1}{2m} \Delta + (1-oldsymbol{\lambda}) oldsymbol{U}_{oldsymbol{\lambda}} \Big] \psi + rac{1}{2} \, \int \psi^{\dagger} \psi oldsymbol{\lambda} oldsymbol{V}_{ ext{low k}} \psi^{\dagger} \psi$$



Example: RG flow of the ground-state density for 1d model ("smeared-out" δ -function interaction)

+ (λV_{3N} will be included later)

density basis expansion scales favorably to large systems

gs energy and density from microscopic nuclear interactions

results for ¹⁶O in prep.

Towards 3N interactions in medium-mass nuclei based on low-momentum $V_{low k}(\Lambda) + V_{3N}(\Lambda)^{-22}$ ⁴He Hagen et al. (2007) developed coupled-cluster theory with E_{ccsD(T)} (MeV) 3N interactions, first benchmark for ⁴He -27 extrapolated: -28.23 MeV exact, FY: -28.20(5)MeV Results show that 0-, 1- and 2-body parts of 3N interaction dominate Ν 2-body only 10 2-body part +---- $1 \Delta E / E_{CCSD}^{-1}$ 0-body 3NF 1-body 3NF occupied orbits estimated triples corrections •.2-body 3NF residual 3N interaction can be neglected 10^{-3} very promising residual 3NF can include via normal-ordered RG 10 (1)(2)(3)(4)(5)



neutron $d_{3/2}$ - proton $d_{5/2}$ interaction pulls down $d_{3/2}$ neutrons in Fluorine

Why do $d_{5/2}$ neutrons not pull down $d_{3/2}$ in oxygen?



Monopole interaction and drip lines

Monopole part of nuclear forces $\mathcal{V}_{st}^T = \frac{\sum_J \mathcal{V}_{stst}^{JT} (2J+1)[1-(-)^{J+T} \delta_{st}]}{\sum_J (2J+1)[1-(-)^{J+T} \delta_{st}]}$

determines interaction of s with t orbit \rightarrow change in d_{3/2} by N_{d5/2} ν_m \Rightarrow small changes in monopoles enhanced by number of neutrons



microscopic results based only on NN interactions require phenomenological repulsive contribution to T=1 monopoles

 \rightarrow neutron d_{3/2} remains high, dripline at N=16 for Oxygen

first results indicate that $v_{m,pheno}$ due to 3N interactions

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Summary

Exciting era with advances on many fronts

For the first time, approaches from light to heavy nuclei and for astrophysics based on the same interactions

Three-nucleon interactions are a frontier

Major investments in new facilities worldwide

Exciting intersections with problems in many related areas