

5 Perturbation Theory - Leading Order Approach

5.1 The S-Matrix

- We have justified $V(\varphi) = \frac{m^2}{2} \varphi^2$ by Taylor exp.
- It is convenient to impose the discrete symm. $\varphi \rightarrow -\varphi$, such that $\lambda/4! \cdot \varphi^4$ is the next term

$$\Rightarrow \begin{aligned} \mathcal{L}_0 &\rightarrow \mathcal{L}_0 + \mathcal{L}_{int} & \mathcal{L}_{int} &= -\frac{\lambda}{4!} \varphi^4 \\ \mathcal{H}_0 &\rightarrow \mathcal{H}_0 + \mathcal{H}_{int} & \mathcal{H}_{int} &= +\frac{\lambda}{4!} \varphi^4 \end{aligned}$$

$$("L = T - V, H = T + V")$$

Comment:

- We could also drop the symm. requirement $\varphi \rightarrow -\varphi$ and consider $V_{int} = \frac{\lambda}{3!} \varphi^3$. This is sometimes dismissed because of instability at $\varphi \rightarrow -\infty$, but pert. theory around the local min. at $\varphi = 0$ is still ok. The real reason for starting with φ^4 is that it is slightly easier to deal with in our context.
- We will use the "Interaction Picture", i.e. keep the free time evolution of operators of the Heisenberg picture and let the additional time evolution act on states:

Schrödinger

$\mathcal{O}$

$$|\psi_t\rangle = e^{-iHt} |\psi\rangle$$

—————→

Interaction ("I")

$$\mathcal{O}_t^I = e^{iH_0 t} \mathcal{O} e^{-iH_0 t}$$

$$|\psi_t^I\rangle = e^{iH_0 t} e^{-iHt} |\psi\rangle$$

- Evolution from $t=0$ to t' is then analogously given by

$$|\psi_{t'}^I\rangle = e^{iH_0 t'} e^{-iH t'} |\psi\rangle.$$

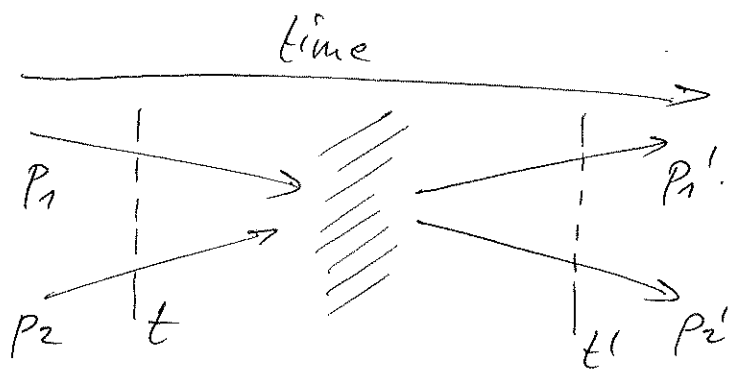
• Thus:

$$|\psi_{t'}^I\rangle = U(t', t) |\psi_t^I\rangle$$

where $U(t', t) = e^{iH_0 t'} e^{-iH(t'-t)} e^{-iH_0 t}$

(The operator U , which you should know from time-dep. pert. th. in QM, describes the unitary evolution of states in the interaction picture.)

- Since $\int d^3x \varphi^4 \supset a_{\vec{p}_1}^\dagger a_{\vec{p}_2}^\dagger a_{\vec{p}_1} a_{\vec{p}_2}$, we expect that at LO our interaction will induce 2-2-scattering (two-to-two...)



Crucial: We assume that, at t and t' , the two particles are separated in space (wave packets rather than plane waves)

(Note: At higher orders in pert. th. or with more complicated interaction terms we could also have 2- n -scattering ($n > 2$).)

- To work this out, let us first split the time evolution from t to t' into n small steps: $\Delta = \frac{t' - t}{n}$;

$$U(t', t) = U(t', t' - \Delta) U(t' - \Delta, t' - 2\Delta) \dots U(t + \Delta, t)$$

• Furthermore,

$$\begin{aligned} U(t + \Delta, t) &= e^{iH_0(t+\Delta)} e^{-iH_0\Delta} e^{-iH_0t} \\ &= e^{iH_0t} e^{iH_0\Delta} e^{-iH_0\Delta} e^{-iH_0t} = e^{iH_0t} e^{-iH_{\text{int}}\Delta} e^{-iH_0t} + O(\Delta^2) \\ &= e^{-iH_{\text{int}}^{\text{I}}(t) \cdot \Delta} + O(\Delta^2), \quad \text{where } H_{\text{int}}^{\text{I}}(t) = e^{iH_0t} H_{\text{int}} e^{-iH_0t} \end{aligned}$$

is the interaction Hamiltonian

H_{int} transformed to the interaction picture.

• Thus,

$$U(t', t) = e^{-iH_{\text{int}}^{\text{I}}(t' - \Delta) \cdot \Delta} e^{-iH_{\text{int}}^{\text{I}}(t' - 2\Delta) \cdot \Delta} \dots e^{-iH_{\text{int}}^{\text{I}}(t) \cdot \Delta}$$

$$\equiv T \exp \left[-i \int_t^{t'} d\tau H_{\text{int}}^{\text{I}}(\tau) \right]$$

↑
time-ordering "operator" (better: "symbol"),
defined by

$$T\varphi(t_1)\varphi(t_2) = \begin{cases} \varphi(t_1)\varphi(t_2) & \text{if } t_1 \geq t_2 \\ \varphi(t_2)\varphi(t_1) & \text{if } t_1 < t_2 \end{cases}$$

• Explicitly, we have

$$H_{\text{int}}^{\text{I}}(t) = e^{iH_0t} \int d^3x \frac{\lambda}{4!} (\varphi^{\text{I}}(x))^4 e^{-iH_0t} = \int d^3x \frac{\lambda}{4!} (\varphi^{\text{I}}(x))^4$$

↑
interaction-picture field
 $\equiv$ Heis.-picl. field of free theory

- We now define the S-matrix:

$$S = \lim_{\substack{t \rightarrow -\infty \\ t' \rightarrow +\infty}} U(t', t) = T e^{-i \int dt H_{\text{int}}^I(t)} = \underline{\underline{T e^{i \int d^4x \mathcal{L}_{\text{int}}^I(x)}}}$$

and the S-matrix-element:

$$\underline{\underline{S_{fi} = \langle p_1' p_2' | T \exp(i \int d^4x \mathcal{L}_{\text{int}}^I) | p_1 p_2 \rangle}}$$

(f/i stands for final/initial)

- Furthermore, we define the transition-matrix or T-matrix (do not confuse with time ordering!) by

$$S = \underline{\underline{1}} - iT \quad \& \quad S_{fi} = \delta_{fi} + iT_{fi}$$

- To work this out, it will be convenient to change a/a⁺-normalization such as to emphasize covariance rather than the harmonic-oscillator analogy:

$$(a_{\vec{p}})_{\text{new}} \equiv \sqrt{2\omega_{\vec{p}}} (a_{\vec{p}})_{\text{old}} \quad ; \quad \omega_{\vec{p}} = p^0$$

i.e.

$$\boxed{\begin{aligned} [a_{\vec{p}}, a_{\vec{q}}^+] &= 2p^0 (2\pi)^3 \delta^3(\vec{p} - \vec{q}) \quad ; \quad |p\rangle = a_{\vec{p}}^+ |0\rangle \\ \varphi^I(x) &= \int \frac{d^3p}{(2\pi)^3 2p^0} (a_{\vec{p}} e^{-ipx} + a_{\vec{p}}^+ e^{ipx}) \end{aligned}}$$

- At LO in λ we then have

$$iT_{fi} = \langle 0 | a_{\vec{p}_1'} a_{\vec{p}_2'} \left(-\frac{i\lambda}{4!}\right) \int d^4x (\varphi^I(x))^4 a_{\vec{p}_1}^+ a_{\vec{p}_2}^+ |0\rangle$$

- The crucial point of all the previous manipulations was the expression of the matrix element in free fields. Based on the last 3 lines, an easy "commutator-style" calculation gives (see problems):

$$iT_{fi} = -i\lambda (2\pi)^4 \delta^4(p_1 + p_2 - p_1' - p_2').$$

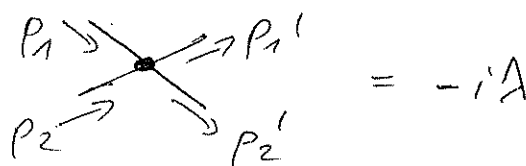
- Since the above "momentum-conservation- δ -fct." always arises in such situations, we define the "invariant matrix elements" $\mathcal{M}_{fi}$ by

$$S_{fi} = \delta_{fi} + i(2\pi)^4 \delta^4(\dots) \mathcal{M}_{fi}$$

where, for 2-2-scattering in " $\lambda\phi$ -theory" we found:

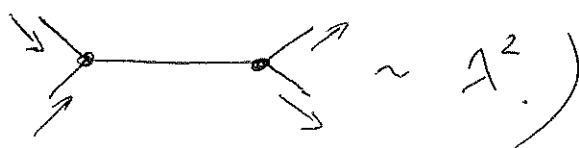
$$i\mathcal{M}_{fi} = \underline{\underline{-i\lambda}}$$

- This is our first "Feynman rule":



$$= -i\lambda$$

(We can now "guess" why $\lambda\phi^3$ -theory would be slightly more complicated:



$$\sim \lambda^2$$

- Not surprisingly (due to the limit $t \rightarrow -\infty / t' \rightarrow +\infty$), our result is singular ($\delta^4(\dots)$). The corresponding well-defined and finite quantity is the...

5.2 Scattering cross section

- Consider a "fixed target experiment" with the beam (consisting of N_B particles of type B) is spread out over a transverse area F and hits a target consisting of one particle of type A:



$$\frac{N_{\text{events}}}{N_B} = \frac{\sigma}{F} \Rightarrow \sigma = \frac{N_{\text{events}}}{(N_B/F)} \leftarrow$$

- We will need localized states and hence wave packets: transverse beam density

$$|f_{\vec{p}}\rangle \equiv \int d\vec{k} \tilde{f}_{\vec{p}}(\vec{k}) |k\rangle \quad \text{where } d\vec{k} \equiv \frac{d^3k}{(2\pi)^3 2k^0} \quad \text{and}$$

$\tilde{f}_{\vec{p}}(\vec{k})$ is peaked near $\vec{k} = \vec{p}$, e.g.

$$\tilde{f}_{\vec{p}}(\vec{k}) \sim \exp(-\alpha |\vec{k} - \vec{p}|^2)$$

- Normalization:

$$\begin{aligned} \langle f_{\vec{p}} | f_{\vec{p}} \rangle &= \int d\vec{k} \int d\vec{k}' \tilde{f}_{\vec{p}}^*(\vec{k}) \tilde{f}_{\vec{p}}(\vec{k}') \langle k' | k \rangle \\ &= \int d\vec{k} |\tilde{f}_{\vec{p}}(\vec{k})|^2 \stackrel{!}{=} 1 \end{aligned}$$

- Next, check that $|f_{\vec{p}}\rangle$ is indeed localized in $\mathbb{R}^3$ for an appropriate choice of the function $\tilde{f}_{\vec{p}}$:

Indeed, using the by now very familiar expression for $\varphi^I(x)$ in terms of a, a^+ , one easily checks that

$$a_{\vec{k}}^+ = -i \int d^3x e^{-ikx} \left. \frac{\partial}{\partial_0} \varphi^I(x) \right|_{x^0=0}$$

Thus $|f_{\vec{p}}\rangle = \int d\vec{k} f_{\vec{p}}(\vec{k}) a_{\vec{k}}^+ |0\rangle$

$$= \int d^3x \left\{ \int d\vec{k} e^{i\vec{k}\vec{x}} (k_0 \varphi^I(\vec{x}) - i \dot{\varphi}^I(\vec{x})) f_{\vec{p}}(\vec{k}) \right\} |0\rangle$$

Now, with $f_{\vec{p}}(\vec{k})$ & $k_0 f_{\vec{p}}(\vec{k})$ smooth and localized in $\vec{k}$, their Fourier transforms

$$\int d\vec{k} e^{i\vec{k}\vec{x}} f_{\vec{p}}(\vec{k}), \quad \text{---} k_0 \text{---}$$

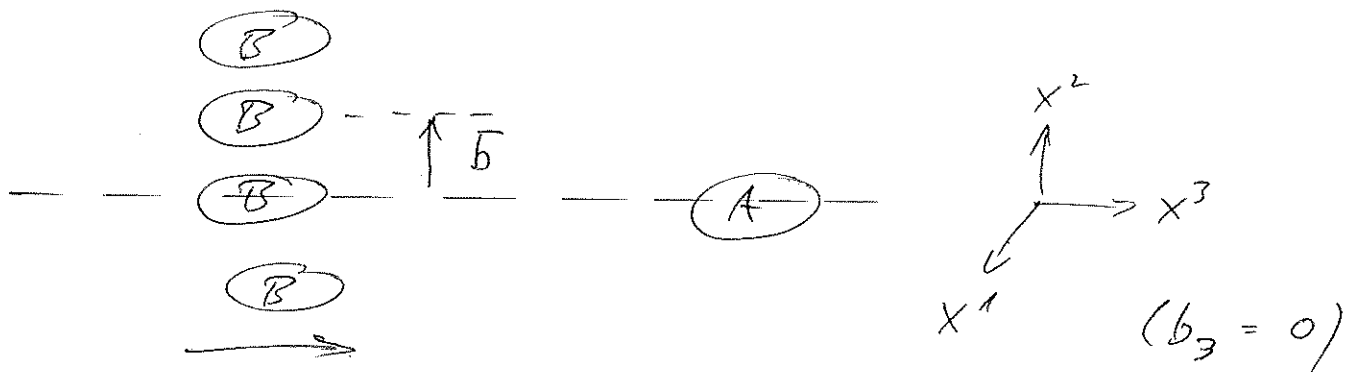
will be localized in $\vec{x}$. $|f_{\vec{p}}\rangle$ is then created by $\varphi^I(\vec{x})$ & $\dot{\varphi}^I(\vec{x})$ acting on $|0\rangle$ in a localized region near $\vec{x} = 0$.

- Thus, the state $|i\rangle = \int d\vec{k}_A d\vec{k}_B f_{\vec{p}_A}(\vec{k}_A) f_{\vec{p}_B}(\vec{k}_B) |\vec{k}_A \vec{k}_B\rangle$ describes the situation



(for appropriate $\vec{p}_A, \vec{p}_B$ and such that the A/B regions overlap at $t=0$)

- More realistically, we need to describe a situation with incoming particles spread over a certain transverse region (i.e. with impact parameters $\vec{b}$):



- As in QM, the operator $\hat{P}$ (cf. Noether Theorem) generates shifts. Thus, the incoming particles with non-zero $\bar{b}$ are given by:

$$e^{-i\hat{P}\bar{b}} \int d\tilde{k} f_{\bar{P}_B}(\tilde{k}_B) |\tilde{k}_B\rangle$$

- Since $\hat{P}|\tilde{k}_B\rangle = \tilde{k}_B |\tilde{k}_B\rangle$, an initial state with impact parameter $\bar{b}$ reads

$$|\bar{b}\rangle = \int d\tilde{k}_A d\tilde{k}_B f_{\bar{P}_A}(\tilde{k}_A) f_{\bar{P}_B}(\tilde{k}_B) e^{-i\tilde{k}_B \bar{b}} |\tilde{k}_A \tilde{k}_B\rangle$$

- We start with this state "at $t = -\infty$ " (when A & B are far apart), evolve forward in time (A & B meet "up to $\bar{b}$ " at $t = 0$), and project on the desired scattered state "at $t = +\infty$ ". The probability for this scattered state (say with momenta p_1, p_2) is

$$|\langle p_1 p_2 | S | \bar{b} \rangle|^2$$

- Summing over a set of particles with different $\bar{b}$ gives

$$N_{\text{events}} = \sum_{\bar{b}} |\langle p_1 p_2 | S | \bar{b} \rangle|^2$$

- For a homogeneous transverse distribution of N_B particles

in an area F we can approximate the sum by an integral:

$$N_{\text{events}} = \frac{N_B}{F} \int d^2b |\langle p_1 p_2 | S | i_b \rangle|^2$$

- The corresponding cross section is

$$\sigma(p_1, p_2) = \frac{N_{\text{events}}}{(N_B/F)} = \int d^2b |\langle p_1 p_2 | S | i_b \rangle|^2$$

- Clearly this is too naive since we can not ask for specific final-state momenta p_1, p_2 . (This also clashes with the finite precision of any detector.)

Instead, we define σ for a finite region of phase space $R^6 \ni (\vec{p}_1, \vec{p}_2)$:

$$\sigma(V_f) = \int_{V_f} d\vec{p}_1 d\vec{p}_2 \int d^2b |\langle p_1 p_2 | S | i_b \rangle|^2$$

- The correctness of the proposed measure $d\vec{p}_1 d\vec{p}_2$ follows immediately from considering the free theory ($S = \mathbb{1}$) and integrating over the whole phase space:

$$1 \stackrel{!}{=} \int d\vec{p}_1 d\vec{p}_2 |\langle p_1 p_2 | f_{\vec{p}_A} f_{\vec{p}_B} \rangle|^2 = \int d\vec{p}_1 d\vec{p}_2 |f_{\vec{p}_A}(\vec{p}_1)|^2 |f_{\vec{p}_B}(\vec{p}_2)|^2$$

(assuming for simplicity distinguishable particles A & B) ✓

- All of this generalizes to n final state particles. Also, theorists generally leave the actual integration open for as long as possible and talk about differential cross sections ($d\sigma/d\dots$):

$$d\sigma = \prod_{j=1}^n d\vec{p}_j \int d^2b |\langle p_1 \dots p_n | S | i_b \rangle|^2$$

- Now plug in $|i_b\rangle$ as given above and use

$$S_{fi} = \delta_{fi} + i(2\pi)^4 \delta^4(p_f - p_i) \mathcal{M}_{fi}.$$

Unless $f = i$, only the $\mathcal{M}_{fi}$ -part contributes. Thus:

$$d\sigma = \prod_j d\tilde{p}_j \int d^2b \int d\tilde{k}_A d\tilde{k}_B \rho_{\bar{P}_A}(\tilde{k}_A) \rho_{\bar{P}_B}(\tilde{k}_B) \int d\tilde{k}'_A d\tilde{k}'_B \rho_{\bar{P}_A}^*(\tilde{k}'_A) \rho_{\bar{P}_B}^*(\tilde{k}'_B) \\ e^{i\bar{b}(\tilde{k}'_B - \tilde{k}_B)} |\mathcal{M}_{fi}|^2 (2\pi)^4 \delta^4(p_f - k_i) (2\pi)^4 \delta^4(p_f - k'_i)$$

where $p_f = \sum_j p_j$, $k_i = k_A + k_B$, $k'_i = k'_A + k'_B$

- The rest is a fairly straight forward calculation:

$$\rightarrow \int d^2b e^{i\bar{b}(\tilde{k}'_B - \tilde{k}_B)} = (2\pi)^2 \delta^2(k'_{B\perp} - k_{B\perp}) \quad \text{"}\perp\text{"} \triangleq \text{transverse (perpendicular)}$$

$$\rightarrow \int d^3k'_A d^3k'_B \delta^4(p_f - k'_i) \delta^2(k'_{B\perp} - k_{B\perp}) \dots$$

$$= \int d(k_A^{3'}) d(k_B^{3'}) \delta(p_f^3 - k_i^{3'}) \delta(p_f^0 - k_i^{0'}) \dots$$

- obviously, here $k_{B\perp}' = k_{B\perp}$ from now on

- less obviously, also $k_{A\perp}' = k_{A\perp}$ from now on

(This follows from $\delta^2(p_{f\perp} - k_{A\perp}' - k_{B\perp}')$ together with $\delta^2(p_{f\perp} - k_{A\perp} - k_{B\perp})$.)

$$= \int d(k_A^{3'}) \delta(p_f^0 - k_A^{0'} - k_B^{0'}) \dots$$

- obviously, now also $k_B^{3'} = p_f^3 - k_A^{3'}$

$$= \int d(k_A^{3'}) \delta(p_f^0 - \sqrt{m_A^2 + k_A^{1'2}} - \sqrt{m_B^2 + k_B^{1'2}}) \dots$$

with $\bar{k}_A^{-12} = k_{A\perp}^2 + (k_A^{3'})^2$; $\bar{k}_B^{-12} = k_{B\perp}^2 + (p_f - k_A^{3'})^2$.

$$\rightarrow = \frac{1}{\left| \frac{k_A^{3'}}{\sqrt{m_A^2 + \bar{k}_A^{-12}}} - \frac{k_B^{3'}}{\sqrt{m_B^2 + \bar{k}_B^{-12}}} \right|} \dots = \frac{1}{\left| \frac{k_A^{3'}}{k_A^{0'}} - \frac{k_B^{3'}}{k_B^{0'}} \right|} \dots \approx \frac{1}{|v_A - v_B|} \dots$$

here we use $k_{A,B} \approx p_{A,B}$, as
approximately implemented
by $f_{\bar{P}_A}(k_A)$; $f_{\bar{P}_B}(k_B)$.

Note: We talked about "fixed target" initially, but
nothing in our analysis depended on $v_A = 0$, so
we may as well keep it general.

- We have now completely carried out the $k_{A,B}^{3'}$
integrations, in the process implementing the relations

$$k_{A\perp}^{3'} = k_{A\perp}^{3'}; \quad k_{B\perp}^{3'} = k_{B\perp}^{3'}; \quad k_B^{3'} + k_A^{3'} = p_f^{3'}$$

together with

$$k_B^{0'} + k_A^{0'} = p_f^{0'}, \text{ i.e., } \sqrt{(k_B^{3'})^2 + k_{B\perp}^2 + m_B^2} + \sqrt{(k_A^{3'})^2 + k_{A\perp}^2 + m_A^2} = p_f^{0'}$$

- We can view the last two of these 4 relations as
two eqs. for the two variables $k_A^{3'}$, $k_B^{3'}$, which are
hence fixed unambiguously.
- Since these two eqs. also hold for k_A^3 , k_B^3 due to the
(not yet used) δ -fct. $\delta^4(p_f - k_A - k_B)$, we have:

$$\bar{k}_A' = \bar{k}_A \quad \& \quad \bar{k}_B' = \bar{k}_B.$$

- Thus, summarizing, we have arrived at

$$d\sigma = \prod_j d\tilde{p}_j \int d\tilde{k}_A d\tilde{k}_B |f_{\bar{p}_A}(\tilde{k}_A)|^2 |f_{\bar{p}_B}(\tilde{k}_B)|^2 |M_{fi}|^2 \cdot \frac{(2\pi)^4 \delta^4(p_f - k_A - k_B)}{4k_A^0 k_B^0 |v_A - v_B|}$$

- We can now view the two " $|f|^2$ " as effective δ -fcts. ensuring $\bar{k}_A = \bar{p}_A$ & $\bar{k}_B = \bar{p}_B$, thus

$$d\sigma = \underbrace{\frac{1}{4p_A^0 p_B^0 |v_A - v_B|}}_{\text{invariant}} |M_{fi}|^2 (2\pi)^4 \delta^4(p_f - p_i) \underbrace{\prod_{j=1}^n \frac{d^3 p_j}{(2\pi)^3 2p_j^0}}_{n\text{-particle phase space}}$$

This prefactor is invar.
under boosts along x^3
and can be written as

$$\frac{1}{2W(s, m_A^2, m_B^2)} \quad \text{where } W(x, y, z) \equiv \sqrt{x^2 + y^2 + z^2 - 2xy - 2xz - 2yz}$$

($\rightarrow$ problems). and $s \equiv (p_A + p_B)^2$ (centre-of-mass energy²)

- For us, at the moment the highly relativistic case will be most relevant. Thus, let $p^0 = m_A$; $\bar{p}_A = 0$

$$|v_A - v_B| = |v_B| = c = 1;$$

$$4p_A^0 p_B^0 = 2(p_A + p_B)^2 = 2s$$

$$\Rightarrow \boxed{d\sigma = \frac{1}{2s} |\mathcal{M}_{fi}|^2 dX^{(n)}} \quad \text{for the highly relativistic case}$$

$$\text{(where } dX^{(n)} = (2\pi)^4 \delta^4(\dots) d\tilde{p}_1 \dots d\tilde{p}_n \text{)}$$

§3 2-particle phase space & simple example

- Look at " $A+B \rightarrow 1+2$ " in $\lambda\phi^4$ -theory
- Focus just on phase space first:

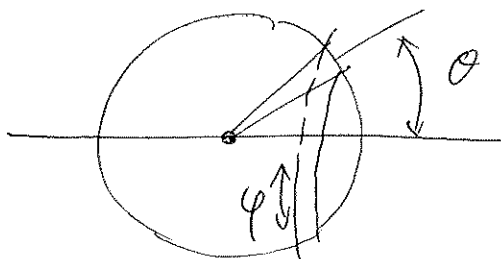
$$\int dX^{(2)} = \int (2\pi)^4 \delta^4(p_1 + p_2 - p_A - p_B) \frac{d^3 p_1}{(2\pi)^3 2p_1^0} \cdot \frac{d^3 p_2}{(2\pi)^3 2p_2^0}$$

Carry out $\int d^3 p_1$ trivially, and use $\vec{p}_1 = -\vec{p}_2$ (c.m.s.-frame)

$$\Rightarrow \int \frac{d^3 p_2}{(2\pi)^2 4p_1^0 p_2^0} \delta(p_1^0 + p_2^0 - \sqrt{s}) = \int \frac{d^3 p_2}{(2\pi)^2 4|\vec{p}_2|^2} \delta(2|\vec{p}_2| - \sqrt{s})$$

$$d^3 p_2 = d\Omega |\vec{p}_2|^2 d|\vec{p}_2|$$

$$d\Omega = d\varphi \sin\theta d\theta$$



$$\Rightarrow \int d|\vec{p}_2| \delta(2|\vec{p}_2| - \sqrt{s}) \cdot \frac{d\Omega}{16\pi^2} = \frac{d\Omega}{32\pi^2} \quad \text{(Interpreting "\int" without angular integration.)}$$

$$\Rightarrow \frac{d\sigma}{d\Omega} = \frac{|\mathcal{M}|^2}{64\pi^2 s} = \frac{\lambda^2}{64\pi^2 s}$$

(Note: $\frac{\lambda^2}{s^2}$ could have been argued without any calculation!)