

## 4. Perturbation Theory (Naive Approach)

20

### 4.1 Heisenberg picture

Since, in QFT, perturbation theory is usually developed in the "Interaction picture" (which is a "mixture" of Schrödinger- & Heisenberg picture), we first discuss the Heisenberg picture of the free theory:

- So far, our observables are (functions of)  $\varphi(\bar{x})$ ,  $\pi(\bar{x}) (= \dot{\varphi}(\bar{x}))$ , while our states are  $|p\rangle = \sqrt{2\omega_p} a_p^\dagger |0\rangle$

$$|p, q\rangle = \sqrt{2\omega_p} \sqrt{2\omega_q} a_p^\dagger a_q^\dagger |0\rangle$$

etc.

- To describe dynamics in the Schrödinger picture, we need to allow for time-evolution of states, e.g.

$$|p_t\rangle \equiv |p_{(0)}\rangle \text{ and } |p_t\rangle = e^{-iHt} |p_{(0)}\rangle \left( \begin{array}{l} \text{in general:} \\ |\psi_t\rangle = e^{-iHt} |\psi\rangle \end{array} \right)$$

We can then analyze the measurement of generic

observables  $O(\varphi, \pi)$  at some time, e.g.  $\langle \psi_t | O | \psi_t \rangle$ .

- For free fields, one could easily do that (we have diagonalized  $H$  anyway), but even here one can see that this is not natural in a Poinc.-invariant theory:
  - The  $t$ -dependence sits in  $|\psi_t\rangle$
  - The  $\bar{x}$ -dependence sits in  $O$  (via  $\varphi(\bar{x})$  &  $\pi(\bar{x})$ ).

- Better: (Heisenberg picture)

$$\langle \psi_t | O | \psi_t \rangle = \langle \psi | O_t | \psi \rangle \quad \text{with } |\psi\rangle - \text{time indep.} \\ \text{(in our case } |p\rangle \text{ etc.)}$$

$$\text{and } O_t = e^{iHt} O e^{-iHt}$$

- In our case, the crucial time-dep. operator is

$$\varphi(x) = \varphi(t, \vec{x}) = e^{iHt} \int \frac{d^3p}{(2\pi)^3} \cdot \frac{1}{\sqrt{2\omega_{\vec{p}}}} \left( a_{\vec{p}} e^{i\vec{p}\vec{x}} + a_{\vec{p}}^\dagger e^{-i\vec{p}\vec{x}} \right) e^{-iHt}$$

- To evaluate this, observe that  $H a_{\vec{p}} |q_1 \dots q_n\rangle = a_{\vec{p}} (H - E_{\vec{p}}) |q_1 \dots q_n\rangle$  since  $a_{\vec{p}}$  destroys a particle with energy  $E_{\vec{p}} = \omega_{\vec{p}}$ .

- Since our multi-particle states are by definition a basis, this holds in complete generality:

$$H a_{\vec{p}} = a_{\vec{p}} (H - \omega_{\vec{p}}).$$

$$\Rightarrow e^{iHt} a_{\vec{p}} e^{-iHt} = a_{\vec{p}} e^{i(H - \omega_{\vec{p}})t} e^{-iHt} = a_{\vec{p}} e^{-i\omega_{\vec{p}}t}$$

$$\& e^{iHt} a_{\vec{p}}^\dagger e^{-iHt} = \dots = a_{\vec{p}}^\dagger e^{i\omega_{\vec{p}}t}$$

$$\Rightarrow \varphi(x) = \int \frac{d^3p}{(2\pi)^3} \cdot \frac{1}{\sqrt{2p^0}} \left( a_{\vec{p}} e^{-ipx} + a_{\vec{p}}^\dagger e^{ipx} \right) \quad \text{with } px = p^0 x^0 - \vec{p}\vec{x}$$

$$\& p^0 = \omega_{\vec{p}} = E_{\vec{p}}$$

We will now always use this definition of  $p^0$  (unless it is an independent variable).

Note: The same procedure can be used to calculate  $\pi(x) = \pi(t, \vec{x})$ .

We will not really need  $\pi(x)$  since it turns out that

$\pi(x) = \dot{\varphi}(x)$  holds at the operator level. Furthermore,

$(\square + m^2) \varphi(x)$  holds at the operator level (obvious from the above).

- As an important application, we can establish

Causality:

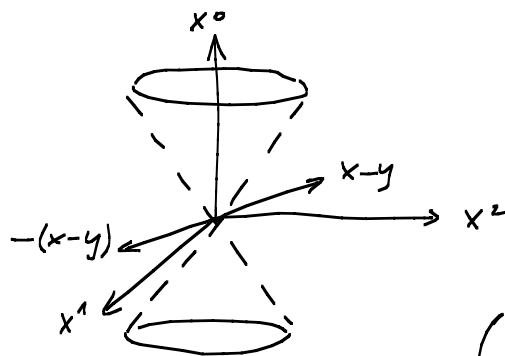
$$[\varphi(x), \varphi(y)] = 0 \quad \text{if } (x-y)^2 < 0 \quad [\text{i.e. the measurements of } \varphi \text{ at}$$

two points with space-like separation do not influence each other] <sup>22</sup>

• Demonstration:

$$\begin{aligned}
 [\varphi(x), \varphi(y)] &= \int \frac{d^3p}{(2\pi)^3 2p^0} \int \frac{d^3q}{(2\pi)^3 2q^0} \left\{ [a_{\vec{p}}, a_{\vec{q}}^+] e^{-ipx + iqy} \right. \\
 &\quad \left. + [a_{\vec{p}}^+, a_{\vec{q}}] e^{ipx - iqy} \right\} \\
 &= \int \frac{d^3p}{(2\pi)^3 2p^0} e^{-ip(x-y)} - \int \frac{d^3p}{(2\pi)^3 2p^0} e^{ip(x-y)} \\
 &= \int \frac{d^4p}{(2\pi)^3} \delta(p^2 - m^2) \Big|_{p^0 > 0} e^{-ip(x-y)} - \int \frac{d^4p}{(2\pi)^3} \delta(p^2 - m^2) \Big|_{p^0 > 0} e^{ip(x-y)}
 \end{aligned}$$

- Each term is manifestly invariant under the special Lorentz group (all  $SO(1,3)$  tfs. which can be linked to  $\mathbb{1}$  by continuous deformations).
- This group allows us to transform  $x-y$  (with  $(x-y)^2 < 0$ ) into  $-(x-y)$ :



(Problem: prove this!)

- Hence, the two terms cancel and  $[\varphi(x), \varphi(y)] = 0$  for  $(x-y)^2 < 0$ .

Note: One could also prove " $[\varphi(x), \varphi(y)] = 0$  if  $(x-y)^2 < 0$ " starting from " $[\varphi(x), \varphi(y)] = 0$  if  $x^0 = y^0$ ", which is known from canonical quantization. The generalization to  $x^0 \neq y^0$  can be realized using Lorentz-symm., more specifically:

$$\hat{\Lambda} \varphi(\Lambda^{-1}x) \hat{\Lambda}^{-1} = \varphi(x).$$

Here  $\hat{\Lambda}$  is the repres. of the Lorentz-trf.  $\Lambda$  on the Hilbert space and the equation above is the obvious requirement of Lorentz covariance of our field operators. However, while physically obvious, it does require demonstration (see e.g. Itzykson/Zuber).

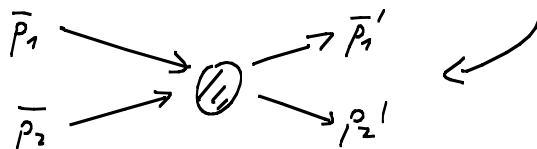
## 4.2 S-Matrix (scattering matrix)

- The simplest modification allowing for scattering is

$$H_0 \rightarrow H_0 + H_{int} ; \quad H_{int} = \frac{\lambda}{4!} \varphi^4 = a_{\vec{p}_2'}^+ a_{\vec{p}_1'}^+ a_{\vec{p}_2}^- a_{\vec{p}_1}^-$$

(equivalently,

$$\mathcal{L}_0 \rightarrow \mathcal{L}_0 + \mathcal{L} ; \quad \mathcal{L}_{int} = -H_{int})$$



(We exclude a possible term  $\sim \varphi^3$

- physically: because  $V(\varphi)$  is not bounded from below
- technically: by demanding that  $\mathcal{L}$  is symmetric w.r.t.  $\varphi \rightarrow -\varphi$ .)

- For the technical treatment, we work in the "Interaction Picture", in which operators evolve as in the Heisenberg picture, but with the free Hamiltonian,

$$O_t = e^{iH_0 t} O_0 e^{-iH_0 t},$$

while states carry the non-trivial  $H_{\text{int}}$ -evolution

$$|\psi_t\rangle = e^{iH_0 t} e^{-iH t} |\psi_0\rangle.$$

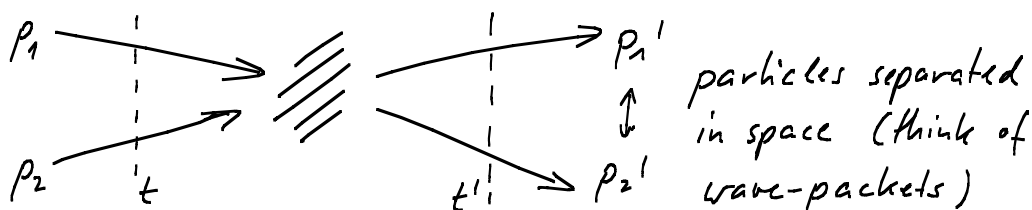
- The evolution of an interaction-picture state  $|\psi_t\rangle$  into an interaction picture state  $|\psi_{t'}\rangle$  is then given by

$$\begin{aligned} |\psi_{t'}\rangle &= e^{iH_0 t'} e^{-iH t'} |\psi_0\rangle = e^{iH_0 t'} e^{-iH t'} (e^{iH_0 t} e^{-iH t})^{-1} |\psi_t\rangle \\ &= \underbrace{e^{iH_0 t'} e^{-iH(t'-t)} e^{-iH_0 t}}_{\equiv U(t', t)} |\psi_t\rangle \\ &\equiv U(t', t) |\psi_t\rangle \end{aligned}$$

This is called the "S-matrix"

(It simply corresponds to the unitary evolution of states in the interaction picture; cf. time-dependent pert. theory in usual QM)

- Our main physical application will be 2-2-scattering (more generally, 2-n-scattering):



- The amplitude for this process is given by the projection of  $U(t', t) |p_1 p_2\rangle$  onto a state  $|p'_1 p'_2\rangle$  (with different momenta!):

"S-matrix element"  $\rightarrow S_{fi} = \langle p'_1 p'_2 | U(t', t) | p_1 p_2 \rangle$

$\uparrow$  final     $\uparrow$  initial

where  $|p_1 p_2\rangle = |p_1 p_2, t\rangle = a_{\vec{p}_1}^+ a_{\vec{p}_2}^+ |0\rangle$

Here we assume that  $H_I$  is irrelevant at early times so that we can use the previously discussed Heisenberg picture

(This is actually not correct (hence "naïve approach") since  $H_I$  affects even the evolution of a spatially separated single particle. We will treat this properly later on (LSZ-formalism).)

Analogously:  $|p'_1 p'_2\rangle = a_{\vec{p}'_1}^+ a_{\vec{p}'_2}^+ |0\rangle$

(Note that we use our "second" normalization convention for the  $a$  &  $a^+$ .)

- To proceed, we rewrite  $U(t', t)$ :

$$\begin{aligned}
 U(t', t) &= e^{iH_0 t'} e^{-iH(t'-t)} e^{-iH_0 t} \\
 &= e^{iH_0 t'} \underbrace{e^{-iH\Delta} e^{-iH\Delta} \dots e^{-iH\Delta}}_{n \text{ times}} e^{-iH_0 t} \quad \left( \Delta = \frac{t'-t}{n} \right) \\
 &= e^{iH_0 t'} e^{-iH\Delta} \underbrace{e^{-iH_0(t'-\Delta)} e^{iH_0(t'-\Delta)}}_{=1} e^{-iH\Delta} \underbrace{e^{-iH_0(t'-2\Delta)} \dots}_{=1 \text{ etc.}}
 \end{aligned}$$

$\nwarrow$  These are products, not arguments!

- Using  $H = H_0 + H_{int}$  and the smallness of  $\Delta$  we write

$$e^{iH_0 t'} e^{-iH_0 \Delta} e^{-iH_0 (t' - \Delta)} = e^{iH_0 t'} e^{-iH_0 \Delta} e^{iH_0 \Delta} e^{-iH_0 t'}$$

$$\approx e^{iH_0 t'} e^{-iH_{int} \Delta} e^{-iH_0 t'} = e^{-\underbrace{iH_{int}(t') \cdot \Delta}_{\text{time-dependent operator in int.-picture}}}$$

$$\Rightarrow U(t', t) = e^{-iH_{int}(t') \Delta} e^{-iH_{int}(t' - \Delta) \Delta} \dots e^{-iH_{int}(t) \Delta}$$

$$= T \exp \left[ -i \int_t^{t'} d\tau H_{int}(\tau) \right]$$

$\uparrow$  time-ordering "operator"       $\uparrow$   $H_{int}$  in interaction picture

(Definition of  $T$ :  $T \varphi(t_1) \varphi(t_2) = \begin{cases} \varphi(t_1) \varphi(t_2) & \text{if } t_1 > t_2 \\ \varphi(t_2) \varphi(t_1) & \text{if } t_2 > t_1 \end{cases}$ )

- Thus, all we need is

$$H_{int}(t) = e^{iH_0 t} \int d^3x \frac{1}{4!} \varphi^4(\vec{x}) e^{-iH_0 t} = \int d^3x \frac{1}{4!} \varphi^4(x)$$

$\uparrow$   
 our familiar space-time  
 dependent quantum field  
 (of the free theory)

- Now we take  $t' \rightarrow +\infty$  &  $t \rightarrow -\infty$  and define

$$S = T e^{-i \int dt H_{int}(t)} = T e^{i \int d^4x \mathcal{L}_{int}(x)}$$

$$\text{and } S_{fi} = \langle 0 | a_{\vec{p}_1} a_{\vec{p}_2} T e^{i \int d^4x \mathcal{L}_{int}} a_{\vec{p}_1}^+ a_{\vec{p}_2}^+ | 0 \rangle$$

Note: only free fields in this formula!

- Our  $S$ -matrix clearly has a trivial part corresponding to the "1" in  $e^{i \dots} = 1 + \dots$ . Hence, we define the transition or

$T$ -matrix via

$$S = \mathbb{1} + iT \quad \& \quad S_{fi} = \delta_{fi} + iT_{fi}.$$

- At leading non-trivial order in  $\lambda$  we get

$$iT_{fi} = \langle 0 | a_{\vec{p}_1'} a_{\vec{p}_2'} \left( -\frac{i\lambda}{4!} \right) \int d^4x \varphi^4(x) a_{\vec{p}_1}^+ a_{\vec{p}_2}^+ | 0 \rangle$$

$$\text{with } \varphi(x) = \int \frac{d^3p}{(2\pi)^3 2p^0} \left( a_{\vec{p}} e^{-ipx} + a_{\vec{p}}^+ e^{ipx} \right)$$

$$\text{and } [a_{\vec{p}}, a_{\vec{q}}^+] = 2p^0 (2\pi)^3 \delta^3(\vec{p} - \vec{q}). \quad (\text{Note the normalization!})$$

- $T_{fi}$  can now be worked out by using the commutator to push the  $a^+$ s to the left (the  $a$ 's to the right) until they annihilate the vacuum. The result (see problems) reads

$$iT_{fi} = -i\lambda (2\pi)^4 \delta^4(p_1 + p_2 - p_1' - p_2').$$

- The momentum-conservation  $\delta$ -fct. (which always appears in this context) motivates the definition of the "invariant matrix element"  $\mathcal{M}$ :

$$S_{fi} = \mathbb{1} + i(2\pi)^4 \delta^4(\dots) \mathcal{M}$$

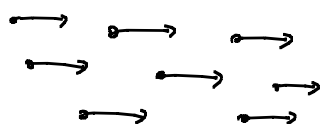
$$\text{with } \underline{\underline{i\mathcal{M}}} = -i\lambda.$$

- This is our first "Feynman rule":

$$\begin{array}{c} p_1 \nearrow \\ p_2 \nearrow \end{array} \begin{array}{c} \diagdown \\ \diagdown \end{array} \begin{array}{c} p_1' \\ p_2' \end{array} \quad " = " \quad -i\lambda$$

### 4.3 Scattering cross section

Consider fixed-target experiment:



particle stream (type B)

- target particle  
(type A)

If the beam is homogeneous (transversely) with transverse area  $F$ ,

$$\frac{N_{\text{events}}}{N_B} = \frac{\sigma}{F} \quad \text{or} \quad \sigma = \frac{N_{\text{events}}}{(N_B/F)}$$

↑  
transverse beam density.

To calculate this, we will need initial states localized in space (at least transversely). Consider wave packets:

$$|f_{\bar{p}}\rangle = \int d\tilde{k} f_{\bar{p}}(\tilde{k}) |\tilde{k}\rangle \quad \text{with} \quad d\tilde{k} \equiv \frac{d^3k}{(2\pi)^3 2k^0}$$

(Here the fct.  $f_{\bar{p}}(\tilde{k})$  is peaked near  $\tilde{k} = \bar{p}$ ; think e.g. of an appropriately shifted Gaussian.)

Normalization:  $\langle f_{\bar{p}} | f_{\bar{p}} \rangle = \int d\tilde{k} d\tilde{k}' f_{\bar{p}}(\tilde{k}) f_{\bar{p}}^*(\tilde{k}') \cdot \langle \tilde{k}' | \tilde{k} \rangle$

$$= \int d\tilde{k} |f_{\bar{p}}(\tilde{k})|^2 \stackrel{!}{=} 1$$

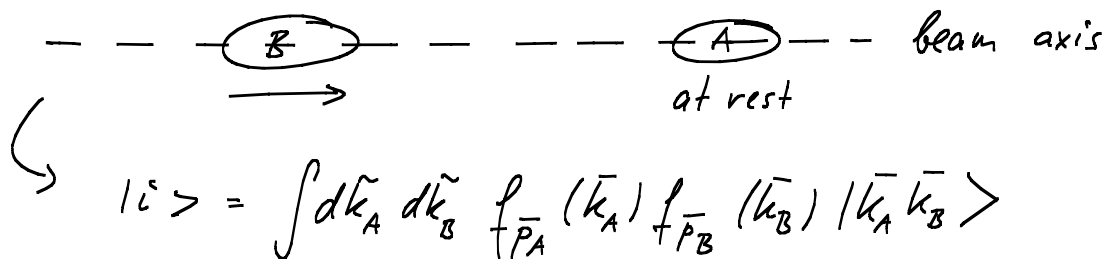
Using  $a_{\tilde{k}}^+ = \int d^3x e^{i\tilde{k}\tilde{x}} \{ k_0 \varphi(\tilde{x}) - i\pi(\tilde{x}) \}$ , we have

$$|f_{\bar{p}}\rangle = \int d^3x \left\{ \int d\tilde{k} e^{i\tilde{k}\tilde{x}} f_{\bar{p}}(\tilde{k}) k^0 \right\} \varphi(\tilde{x}) |0\rangle$$

$$- i \int d^3x \left\{ \int d\tilde{k} e^{i\tilde{k}\tilde{x}} f_{\bar{p}}(\tilde{k}) \right\} \pi(\tilde{x}) |0\rangle,$$

which shows that particles can really be localized in  $\tilde{x}$  while  $f_{\bar{p}}$  is appropriately localized in  $\tilde{k}$ . (Recall that the Fourier transform of a Gaussian is again a Gaussian.)

- Thus, we can find functions  $f_{\bar{p}_A}$  &  $f_{\bar{p}_B}$  characterizing a situation like this:



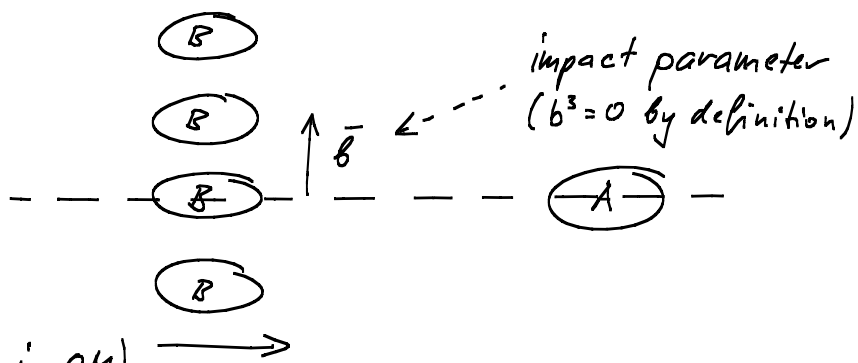
- more realistic:

- the "shifted" B's

can be generated using

$\hat{P}_x$  (which generates

shifts in  $x$ , just as in QM)



- To see this more explicitly, recall:  $H a_k^+ = a_k^+ (H + k_0)$

analogously:  $P^i a_k^+ = a_k^+ (P^i + k^i)$ .

Thus:  $e^{-i\tilde{P}\tilde{b}} \left( \int d\tilde{k} f_{\tilde{q}}(\tilde{k}) a_k^+ |0\rangle \right) = \int d\tilde{k} f_{\tilde{q}}(\tilde{k}) e^{-i\tilde{k}\tilde{b}} a_k^+ |0\rangle$

Using  $a_k^+ = \int d^3x e^{i\tilde{k}\tilde{x}} \{ \varphi(\tilde{x}) k^0 - i\pi(\tilde{x}) \}$

We find  $\int d^3x \left\{ d\tilde{k} f_{\tilde{q}}(\tilde{k}) e^{i\tilde{k}(\tilde{x}-\tilde{b})} \right\} \{ \varphi(\tilde{x}) k^0 - i\pi(\tilde{x}) \} |0\rangle$

We see that the  $\uparrow$  fct. defining the spatial distribution of the state is indeed shifted by  $\tilde{b}$ .

- Thus, an initial state with fixed impact parameter  $\tilde{b}$  is

$$|i_b\rangle = \int d\tilde{k}_A d\tilde{k}_B f_{\tilde{P}_A}(\tilde{k}_A) f_{\tilde{P}_B}(\tilde{k}_B) e^{-i\tilde{k}_B \tilde{b}} |\tilde{k}_A \tilde{k}_B\rangle$$

- Our transition amplitude then reads  $\langle p_1 p_2 | S | i_b \rangle$

(We changed  $p_1' p_2'$  to  $p_1 p_2$  since we now use  $p_A p_B$  for the

initial momenta) and the corresponding probability naively reads  $|\langle p_1 p_2 | S | i_b \rangle|^2$ . For many incoming particles, we expect a # of events given by

$$N_{\text{events}} = \sum_b |\langle p_1 p_2 | S | i_b \rangle|^2$$

or, in a dense beam,  $N_{\text{events}} = \frac{N_B}{F} \int d^2b |\langle p_1 p_2 | S | i_b \rangle|^2$

- The corresponding cross-section would be

$$\sigma(p_1, p_2) = \frac{N_{\text{events}}}{(N_B/F)} = \int d^2b |\langle p_1 p_2 | S | i_b \rangle|^2$$

- However, this clashes with the  $\delta$ -function-normalization of  $\langle p_1$ ,  $\langle p_1 p_2$ , etc. Furthermore, no realistic detector can resolve momenta with infinite precision.
- Hence, we need to sum (integrate) over a small region final-state phase-space:

$$\sigma(V_f) = \int_{V_f} d\tilde{p}_1 d\tilde{p}_2 \int d^2b |\langle p_1 p_2 | S | i_b \rangle|^2.$$

- The fact that  $\int d\tilde{p}_1 = \int \frac{d^3 p_1}{(2\pi)^3 2 p_1^0}$  is the proper normalization

for this integration follows from checking that the probability is 1 for a single particle falling into the right part of phase space:

$$w = \int d\tilde{p} |\langle p | f_{\tilde{\eta}} \rangle|^2 = \int d\tilde{p} |f_{\tilde{\eta}}(\bar{p})|^2 = 1$$

according to our definition of  $f_{\tilde{\eta}}(\bar{k})$ . ✓

Instead of working with some  $\sigma(V_f)$ , we will use a differential formulation

$$d\sigma = \prod_{j=1}^n d\tilde{p}_j \int d^2b \left| \langle p_1 \dots p_n | S | i_b \rangle \right|^2$$

for  $n$  final-state particles.

- With our explicit formula for  $|i_b\rangle$  and  $S_{fi} = \delta_{fi} + i(2\pi)^4 \delta^4(p_f - p_i) \mathcal{M}_{fi}$  we find (for non-trivial scattering):

$$d\sigma = \prod_j d\tilde{p}_j \int d^2b \int d\tilde{k}_A d\tilde{k}_B f_{PA}(\tilde{k}_A) f_{PA}(\tilde{k}_B) \int d\tilde{k}'_A d\tilde{k}'_B f_{PA}^*(\tilde{k}'_A) f_{PB}^*(\tilde{k}'_B) e^{i\bar{b}(\tilde{k}'_B - \tilde{k}_B)} |\mathcal{M}_{fi}|^2 (2\pi)^4 \delta^4(p_f - k_i) (2\pi)^4 \delta^4(p_f - k'_i),$$

where  $p_f = \sum_j p_j$   
 $k_i^{(1)} = k_A^{(1)} + k_B^{(1)}$ .

Calculate...

- $\int d^2b e^{i\bar{b}(\tilde{k}'_B - \tilde{k}_B)} = (2\pi)^2 \delta^2(k'_{B\perp} - k_{B\perp})$
- $\int d^3k'_A d^3k'_B \delta^4(p_f - k'_i) \delta^2(k'_{B\perp} - k_{B\perp}) = \int d^3k'_A d^3k'_B \delta(p_f^3 - k_i^3) \delta(p_f^0 - k_i^0)$   
 [From now on  $k'_{B\perp} = k_{B\perp}$  &  $k'_{A\perp} = k_{A\perp}$ , where we also used  $\delta^2(p_{f\perp} - k_{i\perp})$  from the second  $\delta^4$ -fct. above.]  
 $= \int d^3k'_A \delta(p_f^0 - k_A^0 - k_B^0)$  [from now on  $k_B^3 = p_f^3 - k_A^3$ ]  
 $= \int d^3k'_A \delta(p_f^0 - \sqrt{m_A^2 + \vec{k}_A'^2} - \sqrt{m_B^2 + \vec{k}_B'^2})$   
 $\quad \quad \quad \uparrow \quad \quad \quad \uparrow$   
 $\quad \quad \quad k_{A\perp}^2 + (k_A^3)^2 \quad \quad \quad k_{B\perp}^2 + (p_f^3 - k_A^3)^2$

$$= \frac{1}{\left| \frac{k_A^3}{k_A^0} - \frac{k_B^3}{k_B^0} \right|} \approx \frac{1}{|v_A - v_B|}$$

$k_{A,B} \approx p_{A,B}$  for these purposes

- After all these  $\delta$ -fct. integrations,  $\bar{k}_A'$  &  $\bar{k}_B'$  are fixed by

$$k_{A\perp}' = k_{A\perp}; \quad k_{B\perp}' = k_{B\perp}; \quad p_f^3 - k_A'^3 - k_B'^3 = 0; \quad p_f^0 - k_A'^0 - k_B'^0 = 0.$$

- Together with  $p_f^3 - k_A^3 - k_B^3 = 0$  and  $p_f^0 - k_A^0 - k_B^0 = 0$  from  $\delta^4(p_f - k_i)$ , which have not yet used, this is equivalent to

$$\bar{k}_A' = \bar{k}_A; \quad \bar{k}_B' = \bar{k}_B.$$

- Hence we have

$$d\mathcal{G} = \prod_j d\tilde{p}_j \int d\tilde{k}_A d\tilde{k}_B |f_{\bar{p}_A}(\bar{k}_A)|^2 |f_{\bar{p}_B}(\bar{k}_B)|^2 |\mathcal{M}_{fi}|^2 \cdot \frac{1}{4k_A^0 k_B^0 |v_A - v_B|} \cdot (2\pi)^4 \delta^4(p_f - k_i).$$

- We now perform an integration over a sufficiently large path of final phase space to make effectively the  $\delta^4(p_f - \dots)$ -fct. very smeared. (But we keep the  $d\tilde{p}_j$ -notation for simplicity.)
- Now the localization of  $f_{\bar{p}_A}$  &  $f_{\bar{p}_B}$  is more sharp than the  $\delta^4$ -fct. This allows us to perform the  $\bar{k}_A$  &  $\bar{k}_B$ -integrations and simply replace  $k_i$  in the  $\delta^4$ -fct. by  $p_i$ .

- Using  $\int d\tilde{k}_{A,B} |f_{\bar{P}_{A,B}}(\tilde{k}_{A,B})|^2 = 1$ , we have

$$d\sigma = \underbrace{\frac{1}{4p_A^0 p_B^0 |\vec{v}_A - \vec{v}_B|}}_{\text{prefactor (invariant under boosts along beam axis)}} |M_{fi}|^2 \underbrace{(2\pi)^4 \delta^4(p_f - p_i) \prod_{j=1}^n \frac{d^3 p_j}{(2\pi)^3 2p_j^0}}_{n\text{-particle phase space}}$$

prefactor (invariant under boosts along beam axis)

$n$ -particle phase space

which, for  $S = (p_A + p_B)^2 \gg m_A^2, m_B^2$ ,

can be replaced by

$$\frac{1}{2s}$$

( $\sqrt{s} \hat{=}$  centre-of-mass energy)

$\Rightarrow$  Highly relativistic case:

$$d\sigma = \frac{1}{2s} |M_{fi}|^2 dX^{(n)} \quad \text{with } (2\pi)^4 \delta^4(\dots) \text{ \& } d\tilde{p}_j$$

### Problems:

- Demonstrate that  $\frac{1}{4 p_A^0 p_B^0 |v_A - v_B|} = \frac{1}{4 |p_A^3 p_B^0 - p_B^3 p_A^0|}$  is invariant

under boosts along the beam axis.

Hint:

Characterize 4-vectors not by  $(p^0, p_\perp, p^3)$ , but rather by  $(p^+, p^-, p_\perp)$ , where  $p^\pm \equiv p^0 \pm p^3$ .

( $\equiv$  light-cone momenta). Work out the transformations of  $p^+, p^-$  under boosts along  $x^3$ . Show that

$$p_A^+ p_B^- = p_A \cdot p_B + p_A^3 p_B^0 - p_A^0 p_B^3 \quad (\text{for } p_{A,\perp} = p_{B,\perp} = 0)$$

The claim should then become obvious.

- Show that the above prefactor of the cross section can also be written as  $\frac{1}{2W(s, m_A^2, m_B^2)}$  [This is most easily done in the target-rest-frame. Cf. Nadtman's Book.

$$\text{where } W(x, y, z) \equiv \sqrt{x^2 + y^2 + z^2 - 2xy - 2xz - 2yz},$$

thus bringing the cross section to a manifestly boost-inv. form.

### 4.4 2-particle phase space and a simple example

- Consider the simple case  $A + B \rightarrow 1 + 2$  in  $\lambda\varphi^4$ -theory with  $s \gg m$ . Ignore the matrix element and the prefactor of the cross-section for the moment and focus just on the phase space:

$$\int dX^{(2)} = \int (2\pi)^4 \delta^4(p_1 + p_2 - p_A - p_B) \frac{d^3 p_1}{(2\pi)^3 2p_1^0} \cdot \frac{d^3 p_2}{(2\pi)^3 2p_2^0}$$

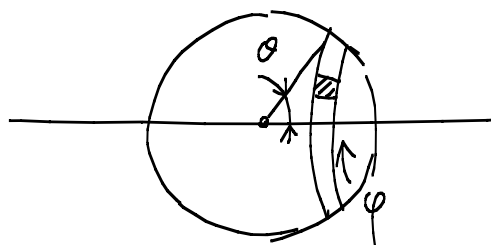
- Perform  $\int d^3 p_1$  using  $\delta^3(\vec{p}_1 + \dots)$ . Working in the cms-system ( $\vec{p}_1 = -\vec{p}_2$ ), we are left with

$$\int \frac{1}{(2\pi)^2 4 p_1^0 p_2^0} d^3 \vec{p}_2 \delta(p_1^0 + p_2^0 - \sqrt{s}) = \int \frac{d^3 \vec{p}_2}{(2\pi)^2 4 |\vec{p}_2|^2} \delta(2|\vec{p}_2| - \sqrt{s})$$

$$= |\vec{p}_1| = |\vec{p}_2| \quad = |\vec{p}_2|$$

Use also  $d^3 p_2 = d\Omega |\vec{p}_2|^2 d|\vec{p}_2|$ , where

$$d\Omega = d\varphi \sin\theta d\theta$$



$$\Rightarrow \int dX^{(2)} = \frac{1}{(2\pi)^2 s} d\Omega \cdot (\sqrt{s}/2)^2 \cdot \frac{1}{2},$$

where we decided not to integrate over the angles.

Remembering  $|\mathcal{M}|^2 = \lambda^2$ , we find

$$\frac{d\sigma}{d\Omega} = \frac{1}{2s} |\mathcal{M}|^2 \cdot \frac{\int dX^{(2)}}{d\Omega} = \frac{\lambda^2}{64\pi^2 s} \quad \text{— indep. of angle!}$$

(because of spin-0 particles).

- Note:  $\frac{\lambda^2}{s}$  could have been argued just on dimensional grounds.