

Varying fundamental constants

-

a key property and key problem  
of quintessence

**Fundamental  
“constants” are not  
constant**

*Have coupling constants in the  
early Universe  
other values than today ?*

**Yes !**

# Fundamental couplings in quantum field theory

*Masses and coupling constants  
are determined by properties  
of **vacuum** !*

Similar to Maxwell – equations in matter

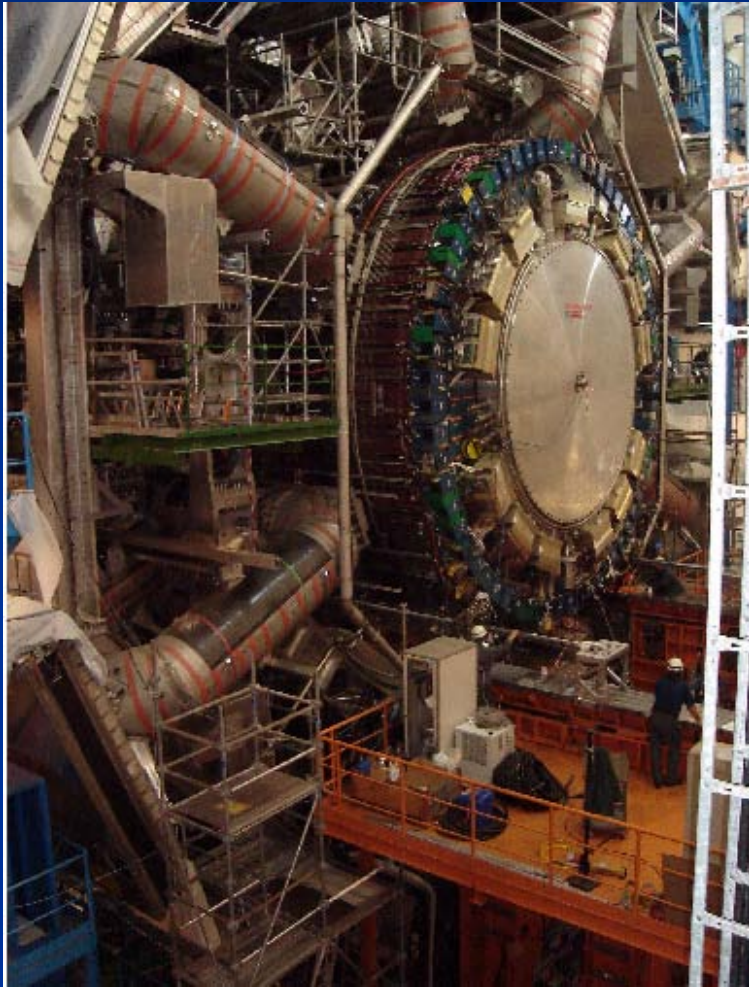
# Condensed matter physics : laws depend on state of the system

- Ground state , thermal equilibrium state ...
- Example : Laws of electromagnetism in superconductor are different from Maxwells' laws

# Standard model of particle physics :

Electroweak gauge symmetry is spontaneously broken by expectation value of Higgs scalar

# Spontaneous symmetry breaking confirmed at the LHC



# Cosmology :

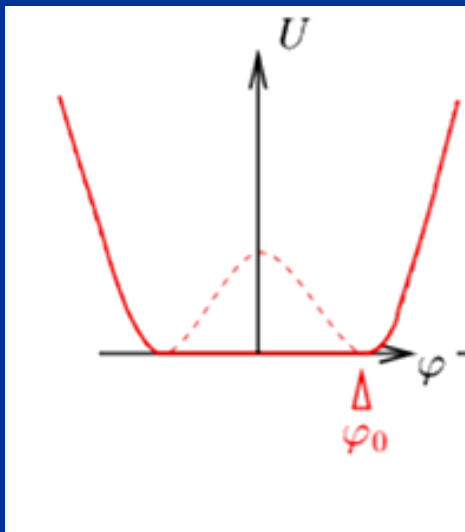
- Universe is not in one fixed state
- Dynamical evolution
- Laws are expected to depend on time



# Restoration of symmetry at high temperature in the early Universe

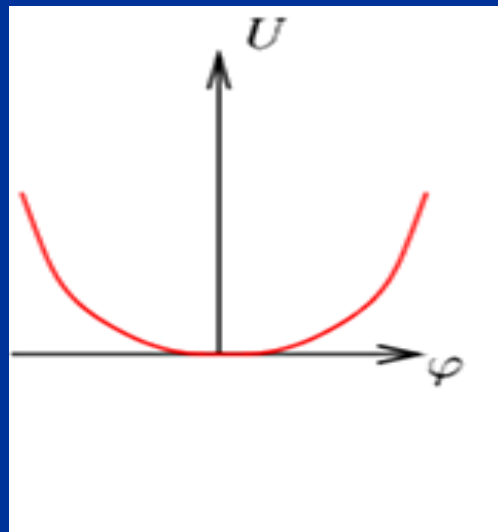
Low T  
SSB

$$\langle \varphi \rangle = \varphi_0 \neq 0$$



High T  
SYM

$$\langle \varphi \rangle = 0$$



high T :  
Less order  
More symmetry

Example:  
Magnets

In hot plasma  
of early Universe :

masses of electron und muon  
not different!

similar strength of electromagnetic  
and weak interaction

# Varying couplings

only question :

How strong is **present** variation of couplings ?

# Can variation of fundamental “constants” be observed ?

Fine structure constant  $\alpha$  (electric charge)

Ratio electron mass to proton mass

Ratio nucleon mass to Planck mass

# Time evolution of couplings and scalar fields

- Fine structure constant depends on value of Higgs field :  $\alpha(\varphi)$
- Time evolution of  $\varphi$ 
  - ➔ Time evolution of  $\alpha$

Jordan,...

# Static scalar fields

In Standard Model of particle physics :

- Higgs scalar has settled to its present value around  $10^{-12}$  seconds after big bang.
- Chiral condensate of QCD has settled at present value after quark-hadron phase transition around  $10^{-6}$  seconds after big bang .
- No scalar with mass below pion mass.
- No substantial change of couplings after QCD phase transition.
- Coupling constants are frozen.

**Observation of time- or space-  
variation of couplings**



**Physics beyond Standard Model**

# Particle masses in quintessence cosmology

can depend on value of cosmon field

similar to dependence on value of Higgs field



# Time evolution of couplings and scalar fields

Fine structure constant depends on value of cosmon field :  $\alpha(\varphi)$

Cosmon field changes in present cosmological epoch

Time evolution of  $\varphi$    
Time evolution of  $\alpha$

Jordan,...

A deep space photograph showing a vast field of galaxies and stars against a black background. The galaxies are of various shapes, including spirals and ellipticals, and are scattered across the frame. Some are bright and clear, while others are faint and distant. The stars appear as small, bright points of light, some with visible diffraction patterns.

**Dark Energy :**  
**Energy density that does not clump**

Photons , gravitons : insignificant



# Matter

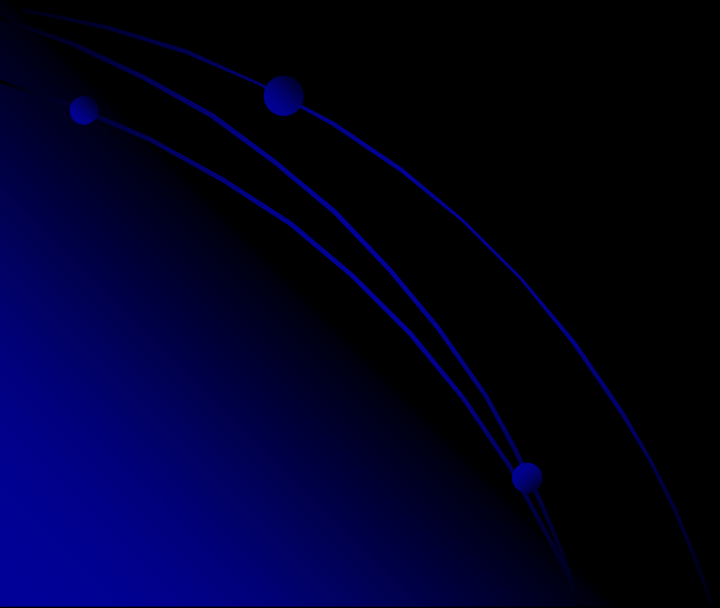
Gravitational Lens in Galaxy Cluster Abell 1689

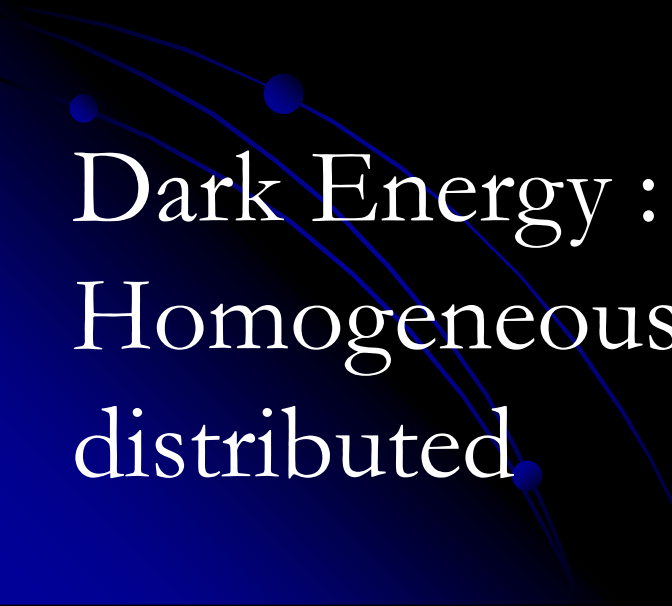


HUBBLESITE.org

Space between clumps  
is not empty :

Dark Energy !

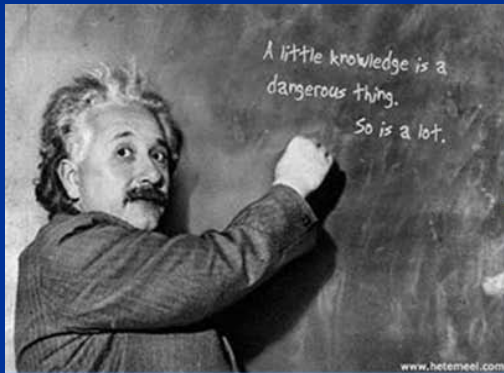




Dark Energy :  
Homogeneously  
distributed.



Einstein's equations :  
almost static Dark Energy predicts  
accelerated expansion of Universe



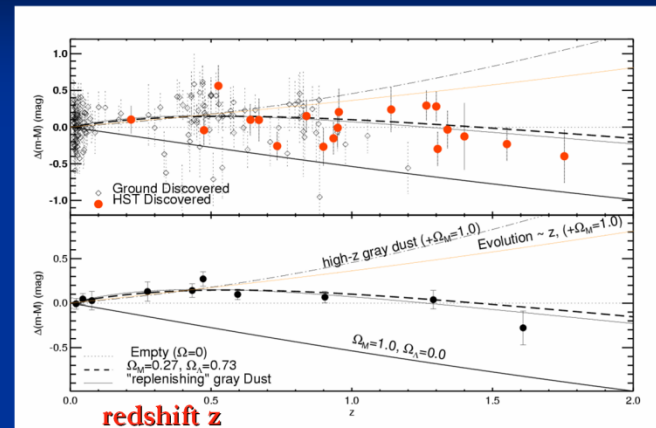
$$G_{\mu\nu} = 8\pi T_{\mu\nu}$$



# Predictions for dark energy cosmologies

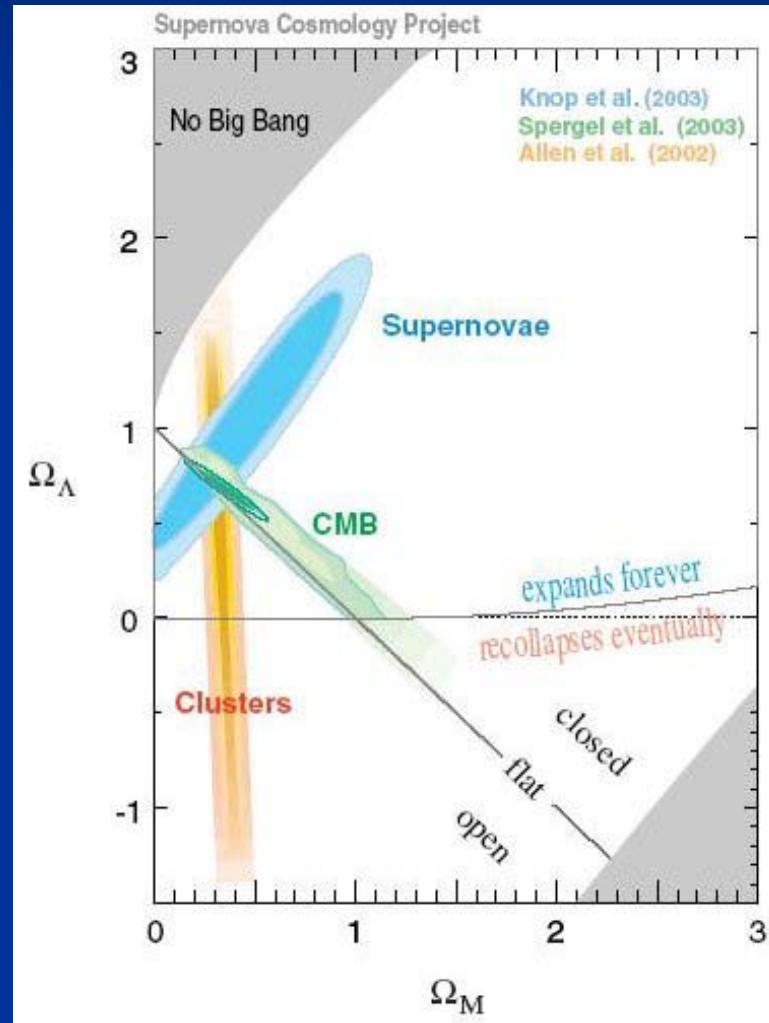
*The expansion of the Universe  
accelerates today !*

## Supernovae 1a Hubble diagram



# Dark Energy :

## observations fit together !





# Composition of the Universe

$$\Omega_b = 0.045$$

visible

clumping

$$\Omega_{dm} = 0.25$$

invisible

clumping

$$\Omega_h = 0.7$$

invisible

homogeneous

# What is Dark Energy ?

Cosmological Constant  
or  
Quintessence ?

# Cosmological Constant

## - Einstein -

- Constant  $\lambda$  compatible with all symmetries
- No time variation in contribution to energy density
- Why so small ?  $\lambda/M^4 = 10^{-120}$
- Why important just today ?

# Cosmological mass scales

- Energy density

$$\rho \sim (2.4 \times 10^{-3} \text{ eV})^{-4}$$

- Reduced Planck mass

$$M = 2.44 \times 10^{27} \text{ eV}$$

- Newton's constant

$$G_N = (8\pi M^2)$$

Only ratios of mass scales are observable !

homogeneous dark energy:  $\rho_h/M^4 = 6.5 \cdot 10^{-121}$

matter:  $\rho_m/M^4 = 3.5 \cdot 10^{-121}$

# Time evolution

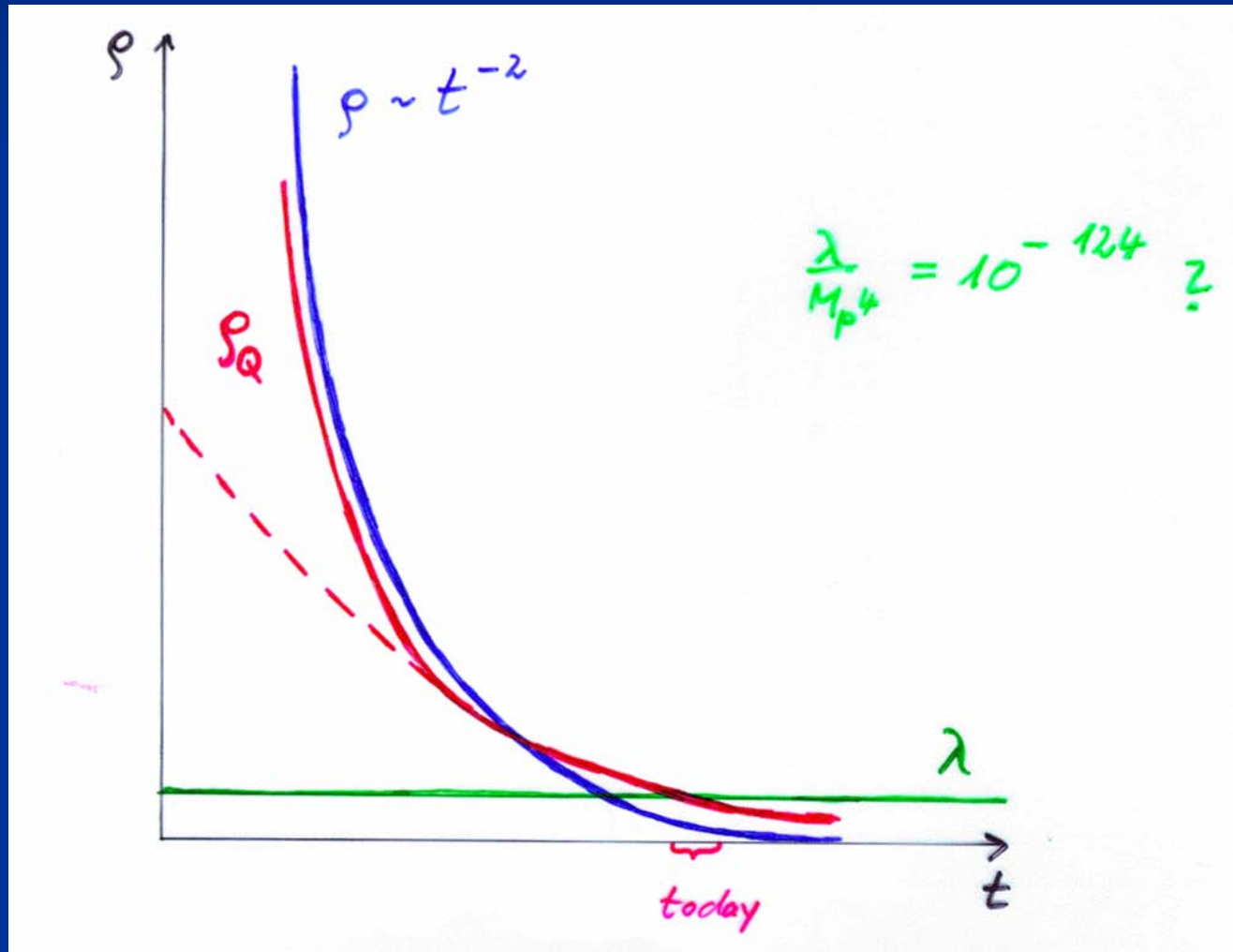
- $\rho_m/M^4 \sim a^{-3} \sim \begin{matrix} t^{-2} & \text{matter dominated universe} \\ t^{-3/2} & \text{radiation dominated universe} \end{matrix}$
- $\rho_r/M^4 \sim a^{-4} \sim t^{-2}$  radiation dominated universe

Huge age  $\Rightarrow$  small ratio

Same explanation for small dark energy?

Cosm. Const.  
static

Quintessence  
dynamical



# Quintessence

Dynamical dark energy ,  
generated by scalar field  
(cosmon)

C.Wetterich,Nucl.Phys.B302(1988)668, 24.9.87  
P.J.E.Peebles,B.Ratra,ApJ.Lett.325(1988)L17, 20.10.87

**Prediction :**

**homogeneous dark energy  
influences recent cosmology**

**- of same order as dark matter -**

Original models do not fit the present observations  
.... modifications



# Quintessence

Cosmon – Field  $\varphi(x,y,z,t)$

similar to electric field , but no direction ( scalar field )

Homogeneous und isotropic Universe :  $\varphi(x,y,z,t)=\varphi(t)$

Potential und kinetic energy of the cosmon -field  
contribute to a dynamical energy density of the Universe !

# Cosmon

- *Scalar field changes its value even in the **present** cosmological epoch*
- *Potential und kinetic energy of cosmon contribute to the energy density of the Universe*
- *Time - variable dark energy :  
 $\varrho_b(t)$  decreases with time !*

# Evolution of cosmon field

Field equations

$$\ddot{\phi} + 3H\dot{\phi} = -dV/d\phi$$

$$3M^2H^2 = V + \frac{1}{2}\dot{\phi}^2 + \rho$$

Potential  $V(\varphi)$  determines details of the model

$$V(\varphi) = M^4 \exp(-\alpha\varphi/M)$$

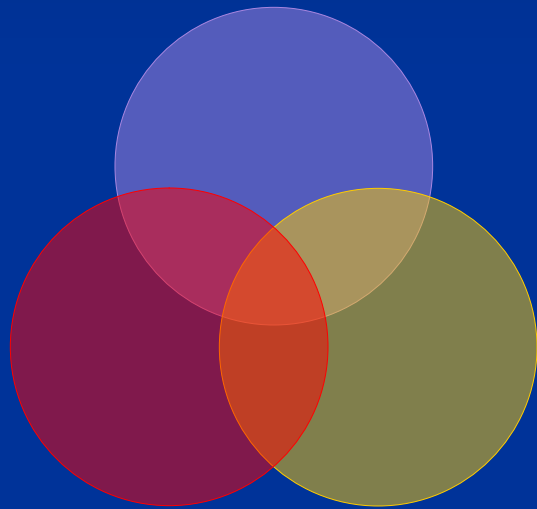
for increasing  $\varphi$  the potential decreases  
towards zero !

# Cosmon

- *Tiny mass*
- $m_c \sim H$  (depends on time !)
- *New long - range interaction*

# “Fundamental” Interactions

Strong, electromagnetic, weak  
interactions



gravitation

cosmodynamics

On astronomical  
length scales:

graviton

+

cosmon

# Time varying constants

- It is not difficult to obtain quintessence potentials from higher dimensional or string theories
- Exponential form rather generic  
( after Weyl scaling)
- But most models show too strong time dependence of fundamental constants !

# Bounds on time varying couplings from nucleosynthesis

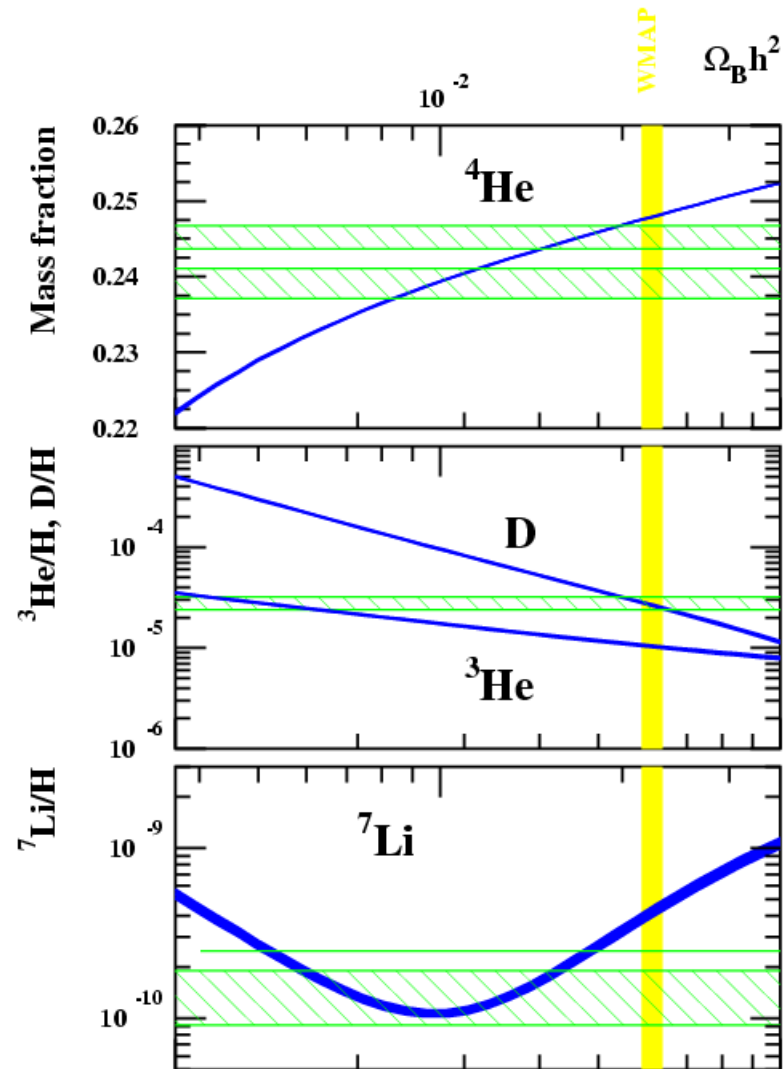
**baryons :**

**the matter of stars and humans**

$$\Omega_b = 0.045$$



# Abundancies of primordial light elements from nucleosynthesis

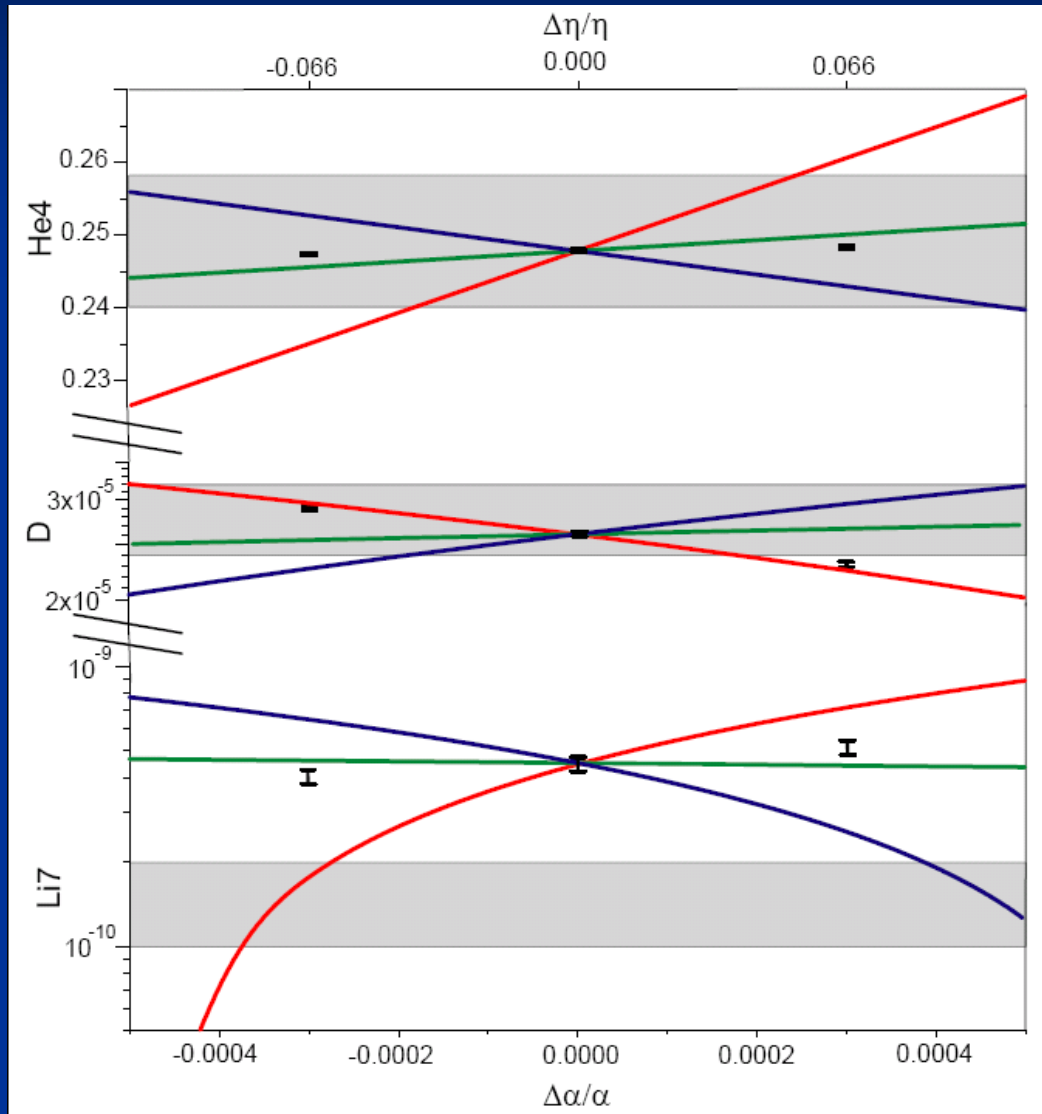


# primordial abundances for three GUT models

He

D

Li



present  
observations :  
 $1\sigma$

T.Dent,  
S.Stern,...

# three GUT models

- unification scale  $\sim$  Planck scale
- 1) All particle physics scales  $\sim \Lambda_{\text{QCD}}$
- 2) Fermi scale and fermion masses  $\sim$  unification scale
- 3) Fermi scale varies more rapidly than  $\Lambda_{\text{QCD}}$

$\Delta\alpha/\alpha \approx 4 \cdot 10^{-4}$  allowed for GUT 2 and 3 , larger for GUT 1

$\Delta\ln(M_n/M_p) \approx 40$   $\Delta\alpha/\alpha \approx 0.015$  allowed

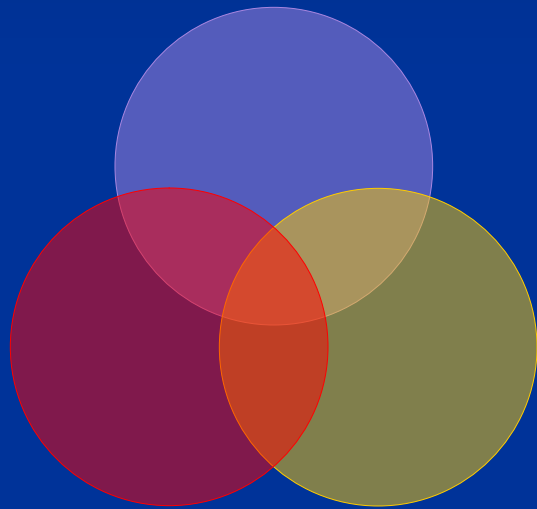
Time variation of coupling constants  
must be tiny –

would be of very high significance !

**Possible signal for Quintessence**

# “Fundamental” Interactions

Strong, electromagnetic, weak  
interactions



gravitation

cosmodynamics

On astronomical  
length scales:

graviton

+

cosmon

# “Fifth Force”

- Mediated by scalar field

R.Peccei,J.Sola,C.Wetterich,Phys.Lett.B195,183(1987)

- Coupling strength: weaker than gravity  
( nonrenormalizable interactions  $\sim M^{-2}$  )
- Composition dependence  
    → violation of equivalence principle
- Quintessence: connected to time variation of  
fundamental couplings

C.Wetterich , Nucl.Phys.B302,645(1988)

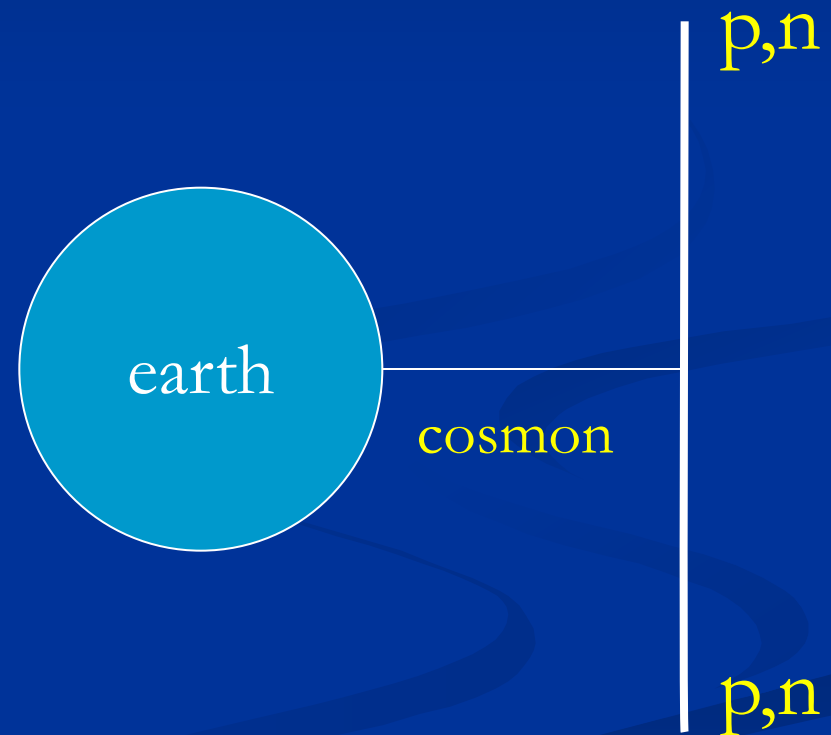
# Violation of equivalence principle

Different couplings of  
cosmon to proton and  
neutron

Differential acceleration

“Violation of  
equivalence principle”

only apparent : new “fifth force” !



# Differential acceleration

Two bodies with equal mass experience  
a different acceleration !

$$\eta = (a_1 - a_2) / (a_1 + a_2)$$

$$\text{bound : } \eta < 3 \cdot 10^{-14}$$



*Apparent violation of equivalence principle*

*and*

*time variation of fundamental couplings*

*measure both the*

*cosmon — coupling to ordinary matter*

# Differential acceleration $\eta$

For unified theories ( GUT ) :

$$\eta = -1.75 \cdot 10^{-2} \Delta R_z \left( \frac{\partial \ln \alpha}{\partial z} \right)^2 \frac{1 + \tilde{Q}}{\Omega_h (1 + w_h)}$$

$$\Delta R_z = \frac{\Delta Z}{Z + N} \approx 0.1$$

$$\eta = \Delta a / 2a$$

Q : time dependence of other parameters

*Link between time variation of  $\alpha$   
and violation of equivalence principle*

typically :  $\eta = 10^{-14}$

if time variation of  $\alpha$

near Oklo upper bound

to be tested ( MICROSCOPE , ...)



# Cosmon coupling to atoms

- Tiny !!!
- Substantially weaker than gravity.
- Non-universal couplings bounded by tests of equivalence principle.
- Universal coupling bounded by tests of Brans-Dicke parameter  $\omega$  in solar system.
- Only very small influence on cosmology.

( All this assumes validity of linear approximation )

# key problem for realistic quintessence

**tiny cosmon - atom coupling**

Planck mass may increase, but particle masses must be (almost) proportional to it

C.Wetterich, Nucl.Phys.B302, 645 and 668 (1988)

# Possible mechanism for tiny cosmon-atom couplings

- asymptotic approach to fixed point for dimensionless couplings and mass ratios
- at fixed point : no cosmon coupling to atoms – no time variation of fundamental constants
- very near fixed point : tiny coupling
- how small ?

# Approach to fixed point for $\varphi \rightarrow \infty$

$$V = M^4 \exp\left(-\alpha \frac{\varphi}{M}\right)$$

$$\frac{m_n}{M} = f_0 + c_n \exp\left(-\beta_n \frac{\varphi}{M}\right)$$

requires for small time  
variation of couplings :

$$f_0 \neq 0$$

or

$$\beta_n \ll \alpha$$

similar for fine structure constant

$$\mathcal{L}_F = \frac{1}{4} \left[ \frac{1}{e_0^2} + c_\alpha \exp\left(-\beta_\alpha \frac{\varphi}{M}\right) \right] F_{\mu\nu} F^{\mu\nu}$$

$$\frac{1}{e^2} = \frac{1}{e_0^2} + c_\alpha \exp\left(-\beta_\alpha \frac{\varphi}{M}\right) = \frac{1}{4\pi\alpha}$$

# Neutrino cosmon coupling

- Strong bounds on atom-cosmon coupling from tests of equivalence principle or time variation of couplings.
- No such bounds for neutrino-cosmon coupling.
- In particle physics : Mass generation mechanism for neutrinos differs from charged fermions. Seesaw mechanism involves heavy particles whose mass may depend on the value of the cosmon field.



# neutrino mass

$$M_\nu = M_D M_R^{-1} M_D^T + M_L$$

$$M_L = h_L \gamma \frac{d^2}{M_t^2}$$

seesaw and  
cascade  
mechanism

triplet expectation value  $\sim$  doublet squared

$$m_\nu = \frac{h_\nu^2 d^2}{m_R} + \frac{h_L \gamma d^2}{M_t^2}$$

omit generation  
structure

# neutrino mass

$$M_\nu = M_D M_R^{-1} M_D^T + M_L$$

(?) ....

C.Wetterich, Nucl.Phys.B187 (1981) 343

$$M_L = h_L \gamma \frac{d^2}{M_t^2}$$

cascade ( seesaw II )  
mechanism

M.Magg, C.W. 1980

# Neutrino cosmon coupling

- realized by dependence of neutrino mass on value of cosmon field

$$\beta(\varphi) = -M \frac{\partial}{\partial \varphi} \ln m_\nu(\varphi)$$

- $\beta \approx 1$  : cosmon mediated attractive force between neutrinos has similar strength as gravity

growing neutrino  
quintessence

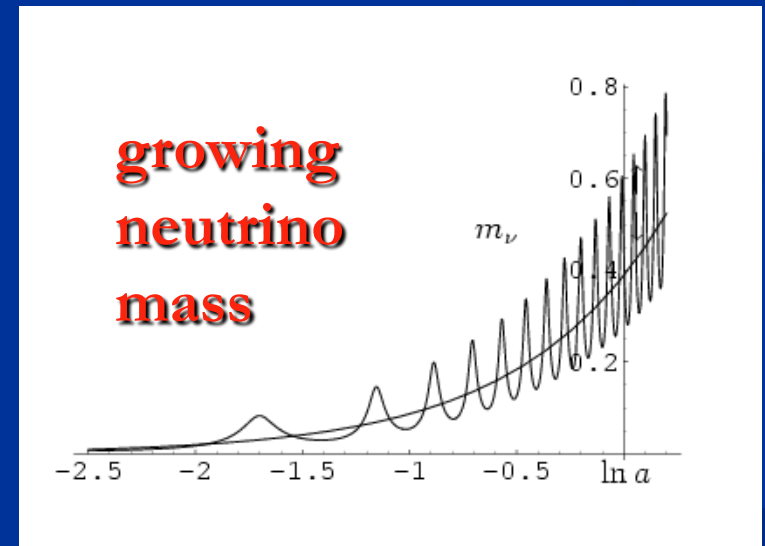
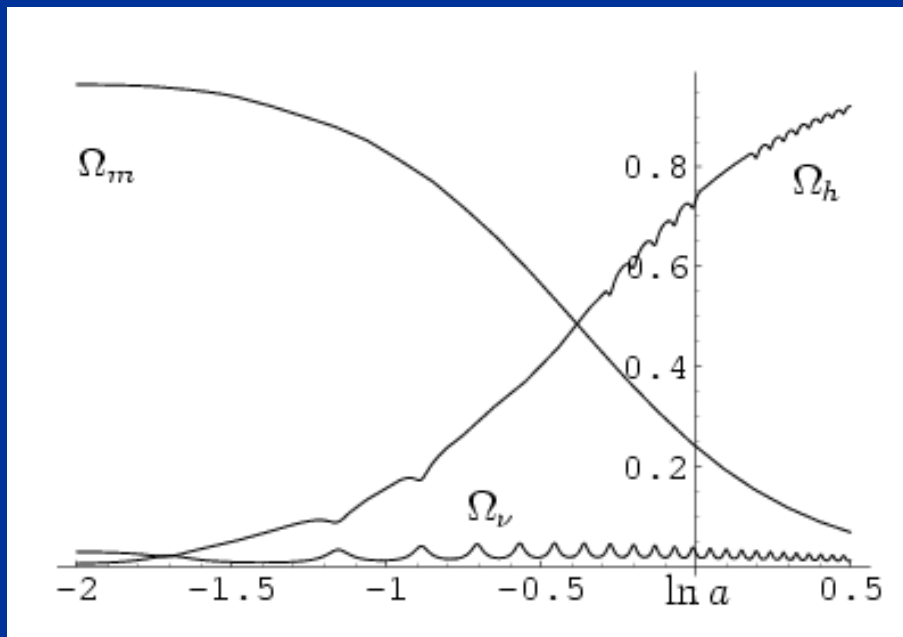
# growing neutrinos change cosmological evolution

$$\ddot{\varphi} + 3H\dot{\varphi} = -\frac{\partial V}{\partial \varphi} + \frac{\beta(\varphi)}{M}(\rho_\nu - 3p_\nu),$$
$$\beta(\varphi) = -M \frac{\partial}{\partial \varphi} \ln m_\nu(\varphi) = \frac{M}{\varphi - \varphi_t}$$

modification of conservation equation for neutrinos

$$\begin{aligned}\dot{\rho}_\nu + 3H(\rho_\nu + p_\nu) &= -\frac{\beta(\varphi)}{M}(\rho_\nu - 3p_\nu)\dot{\varphi} \\ &= -\frac{\dot{\varphi}}{\varphi - \varphi_t}(\rho_\nu - 3p_\nu)\end{aligned}$$

# growing neutrino mass triggers transition to almost static dark energy



L. Amendola, M. Baldi, ...

effective cosmological trigger  
for stop of cosmon evolution :  
neutrinos get non-relativistic

- this has happened recently !
- sets scales for dark energy !

# connection between dark energy and neutrino properties

$$[\rho_h(t_0)]^{\frac{1}{4}} = 1.27 \left( \frac{\gamma m_\nu(t_0)}{eV} \right)^{\frac{1}{4}} 10^{-3} eV$$

present dark energy density given by neutrino mass

present equation  
of state given by  
neutrino mass !

$$w_0 \approx -1 + \frac{m_\nu(t_0)}{12eV}$$



# Induced small cosmon-atom coupling

- change of triplet expectation value induces tiny change of doublet expectation value
- change of Fermi scale
- change of ratio electron to proton mass
- change of fine structure constant

Time variation of coupling constants  
is tiny , but would be of very  
high significance !

Size not predicted

**Possible signal for Quintessence**

Three decorative, wavy, light blue lines that sweep across the bottom right portion of the slide, starting from the right edge and moving towards the center.

End