

6 QCD - Running Coupling & β -Function

6.1 General Structure & Running coupling

- $SU(3)$ gauge-th. with (several generations) of fermions (quarks) in fund. repr.:

$$\mathcal{L} = -\frac{1}{2g^2} \text{tr} F_{\mu\nu} F^{\mu\nu} + \sum_f \bar{\Psi}_f (i\not{D} - m) \Psi_f$$

$$f \in \underbrace{\{u, d, s, c, b, t\}}_{\text{"light"}} \quad \underbrace{\quad}_{\text{"heavy"}}$$

- $A_\mu = A_\mu^a T^a$; $T^a = \frac{1}{2} \lambda^a$ "Gell-Mann matrices"

$$\lambda^a = \left(\begin{array}{c|c} \sigma^a & \\ \hline & \end{array} \right) ; \quad a = 1 \dots 3$$

$$\lambda^4 = \left(\begin{array}{cc} 1 & 0 \\ 0 & -1 \end{array} \right) ; \quad \lambda^5 = \left(\begin{array}{cc} 0 & 1 \\ 1 & 0 \end{array} \right) ; \quad \lambda^6 = \left(\begin{array}{cc} 1 & 0 \\ 0 & -1 \end{array} \right) ; \quad \lambda^7 = \left(\begin{array}{cc} 0 & 1 \\ 1 & 0 \end{array} \right)$$

$$\lambda^8 = \frac{1}{\sqrt{3}} \left(\begin{array}{cc} 1 & 0 \\ 0 & -2 \end{array} \right) \quad ([\lambda^3, \lambda^8] = 0 ; \text{"Cartan subalgebra"})$$

- Let us do pert. theory (according to Feynman rules of Ch. 5), using dimensional regularization (cf. also Ch. 11 of QFT I)
- In d dims., g^2 is not dim. less:

$$D_\mu = \partial_\mu + i A_\mu \Rightarrow [A_\mu] = 1$$

$$S = \int d^d x \frac{1}{2g^2} \text{tr}(F^2) \Rightarrow -d - [g^2] + 2 + 2[A_\mu] = 0$$

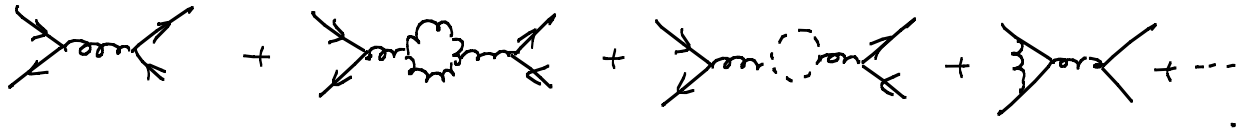
$$\Rightarrow [g^2] = 4 - d$$

- We write $d = 4 - \epsilon$ and redefine $g^2 \rightarrow g^2 \mu^\epsilon$, where μ with $[\mu] = 1$ is a new parameter.

- Calculate some dim. less phys. observable (e.g. σ for

$q\bar{q} \rightarrow q\bar{q}$, for heavy quarks of mass m at c.m.s.

energy Q ($s = Q^2$), normalized by appropriate power of Q)



- Result:

$$R = R(Q^2/\mu^2, g^2)$$

$$= \left[g^2 \mu^\epsilon + g^4 \mu^{2\epsilon} f(Q^2, \epsilon) + O(g^6) \right] Q^{-\epsilon}$$

Prefactor 1
for simplicity

This follows from Feynman diagrams (1-loop)
in $4-\epsilon$ dims.; μ does not appear!

$$\Rightarrow f(Q^2, \epsilon) = Q^{-\epsilon} \left(c_1 \frac{1}{\epsilon} + c_2 + c_3 \epsilon + \dots \right)$$

[A typical loop-integral is

$$\int \frac{d^d k}{(k^2 + Q^2)^2} \sim Q^{d-4} \int \frac{d^d x}{(x^2 + 1)^2} \sim Q^{-\epsilon} \left(\dots \frac{1}{\epsilon} + \dots \right)]$$

$$\Rightarrow R = \left[g^2 + g^4 e^{\epsilon \ln \mu/Q} \left(c_1 \frac{1}{\epsilon} + c_2 + O(\epsilon) \right) \right] \left(\frac{\mu}{Q} \right)^\epsilon$$

$$= \left[g^2 + g^4 \left(c_1 \frac{1}{\epsilon} + c_2 + c_1 \ln \mu/Q + O(\epsilon) \right) \right] \left(\frac{\mu}{Q} \right)^\epsilon$$

- We now need to renormalize. A very convenient

renormalization scheme is "minimal subtraction":

- Redefine $g = g_{\text{bare}}$ as $g_{\text{bare}} = g_{\text{phys}} Z_g$.

Z_g must be a fct. of ϵ , chosen such that R remains finite as $\epsilon \rightarrow 0$. We choose:

$$\boxed{Z_g = 1 + \frac{\#}{\epsilon}} \quad (\text{"minimal", since just the pole is subtracted.})$$

- In our case, obviously $Z_g = 1 - \frac{1}{2} c_1 \frac{1}{\epsilon}$.

- We find: $R = [g^2 (1 + g^2 (c_2 + c_1 \ln \mu/Q)) + O(g^4)] \underbrace{\left(\frac{\mu}{Q}\right)^\epsilon}_{=1 \text{ since } \epsilon \rightarrow 0}$
 (at $\epsilon \rightarrow 0$ and with $g_{\text{phys}} \rightarrow g$)
 (Beware, from now on g is our phys. coupling.)

- This works to all orders in g and we find

$$\underline{\underline{R = R(g^2, Q^2/\mu^2)}}$$

[Note: We have chosen not to renormalize our fields, i.e. not to substitute $A_\mu \rightarrow A_\mu Z_A^{1/2}$. In this way, we don't get finite Green's fcts., but phys. observables can nevertheless all be made finite. A redefinition of fields would not affect the result since it affects both the vertices and the Z -factors of LSZ. These effects cancel in cross sects.]

- Obviously, the $\overline{\text{MS}}$ -scheme is, in fact, a 1-parameter family of schemes (with the parameter being μ).

- Of course, R can not depend on which scheme one uses, i.e., it can not depend on μ . This can only be the case if $g = g_{\text{phys.}}$ depends on μ , i.e.

$$0 = \underbrace{\mu^2 \frac{d}{d\mu^2}}_{\frac{d}{d \ln \mu^2}} R(g^2(\mu^2), Q^2/\mu^2)$$

$$\Rightarrow 0 = \mu^2 \left(\frac{\partial}{\partial \mu^2} + \frac{dg^2}{d\mu^2} \frac{\partial}{\partial g^2} \right) R(g^2, Q^2/\mu^2)$$

$$\text{or } \mu^2 \frac{dg^2}{d\mu^2} = - \left(\mu^2 \frac{\partial}{\partial \mu^2} R \right) / \left(\frac{\partial}{\partial g^2} R \right)$$

This expression does not depend on Q^2/μ^2 (since it describes the change of the coupling with the change of the renorm. scheme – this should not depend on the observable we were calculating.)

$$\Rightarrow \mu \frac{dg}{d\mu} = \beta(g)$$

calculable order-by-order in g .

- In our example:

$$\frac{dg^2}{d \ln \mu^2} = - \frac{g^4 \cdot \frac{c_1}{2}}{1} + O(g^6)$$

↖ we can only use the leading order since numerator is leading order.

$$\Rightarrow \beta(g) = -\frac{c_1}{2} g^3 + \dots$$

Fact: (to be derived later)

For $SU(N_c)$ gauge th. with N_f flavors:

$$\beta(g) = -\frac{g^3}{(4\pi)^2} \left(\frac{11}{3} N_c - \frac{2}{3} N_f \right)$$

$$\text{(cf. } \beta(e) = \frac{e^3}{12\pi^2} \text{)}$$

Problem: Can you get $\beta(e)$ if you are given $\beta(g)$, without doing any loop calc.?

- To get an optimal result at fixed order, choose $\mu = Q$ (to avoid "large logs").

$$\Rightarrow R = R(g^2(Q^2), 1)$$

- Actually, the true expansion parameter is

$$\frac{\alpha_s}{4\pi} \cdot N_c = \underbrace{\frac{g^2}{16\pi^2}}_{\text{"loop factor"}} \underbrace{N_c}_{\text{each loop contributes an extra } N_c \text{ for } N_c \gg 1.}$$

$$\int \frac{d^4 k}{(2\pi)^4} = \int \frac{k^3 dk \cdot \text{Vol}_{S^3}}{(2\pi)^4} = \frac{1}{16\pi^2} \int k^2 d(k^2)$$

$\uparrow$ $\leftarrow 2\pi^2$

$$\Rightarrow R = R(\alpha_s(Q^2), 1)$$

$$\alpha_s \sim \frac{1}{\ln Q^2/\Lambda_{\text{QCD}}^2} \quad \begin{array}{l} \text{(check that this solves} \\ \text{the above RGE; easiest} \\ \text{by looking at RGE for } \alpha_s^{-1}) \\ \uparrow \\ \text{comes from boundary} \\ \text{condition of the diff. eq.} \end{array}$$

- We can trust all this only for large Q^2 , where α_s is small. (e.g. for $Q^2 = m_Z^2$; $\alpha_s(m_Z) \approx 0.12$)
- $\Lambda_{\text{QCD}} \approx 0.2 \text{ GeV}$ (depends strongly on scheme).
- The leading-order β -fct. is scheme-indep. (cf. our QED-derivation based only on discussion of divergences).
So is the qualitative behaviour $\alpha_s \rightarrow 0$
 $Q^2 \rightarrow \infty$
"Asymptotic freedom".

6.2 The 1-loop β -fct. of QCD

- To explicitly calculate the β -fct., it is more convenient not to focus on a specific observable but to renormalize the Lagrangian. This includes renormalizing fields and the gauge-fixing parameter:
- $(A_\mu^a)^0 = z_3^{1/2} A_\mu^a$; $\xi_0 = z_3 \xi$

$$\Rightarrow (A_\mu)^0 (k^2 \eta^{\mu\nu} - k^\mu k^\nu (1 - \frac{1}{\xi_0})) (A_\nu)^0$$

$$= Z_3 A_\mu (k^2 \eta^{\mu\nu} - k^\mu k^\nu (1 - \frac{1}{Z_3 \xi})) A_\nu$$

$$= A_\mu (k^2 \eta^{\mu\nu} - k^\mu k^\nu (1 - \frac{1}{\xi})) A_\nu - \underbrace{\delta Z_3}_{\equiv 1 - Z_3} \cdot \underbrace{A_\mu (k^2 \eta^{\mu\nu} - k^\mu k^\nu) A_\nu}_{\sim F^2} \quad \text{at leading order}$$

$$\bullet \psi_0 = Z_2^{1/2} \psi$$

$$\bullet c_0 = Z_c^{1/2} c$$

$$\bullet m_0 = m - \delta m$$

$$\bullet g_0 = Z_g g$$

This is important since divergences in the gluon self-energy will be only of the "transverse form"
 $\sim k^2 \eta^{\mu\nu} - k^\mu k^\nu$.

The above implies a counterterm (among others)

$$g \bar{\psi} A \psi \cdot (\underbrace{Z_2 Z_3^{1/2} Z_g}_{\equiv Z_{1,F}} - 1) \quad (\text{renormalization of } \bar{\psi} A \psi\text{-vertex}),$$

$$\text{i.e. } Z_g = Z_{1,F} Z_2^{-1} Z_3^{-1/2}.$$

$$\text{Thus, we can write } g = \underbrace{Z_{1,F}^{-1} Z_2 Z_3^{1/2}} g_0 = Z_g^{-1} g_0$$

These are fixed by the counterterms needed to make the $\bar{\psi} A \psi$ -vertex, the fermion-selfenergy & the gluon-selfenergy finite. (We will need to calculate them!)

Note: In QED, we had $e = Z_1^{-1} Z_2 Z_3^{1/2} e_0$ and

$Z_1^{-1} = Z_2$. Thus, the photon-selfenergy was sufficient.

That's not true any more in QCD.

- As in QED, the leading order β -fact. is

$$\beta(g) = \frac{d}{d \ln \Lambda} g_0 = g \frac{d}{d \ln \Lambda} (z_g).$$

- Thus, all we need to do is to extract the coeff. of $1/\epsilon$ from $z_{1,F}$, z_2^{-1} , $z_3^{1/2}$, to add them & multiply by g .

① Gluon self-energy (z_3)

- The field redefinition $A_\mu \rightarrow A_\mu z_3^{1/2}$ induces a counterterm

$$-\frac{1}{4} (\partial_\mu A_\nu^a - \partial_\nu A_\mu^a) (\partial^\mu A^{\nu a} - \partial^\nu A^{\mu a}) \cdot \delta z_3,$$

which gives rise to a Feynman rule

$$i \Pi_{\mu\nu}^{ab}(q) = i (-\eta^{\mu\nu} q^2 + q^\mu q^\nu) \delta^{ab} \delta z_3$$

- It has to cancel the divergence in

$$i \Pi_{\mu\nu}^{(1)ab} = \text{a)} + \text{b)} + \text{c)} + \text{d)}$$

$$= i (q^2 \eta^{\mu\nu} - q^\mu q^\nu) \Pi_{(1)}(q^2) \delta^{ab}. \quad (\text{as in QED})$$

- In this notation: $\delta z_3 = \Pi_{(1)}(0) \Big|_{\frac{1}{\epsilon}\text{-term}}$

- Contribution c) is as in QED, up to the group-theoretic factor

$$T_{ij}^a T_{ji}^b = \text{tr}(T^a T^b) = \frac{1}{2} \delta^{ab}. \quad \text{For } N_f \text{ flavours, we get}$$

$$\Pi_{(1)}^{a)} = - \frac{g^2 N_f}{12\pi\epsilon} \quad (\text{cf. QFT I})$$

- It is useful to think about this a bit more generally:

The fermions could transform in any repres. (called r), generated by matrices $(T_r^a)_{ij}$. All that is required is that the T_r^a satisfy the same commut. relations as $T_F^a = T^a = \frac{\lambda^a}{2}$. We then have:

$$\text{tr}(T_r^a T_r^b) = C(r) \delta^{ab}.$$

Our factor $\frac{1}{2}$ is actually $\frac{1}{2} = C(F)$. group theoretically this is related to the "Dynkin index". One often sees $C(r) = T(r)$

- Contribution b): (using Feynman gauge, $\xi = 1$)

Recall:  $= -g f^{a_1 a_2 a_3} \{ (k_1 - k_2)_{\mu_3} \eta_{\mu_1 \mu_2} + \text{cyclic} \}$

$$\Rightarrow \begin{array}{c} \text{d, } \delta \\ \text{q, } \mu \end{array} \xrightarrow{q} \begin{array}{c} \text{q+p} \\ \text{p} \\ \text{c, } \delta \end{array} \xrightarrow{q} \begin{array}{c} \text{b, } \nu \end{array} = \frac{1}{2} \int \frac{d^d p}{(2\pi)^d} \cdot \frac{-i}{p^2} \cdot \frac{-i}{(p+q)^2} g^2 f^{acd} f^{bcd}.$$

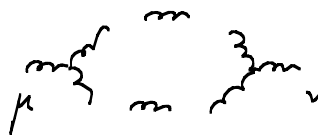
$$\cdot \left\{ \begin{array}{ccc} (q-p)_\delta \eta_{\delta\mu} & + & (p - (-q-p))_\mu \eta_{\delta\delta} & + & ((-q-p) - q)_\delta \eta_{\mu\delta} \end{array} \right\}.$$

$\begin{array}{ccc} \uparrow & \uparrow & \uparrow \\ 1 & 2 & 3 \end{array} \quad \begin{array}{ccc} \uparrow & \uparrow & \uparrow \\ 2 & 3 & 1 \end{array} \quad \begin{array}{ccc} \uparrow & \uparrow & \uparrow \\ 3 & 1 & 2 \end{array}$

$$\cdot \left\{ \begin{array}{ccc} (-q+p)^\delta \eta_{\nu}{}^\delta & + & (-p - (q+p))_\nu \eta^{\delta\delta} & + & (q+p - (-q))^\delta \eta_{\nu}{}^\delta \end{array} \right\}$$

$\begin{array}{ccc} \uparrow & \uparrow & \uparrow \\ 1 & 2 & 3 \end{array} \quad \begin{array}{ccc} \uparrow & \uparrow & \uparrow \\ 2 & 3 & 1 \end{array} \quad \begin{array}{ccc} \uparrow & \uparrow & \uparrow \\ 3 & 1 & 2 \end{array}$

- The prefactor $\frac{1}{2}$ is a symmetry factor:

We start with $\left(\frac{1}{2}\right)^2 \frac{1}{2} \cdot \frac{1}{2}$ 

$\begin{array}{ccc} \uparrow & \uparrow & \uparrow \\ \text{each prop.} & 2 \text{ props.} & \text{two vertices} \end{array}$

(from Taylor expansion of $\exp(\dots)$)

- Note that, as opposed to $\lambda\varphi^3$ -theory, the Feynman rule already incorporates the $3!$ different ways to act with

$$\frac{\delta}{\delta j_1} \frac{\delta}{\delta j_2} \frac{\delta}{\delta j_3} \quad \text{on } AAA. \quad (\text{No explicit factor } 3! \text{ appears.})$$

- The factors $(\frac{1}{2})^2$, $\frac{1}{2}$, $\frac{1}{2}$ are supposed to cancel:

- flip of two ends of each propagator
- exchange of two propagators
- exchange of two vertices.

- However, exchanging the upper & lower propagator is equivalent to rotating the diagram:



This is not a new way of connecting the lines. Hence, the corresponding factor $\frac{1}{2}$ is left over.

- One can prove that $f^{acd} f^{bcd} \sim \delta^{ab}$. In fact,
 - if f^{acd} , interpreted as a set of $(N^2-1) \times (N^2-1)$ matrices,
 - $(-if^a)^{cd}$, also generates a repres. of $SU(N)$. This repres. is known as the adjoint repres. (It corresponds to the Lie-alg. acting on itself via commutator). Thus:

$$f^{acd} f^{bcd} = (-if^a)^{cd} (-if^b)^{dc} = \text{tr}(T_A^a T_A^b) = C(A) \delta^{ab}.$$

(We used total antisymmetry of f^{abc} .)

Alternatively, we can write

$$f^{acd} f^{bcd} = f^{dca} f^{dcb} = (-if^d)^{ac} (-if^d)^{cb} = (T^d T^d)_{ab} \quad 72$$

$$= C_2(A) \delta_{ab}.$$

This last definition of " C_2 " can also be made for any other representation:

$$(T_r^a T_r^a)_{ij} = C_2(r) \delta_{ij}.$$

$C_2(r) \cdot 1$ is the "quadratic Casimir operator" of the repres. " r ".
(This name is then often also used for the constant $C_2(r)$.)

- We have just proved that $C_2(A) = C(A)$. More generally:

$$d(r) C_2(r) = d(A) C(A).$$

- One can show that, for $SU(N)$, $C_2(A) = C(A) = N$. For details see Peskin/Schroeder
- This is the group-theoretic factor of our contribution 8).
- We now rewrite the denominator using the "Feynman parameter" as in QED:

$$\frac{1}{p^2} \cdot \frac{1}{(p+q)^2} = \int_0^1 dx \frac{1}{((1-x)p^2 + x(p+q)^2)^2} = \int_0^1 dx \frac{1}{(k^2 - \Delta)^2},$$

$$\text{with } k = p + xq \quad ; \quad \Delta = -x(1-x)q^2.$$

- We can replace $\int d^d p \rightarrow \int d^d k$.
- Let us now turn to the "numerator" (the two big curly brackets above, with open indices μ, ν).
- We can rewrite this structure in terms of k, q & x .

- Furthermore, since it appears in the integral

$$\int d^d k \frac{1}{(k^2 - \Delta)^2} \dots,$$

any k -linear term gives zero after integration.

- Also: $\int d^d k f(k^2) k^\mu k^\nu = a \eta^{\mu\nu}$ by symmetry.

$$\Downarrow$$

$$\int d^d k f(k^2) k^2 = a \cdot d$$

$$\Downarrow$$

$$a = \int d^d k f(k^2) \frac{k^2}{d}.$$

- Thus, we can replace $k^\mu k^\nu \rightarrow \frac{k^2}{d} \eta^{\mu\nu}$.

- Doing all this, we find

$$\begin{aligned} & \int \frac{d^d p}{(2\pi)^d} \cdot \frac{1}{p^2 (p+q)^2} \cdot \{ \dots \} \cdot \{ \dots \} = \\ & = \int_0^1 dx \int \frac{d^d k}{(2\pi)^d} \frac{1}{(k^2 - \Delta)^2} \left\{ -\eta_{\mu\nu} k^2 \cdot 6 \cdot \left(1 - \frac{1}{d}\right) - \eta_{\mu\nu} q^2 \left[(2-x)^2 + (1+x)^2 \right] \right. \\ & \quad \left. + g_\mu g_\nu \left[(2-d)(1-2x)^2 + 2(1+x)(2-x) \right] \right\} \end{aligned}$$

(Beware! This and much of the following is just copied from Peskin/Schweder - errors are now less likely :))

We now Wick-rotate ($k_0 \rightarrow i k_0$) and use

$$\int \frac{d^d k}{(2\pi)^d} \cdot \frac{1}{(k^2 + \Delta)^n} = \frac{1}{(4\pi)^{d/2}} \cdot \frac{\Gamma(n - \frac{d}{2})}{\Gamma(n)} \left(\frac{1}{\Delta} \right)^{n - \frac{d}{2}}$$

$$\& \int \frac{d^d k}{(2\pi)^d} \frac{k^2}{(k^2 + \Delta)^n} = \frac{1}{(4\pi)^{d/2}} \frac{d}{2} \frac{\Gamma(n - \frac{d}{2} - 1)}{\Gamma(n)} \left(\frac{1}{\Delta}\right)^{n - \frac{d}{2} - 1}$$

(We already used the first integral for $n=2$ in QED.
Here we also only need $n=2$. But we also need the
second line, with the extra k^2 .)

To prove the first line, rewrite (euclidean!!)

$$d^d k = \text{Vol}(S^{d-1}) k^{d-1} dk$$

and use

$$\int_0^\infty \frac{k^{d-1} dk}{(k^2 + \Delta)^n} = \frac{1}{2} \int_0^\infty d(k^2) \frac{(k^2)^{\frac{d}{2}-1}}{(k^2 + \Delta)^n}$$

$$= \frac{1}{2} \left(\frac{1}{\Delta}\right)^{n - \frac{d}{2}} \int_0^1 dx x^{n-1-\frac{d}{2}} (1-x)^{\frac{d}{2}-1}$$

$$(k^2 + \Delta) \equiv \frac{\Delta}{x}, \quad \text{i.e.} \quad d(k^2) = -\frac{\Delta}{x^2} dx \quad \& \quad k^2 = \Delta \cdot \left(\frac{1}{x} - 1\right)$$

$$= \frac{\Delta \cdot (1-x)}{x}$$

Then use

$$\int_0^1 dx x^{a-1} (1-x)^{b-1} \equiv \mathcal{B}(a, b) = \frac{\Gamma(a) \Gamma(b)}{\Gamma(a+b)}$$

Def. of the
"Beta-fct."

(of Mathematics, not to be
confused with the β -fct. of QFT.)

important identity

(can also be proved directly by
doing the integral for integer a, b
and analytically continuing)

The case with the extra k^2 in the numerator is
also easy using $k^2 = (k^2 + \Delta) - \Delta$.

Collecting all of this, we finally get:

$$\text{cloud} = \frac{ig^2}{(4\pi)^{d/2}} C_2(A) \delta^{ab} \int_0^1 dx \frac{1}{\Delta^{2-d/2}}.$$

$$\cdot \left\{ \Gamma(1 - \frac{d}{2}) \gamma^{\mu\nu} q^2 \left[\frac{3}{2} (d-1) x(1-x) \right] + \Gamma(2 - \frac{d}{2}) \gamma^{\mu\nu} q^2 \left[\frac{1}{2} (2-x)^2 + \frac{1}{2} (1+x)^2 \right] \right. \\ \left. - \Gamma(2 - \frac{d}{2}) q^\mu q^\nu \left[(1 - \frac{d}{2}) (1-2x)^2 + (1+x)(2-x) \right] \right\}$$

• Contribution c)

No new ideas, in fact much simpler than b)

$$\text{cloud} = -g^2 C_2(A) \delta^{ab} \int \frac{d^d p}{(2\pi)^d} \cdot \frac{1}{p^2} \gamma^{\mu\nu} (d-1).$$

Note: This appears to be a quadratic divergence (like in ϕ^4 of $\lambda\phi^4$ -theory). We need to see that this divergence is cancelled by the other diagrams. On the other hand,

$$\int \frac{d^d p}{(p^2)^n} = 0 \text{ in dim. reg. (Obvious from rescaling}$$

$p \rightarrow \alpha p$. Thus, in any axiomatic definition of dim. reg. which is consistent with the usual rules of integration, this identity must hold.)

This is not a contradiction: In fact, one has at

$$d=2: \int \frac{dp^2}{p^2} = \frac{1}{\epsilon} - \frac{1}{\epsilon} = 0$$

$\uparrow \qquad \uparrow$
 UV-pole IR-pole.

While it is thus completely consistent to simply set this diagram to zero, one can try to separate the IR/UV

divergences, such that one can see that, in the end the latter cancel among the diagrams. This is done multiplying with $\frac{(q+p)^2}{(q+p)^2}$, introducing the Feynman parameter and working the algebra out as in b).

The result is

$$\text{cloud} = \frac{i g^2}{(4\pi)^{d/2}} C_2(A) \delta^{ab} \int_0^1 \frac{dx}{\Delta^{2-d/2}} \cdot \eta^{\mu\nu} q^2$$

$$\cdot \left\{ -\Gamma\left(1-\frac{d}{2}\right) \frac{1}{2} d(d-1) x(1-x) - \Gamma\left(2-\frac{d}{2}\right) (d-1) (1-x)^2 \right\}$$

↑
Here you see the pole at $d=2$. Such

a pole means a log.-divergence in $d=2$, i.e., a quadratic divergence (UV) at $d=4$.

↑
Here is the pole (log.-divergence) at $d=4$.

• The IR-divergence now hides in $\int_0^1 \frac{dx}{[-q^2 x(1-x)]^{2-d/2}},$

which is $\sim \int_0^1 \frac{dx}{x}$ at $d=2$ (uncancelled in the second

term, $\sim \Gamma(2-\frac{d}{2})$, since there is no numerator-factor x).

• Contribution d)

$$\text{loop} = (-1) \int \frac{d^d p}{(2\pi)^d} \frac{i}{p^2} \frac{i}{(p+q)^2} g^2 f^{dac} f^{cbd} (p+q)^\mu p^\nu$$

$$= \dots = \frac{i g^2}{(4\pi)^{d/2}} C_2(A) \delta^{ab} \int_0^1 \frac{dx}{\Delta^{2-d/2}}.$$

$$\cdot \left\{ -\Gamma\left(1-\frac{d}{2}\right) \eta^{\mu\nu} q^2 \frac{1}{2} x(1-x) + \Gamma\left(2-\frac{d}{2}\right) q^\mu q^\nu x(1-x) \right\}.$$

- In the combination of b), c) & d), the coeff. of $\Gamma\left(1-\frac{d}{2}\right)$ is $\sim (3d-3-d^2+d-1) = -2\left(1-\frac{d}{2}\right)(2-d)$

↑
pole $\frac{1}{1-\frac{d}{2}}$ of Γ -fct.

is cancelled! No log-div.
at $d=2 \Rightarrow$ No quadratic
divergence at $d=4$.

- Using the freedom to replace $x \leftrightarrow (1-x)$ in any term in numerator, with some algebra, everything can now be combined into:
b), c) & d)

$$\frac{i g^2}{(4\pi)^{d/2}} C_2(A) \delta^{ab} \int_0^1 dx \frac{\Gamma\left(2-\frac{d}{2}\right)}{\Delta^{2-d/2}} (\eta^{\mu\nu} q^2 - q^\mu q^\nu) \left[\left(1-\frac{d}{2}\right)(1-2x)^2 + 2 \right].$$

- This can, in fact, be fully evaluated in terms of Γ -facts only, for generic d . (see Itzykson/Zuber, where however very little calculational detail is given).

- Extracting the pole is, however easy ($\Delta^{2-d/2} \rightarrow 1$ for $d \rightarrow 4$).

[Note: To get the finite term right, one has to expand $\Delta^{-d/2} \approx 1 - \frac{d}{2} \ln(-q^2 x(1-x)) + \dots$ before doing the x -integration.]

- In combination with $\Pi_{(1)}^{(q)}$ (see above), we find:

$$\delta Z_3 = \Pi_{(1)} \Big|_{1/\epsilon\text{-part}} = - \frac{g^2 N_f}{6\pi^2 \epsilon} \cdot C(r) - \frac{g^2}{(4\pi)^2} \left(-\frac{5}{3}\right) C_2(A) \Gamma\left(2-\frac{d}{2}\right) + \dots \Big|_{1/\epsilon\text{-part}}$$

$$\sim \frac{1}{2-\frac{d}{2}} = \frac{2}{\epsilon}$$

$$= - \frac{2g^2}{(4\pi)^2 \epsilon} \left[\frac{4}{3} N_f C(r) - \frac{5}{3} C_2(A) \right]$$

↑
if we replace $C(r)$
by 1, this is
the QED vacuum-polariz.
contribution, which is all
there is in the QED β -fct.

↑ here you already see
the crucial opposite-
sign effect.
(to be enhanced
to 11/3 by
 δZ_2 & $\delta Z_{1,F}$).

Z_2 (quark self-energy)

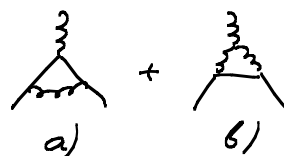
- δZ_2 has to cancel the divergence in

$$\text{quark self-energy diagram} = (ig)^2 \int \frac{d^d p}{(2\pi)^d} \gamma^\mu T^a \frac{i}{\not{p} - \not{k}} \gamma^\mu T^a \frac{-i}{\not{p}^2}.$$

- This is a good exercise in Loop-integrals & γ -matrix-algebra.
(No new ideas are required.)
- Use $(T^a T^a)_{ij} = C_2(r) \delta_{ij}$.
- The result is $\delta Z_2 = - \frac{g^2}{(4\pi)^2} \cdot \frac{2}{\epsilon} \cdot C_2(r)$.

$Z_{1,F}$ (vertex)

The divergences of the two contributing diagrams



have to be cancelled by the counterterm

$$(\mathcal{Z}_{1,F} - 1) \bar{\psi} i \not{A} \psi = \delta \mathcal{Z}_{1,F} \bar{\psi} i \not{A} \psi$$

of the lagrangian.

• Contribution a)



$$= g^3 \int \frac{d^d p}{(2\pi)^d} T^b T^a T^b \gamma^\nu \frac{1}{\not{p} + \not{k}'} \gamma^\mu \frac{1}{\not{p} + \not{k}} \gamma_\nu \cdot \frac{1}{p^2}$$

- Since there is no quadratic divergence which we would need to keep track of (unlike m_{cloud} etc.), it is sufficient to evaluate the diagram at $k = k' = 0$.

- The only new feature is in the group-theory factor:

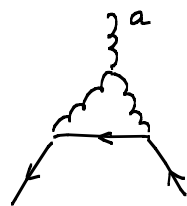
$$\begin{aligned} T^b T^a T^b &= T^b T^b T^a + T^b [T^a, T^b] \\ &= C_2(r) T^a + \underbrace{i T^b f^{abc} T^c}_{\substack{\text{if } abc \frac{1}{2} [T^b, T^c] = \frac{1}{2} \underbrace{if^{abc} if^{bcd}}_{-C(A) \delta^{ad}} T^d \\ -C(A) \delta^{ad} = -C_2(A) \delta^{ad}}} \\ &= \left[C_2(r) - \frac{1}{2} C_2(A) \right] T^a \end{aligned}$$

This is the color structure of the vertex at tree level.

- Use also $\int \frac{d^d p}{(2\pi)^d} \frac{\gamma^\nu \not{p} \gamma^\mu \not{p} \gamma_\nu}{(p^2)^3} = \int \frac{d^d p}{(2\pi)^d} \cdot \frac{\gamma^\nu \gamma^8 \gamma^\mu \gamma^8 \gamma_\nu}{(p^2)^2}$
 $\uparrow$
 $p^8 p^8 \rightarrow p^2 \frac{\gamma^{86}}{d}$

- Doing the γ -algebra and extracting the pole is straightforward.

Contribution b)



= ...

everything is very much as in a), just with a bit different explicit formulae...

- The most interesting part is again the color factor:

$$f^{abc} T^b T^c = \frac{1}{2} f^{abc}, \quad f^{bcd} T^d = \frac{i}{2} C_2(A) T^a.$$

- Combining a) & b), one finds an overall $1/\epsilon$ -divergence and the corresponding counterterm

$$\delta Z_{1,F} = - \frac{g^2}{(4\pi)^2} \cdot \frac{2}{\epsilon} (C_2(r) + C_2(A)).$$

- Finally:

$$\beta(g) = g \cdot \left(-\frac{1}{2} \delta Z_3 - \delta Z_2 + \delta Z_{1,F} \right) \Big|_{\text{coeff. of } 1/\epsilon\text{-term}}$$

$$= \frac{g^3}{(4\pi)^2} \left(\underbrace{\frac{4}{3} N_f C(r)}_{\text{from } Z_3} - \underbrace{\frac{5}{3} C_2(A)}_{Z_2} + \underbrace{2 C_2(r) - 2 C_2(r) - 2 C_2(A)}_{Z_{1,F}} \right)$$

$$= \underline{\underline{\frac{g^3}{(4\pi)^2} \left(\frac{4}{3} N_f C(r) - \frac{11}{3} C_2(A) \right)}}$$

Brief conceptual summary

renormaliz. scale M ; cutoff scale Λ

$$\beta(g_{\text{phys.}}) = \frac{d}{d \ln M} g_{\text{phys.}}(M)$$

$$g_{\text{bare}}(\Lambda) = Z_g(\Lambda/M) g_{\text{phys.}}(M)$$

$$\begin{aligned}
\frac{d}{d \ln M} g_{\text{phys.}}(M) &= \frac{d}{d \ln M} z_g^{-1}(1/M) g_{\text{bare}}(1) \\
&= - \frac{d}{d \ln 1} (z_g^{-1}(1/M)) g_{\text{bare}}(1) \\
&= z_g^{-1}(1/M) \frac{d}{d \ln 1} g_{\text{bare}}(1) \approx \frac{d}{d \ln 1} g_{\text{bare}}(1) \\
&\quad \uparrow \\
&\quad \text{1-loop} \\
&\approx \frac{d z_g(1/M)}{d \ln 1} g_{\text{phys.}}(M)
\end{aligned}$$

Now in dim. reg.: $M \rightarrow \mu$; $\frac{d}{d \ln 1} z_g = (\text{coeff. of } \frac{1}{\epsilon} \text{ in } \delta z_g)$

$$\begin{aligned}
\text{Also: } z_g &= z_{1,F} z_2^{-1} z_3^{-1/2} = (1 + \delta z_{1,F})(1 - \delta z_2)(1 - \frac{1}{2} \delta z_3) \\
&\approx 1 + (\delta z_{1,F} - \delta z_2 - \frac{1}{2} \delta z_3)
\end{aligned}$$

Hence we just need to combine the coeffs. of the $\frac{1}{\epsilon}$ -parts of these δz 's, as executed above.