

6.4 KKLT

(Kachru, Kallosh, Linde, Trivedi '03)

- Focus on $h^{1,1} = 1$, i.e.

$K = -3 \ln(T + \bar{T})$; $W = W_0 \leftarrow$ const. det. by flux choice

$V = 0$; SUSY ($m_{3/2} \sim e^{K/2} W_0$)

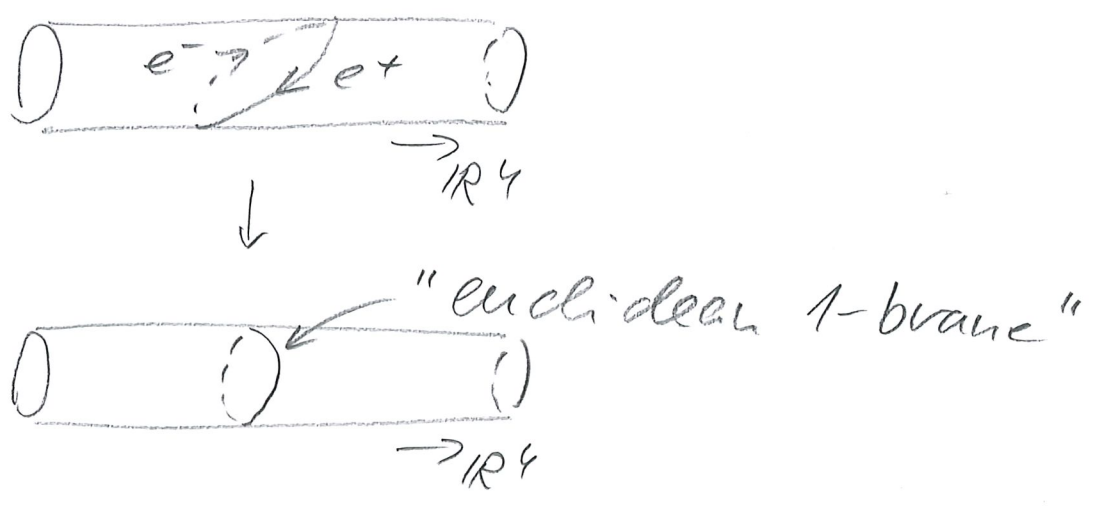
- Quantum corrections will break this "no-scale structure" ($V = 0$)

E.g. Stringy / Exotic / D-brane-Instantons

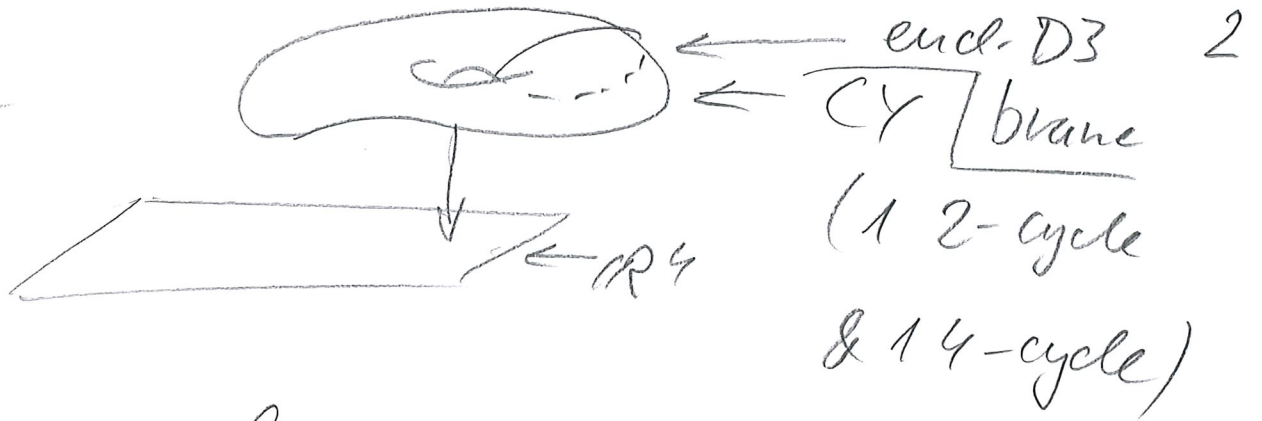
$$W_0 \longrightarrow W_0 + A e^{-2\pi T}$$

\uparrow \uparrow
 dep. on z

- Idea same as for our toy model
instantons in ED in 5d \rightarrow 4d on S^1



• Here



action: $\int_{C_4} \sqrt{g_6} \sim R_{CY}^4 \sim \text{Re} T \sim \tau$

↑
 4-cycle-volume

couples to C_4 in 10d;

$\int_{C_4} C_4 \sim c$ of $T = \tau + ic$

\Rightarrow single instanton effect: $\sim e^{-2\pi\tau} e^{-ic}$

action
 suppression
 of tunneling
 (e^{-S})
 ↑
 coupling
 to axion

Sum over all
 numbers of inst.

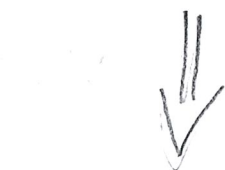
\Rightarrow exponentiation

Non-Susy $e^{-2\pi\tau} \cos(c)$

Susy: $W_0 \rightarrow W_0 + Ae^{-T}$
 (or $k \rightarrow k + e^{-T} + c.c.$)

Now: $K = -3 \ln(T + \bar{T})$

$$W = W_0 + e^{-T} \quad (A = a = 0 \text{ for simplicity})$$



$$V = V(\tau) \quad (\text{assume } V(\tau, c) \text{ is minimized in } c; \text{ ignore } c)$$



← Ads (neg. -2) minimum

(Stability restored; can check:

$DW = 0$ at this minimum)

↑ technically

$$DW = -e^{-T} - \frac{3}{T + \bar{T}} (W_0 + e^{-T}) \stackrel{!}{=} 0$$

\Updownarrow (assume $c = 0$)

$$W_0 = -\left(1 + \frac{2}{3}T\right) e^{-T}$$

Need $W_0 < 1$ (& real, neg.) to be in parametric control ($Re \lambda > 1$).

- Note: This works for any phase of W_0 , just $c=0$ will be modified.

\Rightarrow so far: landscape of \uparrow AdS vacua.
(susy)

- Want: dS; ~~susy~~

- "Macroscopic" description

$$K \rightarrow K(\tau, \bar{\tau}) + \delta K(x, \bar{x})$$

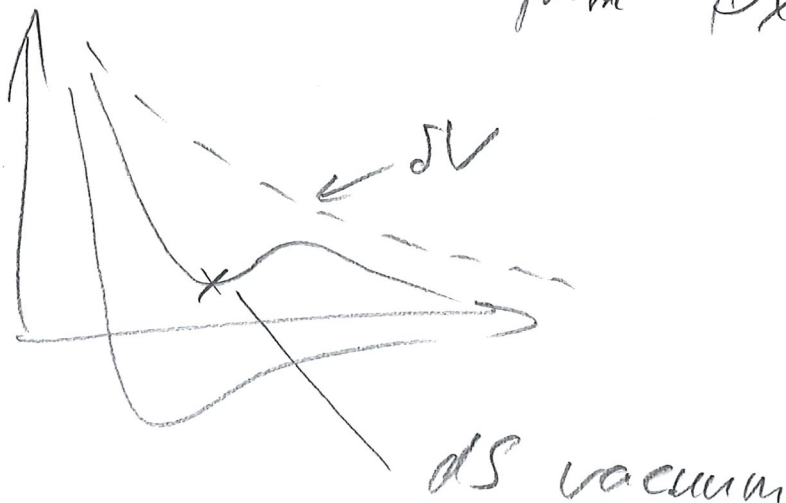
$$W \rightarrow W(\tau) + \delta W(x)$$



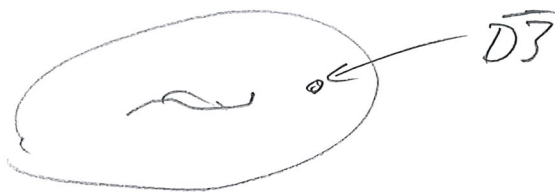
"O'Riada type - type
~~susy~~ sector"

$$\Rightarrow V \rightarrow V + \delta V$$

\uparrow
from $D_x W / 2 \neq 0$



- Presumably many such optima in (huge) landscape
- Explic. realization not trivial (exist even opinion that impossible for fundamental reasons)
- Hist. first & simplest idea (KLELT):
Add $\overline{D3}$ -brane (breaks SUSY relative "anti" to the $O3/D3$ planes/branes which are already present)
- For us: Add point-like "pos. tension" object in CY ↑
"pos. 2"



$\Rightarrow \mathcal{N} = \text{const.}$ in "Brans-Dicke-frame"

$$(S \sim V_6 R_4 + \dots)$$

go to "Einstein frame" $(S \sim R_4 + \dots)$

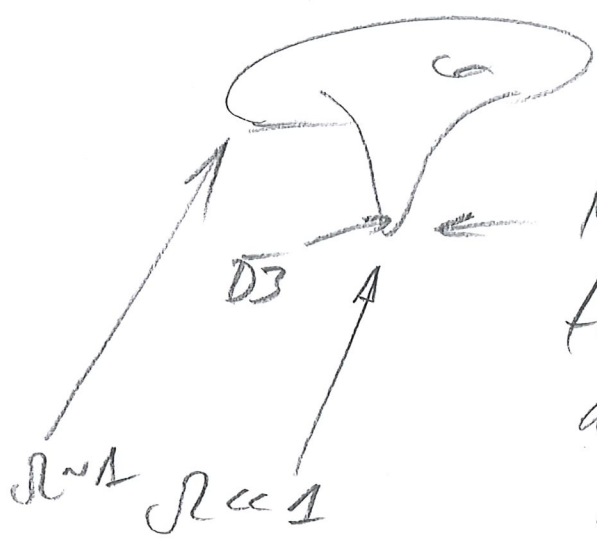
$$\Rightarrow \mathcal{N} \sim \frac{1}{V_6^2} \sim \frac{1}{L^3} \quad \text{by } g_{\mu\nu} \rightarrow g_{\mu\nu}/V_6$$



no dS minimum since uplift too strong.

Idea: (Based on a feature already present in GGP)

Place D3 in "strongly warped region"



Metric is such that in this region all energies are scaled down exponentially

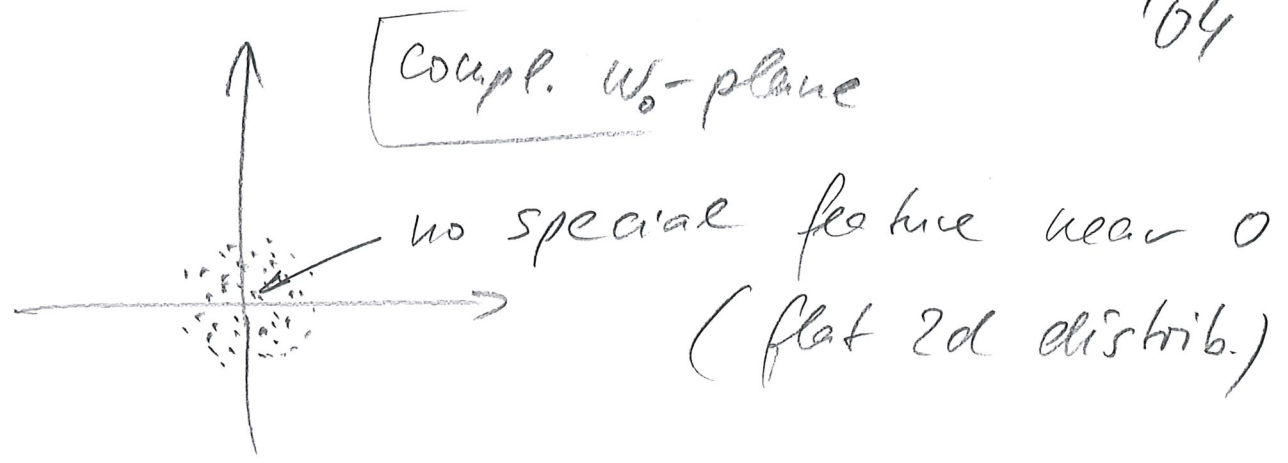
("Klebanov-Strassler-Throat")

Indeed: general compactification ansatz:

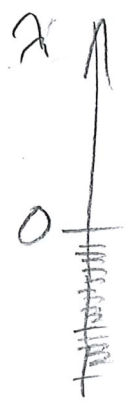
Warp factor $\rightarrow ds^2 = R^2(\gamma) \eta_{\mu\nu} dx^\mu dx^\nu + g_{ij} dy^i dy^j$
Max. symm.

Assuming this works and using W_0 -distribution found by Derif/Douglas

'04



⇒ MLCT, step 1



uplift
→

MLCT, step 2



very small λ will be present with statistical certainty.

Overall number of vacua:

due to $\int C_4 \wedge F_3 \wedge H_3 \rightarrow$ fluxes source

C_4 - need opp. charge from O3-planes

or similar effects, limit at

$$(h, f)^2 \leq L_* \approx 10^4$$

↑
max "tadpole"

(feature of CY orientifold)

$$W_{\text{vac.}} \sim \frac{L_*^k}{k!} \sim \left(\frac{L_*}{k}\right)^k$$

↑
of 3-cycles

Say, $k=300$, $L_* = 3000 \Rightarrow N_{\text{vac.}} \sim 10^{300}$

(Wells, much bigger!)

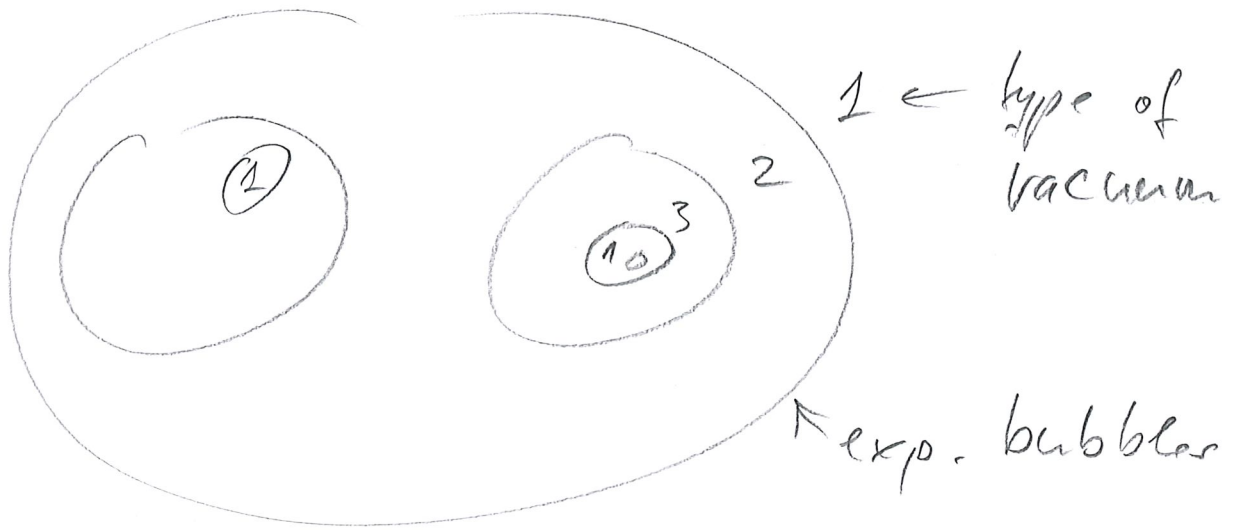
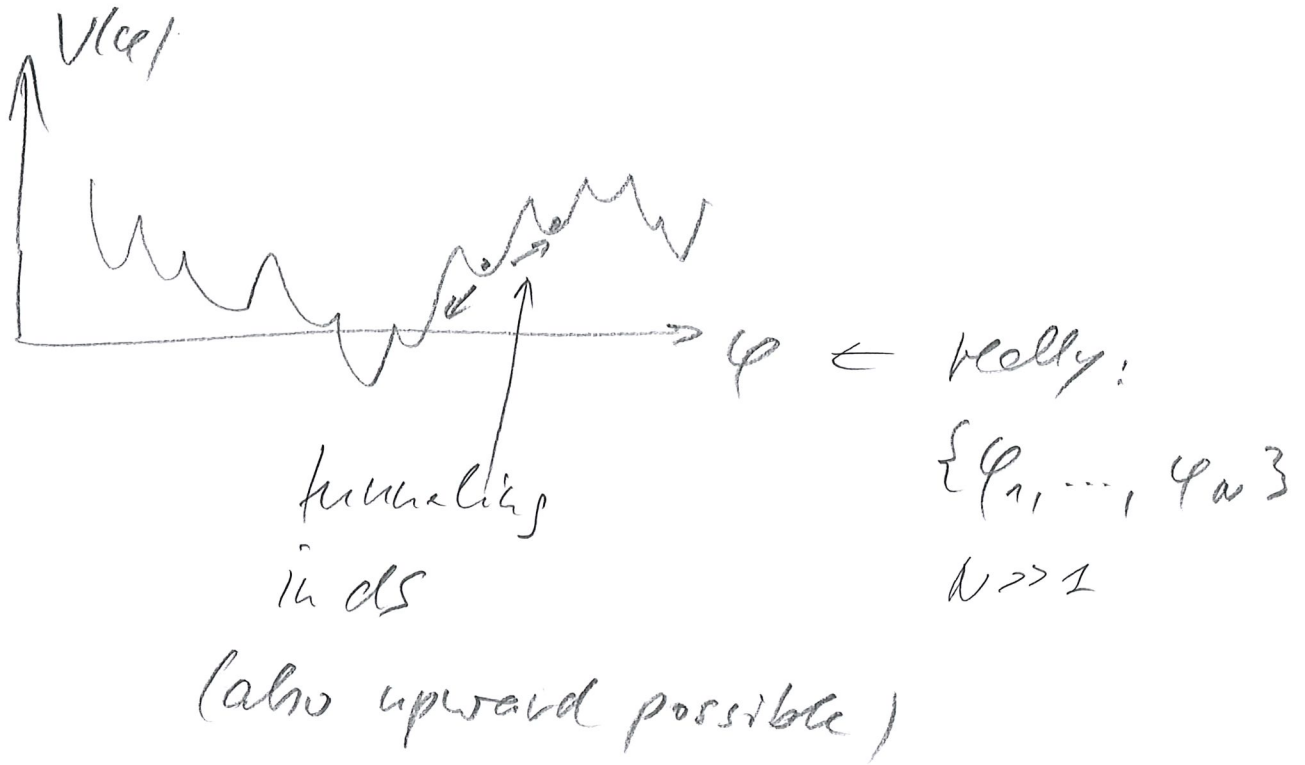
Rough other estimate:

300 cycles; 2 flux from $\underbrace{-10 \dots 10}_{20}$

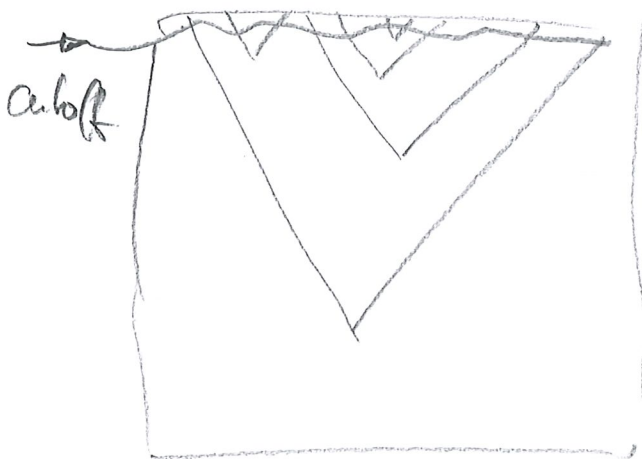
40

$\Rightarrow 40^{300} \sim 10^{500}$

7 Eternal Infl. & Measure problem



In Penrose diagram



stat. predict.
 \Rightarrow count observers
 (galaxies, planets?)
 \Rightarrow Cut-off / Prescript. dep.
 is open problem
 ("measure problem")