

Transforming Particle Physics with AI

Tilman Plehn

Universität Heidelberg

DAMTP, March 2025



LHC: precision & uncertainties

LHC

Neural networks

Examples

Amplitudes

Generative AI

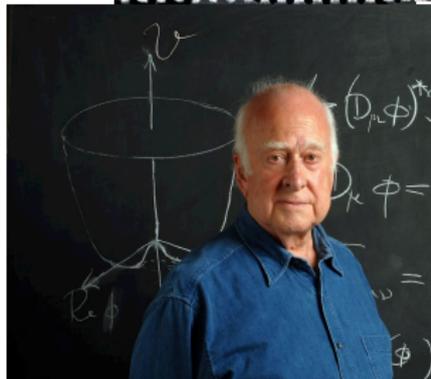
Transformation

Classic motivation

- dark matter?
- matter vs antimatter?
- origin of Higgs boson?

Strengths

- fundamental questions
- huge data set
- first-principle, precision simulations



LHC: precision & uncertainties

LHC

Neural networks

Examples

Amplitudes

Generative AI

Transformation

Classic motivation

- dark matter?
- matter vs antimatter?
- origin of Higgs boson?

Strengths

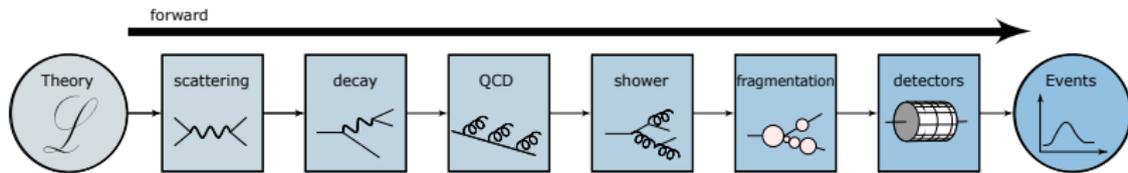
- fundamental questions
- huge data set
- first-principle, precision simulations

First-principle simulations

- start with Lagrangian
 - calculate scattering using QFT
 - simulate collisions
- LHC events in virtual worlds

Searches and measurements

- compare simulations and data
 - infer underlying theory [SM or BSM]
 - publish data to re-interpret
- Understand LHC data systematically



Brief ML-intro

Similar to fit

- approximate $f_{\theta}(x) \approx f(x)$
- no function, but very many θ
- data representation θ

Applications

- regression $x \rightarrow f_{\theta}(x)$
- classification $x \rightarrow p_{\theta}(x) \in [0, 1]$
- generation $r \sim \mathcal{N} \rightarrow p_{\theta}(r)$
- conditional generation $r \sim \mathcal{N} \rightarrow p_{\theta}(r|x)$

LHC

- x always phase space
 - symmetries, locality, etc known
- Is LHC data images or language?



Network training

Learned scalar field $f_\theta(x) \approx f(x)$

- maximize parameter probability given (f_j, σ_j)

$$\theta = \operatorname{argmax} p(\theta|x) = \operatorname{argmax} \frac{p(x|\theta) p(\theta)}{p(x)}$$

→ Gaussian likelihood loss

$$p(x|\theta) \propto \prod_j \exp\left(-\frac{|f_j - f_\theta(x_j)|^2}{2\sigma_j^2}\right)$$
$$\Rightarrow \mathcal{L} \equiv -\log p(x|\theta) = \sum_j \frac{|f_j - f_\theta(x_j)|^2}{2\sigma_j^2}$$



Network training

Learned scalar field $f_\theta(x) \approx f(x)$

- maximize parameter probability given (f_j, σ_j)

$$\theta = \operatorname{argmax} p(\theta|x) = \operatorname{argmax} \frac{p(x|\theta) p(\theta)}{p(x)}$$

→ Gaussian likelihood loss

$$p(x|\theta) \propto \prod_j \exp\left(-\frac{|f_j - f_\theta(x_j)|^2}{2\sigma_j^2}\right)$$
$$\Rightarrow \mathcal{L} \equiv -\log p(x|\theta) = \sum_j \frac{|f_j - f_\theta(x_j)|^2}{2\sigma_j^2}$$

Unknown uncertainties

- loss including normalization

$$\mathcal{L} = \frac{|f(x) - f_\theta(x)|^2}{2\sigma_\theta(x)^2} + \log \sigma_\theta(x) + \dots$$

- if needed replace with Gaussian mixture model

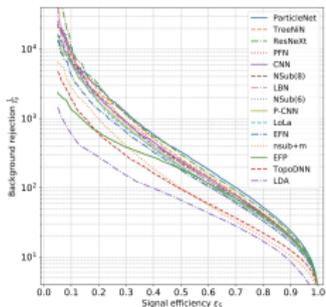
→ Learning function and (systematic) uncertainty



ML in experiment

Top tagging [classification, 2016-today]

- 'hello world' of LHC-ML
 - end of QCD-taggers
 - ever-improving [Huilin Qu]
- Driving NN-architectures



SciPost Physics

Submission

The Machine Learning Landscape of Top Taggers

G. Kasieczka (ed)¹, T. Plehn (ed)², A. Bartsch³, K. Cranmer⁴, D. Deason⁵, B. M. Dillon⁶, M. Fairbairn⁶, D. A. Farazdaghi⁷, W. Fisher⁸, C. Gao⁹, L. Goodson⁹, J. F. Kanieta^{10,11}, P. T. Komiske¹², S. Lee¹³, A. Lester¹⁴, S. Mariani¹⁵, E. M. Mitchell¹⁶, L. Monni¹⁷, B. Nachman^{18,19}, K. Nandoriya^{19,20}, J. Penzo²¹, H. Qi²², V. Rath²³, M. Rogers²⁴, D. Sill²⁵, J. M. Thompson²⁶, and S. Verra²⁷

¹ Institut für Experimentalphysik, Universität Hamburg, Germany

² Institut für Theoretische Physik, Universität Heidelberg, Germany

³ Center for Cosmology and Particle Physics and Center for Data Science, NYU, USA

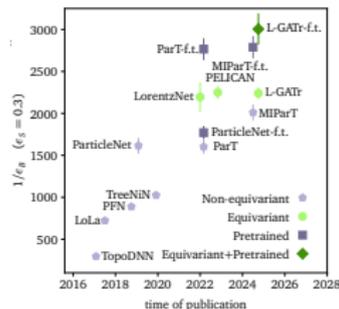
⁴ NHEK, Dept. of Physics and Astronomy, Rutgers, The State University of NJ, USA

⁵ Jozef Stefan Institute, Ljubljana, Slovenia

⁶ Theoretical Particle Physics and Cosmology, King's College London, United Kingdom

⁷ Department of Physics and Astronomy, The University of British Columbia, Canada

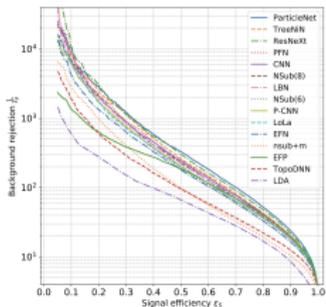
⁸ Department of Physics, University of California, Santa Barbara, USA



ML in experiment

Top tagging [classification, 2016-today]

- 'hello world' of LHC-ML
- end of QCD-taggers
- ever-improving [Huilin Qu]
- **Driving NN-architectures**



Particle flow [2020-today]

- basis of jet analyses
- combining detectors with different resolution
- **Optimality the key**

Towards a Computer Vision Particle Flow *

Francesco Aronzo Di Bello^{1,2}, Sanmay Ganguly^{3,4}, Eilan Gross⁵, Marumi Kado^{6,7}, Michael Pitt⁸, Lorenzo Sutti⁹, Jonathan Shlomi¹⁰

¹Wizman Institute of Science, Rehovot 76100, Israel
²CERN, CH-1211, Geneva 23, Switzerland
³Università di Roma Sapienza, Piazza Aldo Moro, 2, 00185 Roma, Italy e INFN, Italy
⁴Université Paris-Saclay, CNRS/IN2P3, DCLab, 91405, Orsay, France

Progress towards an improved particle flow algorithm at CMS with machine learning

Farook Mukhtar¹, Josep Pata², Javier Duarte³, Eric Walt⁴, Maricela Peraldo⁵ and Jose-Rocio Vilmar⁶

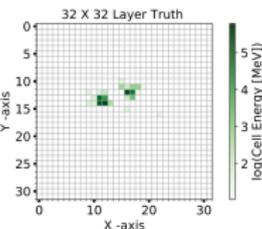
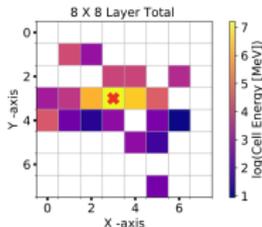
(on behalf of the CMS Collaboration)
¹University of California San Diego, La Jolla, CA 92036, USA
²NSF, Brooks pd 01, 50512 Tallahassee, Florida
³University of California San Diego, 9500 Gilman Drive, San Diego, CA 92161, USA
⁴University of California, Berkeley, CA 94720, USA
⁵University of California, Berkeley, CA 94720, USA
⁶University of California, Berkeley, CA 94720, USA

SciPost Physics Submission

The Machine Learning Landscape of Top Taggers

G. Kasieczka (ed)¹, T. Plehn (ed)², A. Bartsch³, E. Cruzado⁴, D. Debanic⁵, B. M. Dillon⁶, M. Fairhead⁷, D. A. Ferguson⁸, M. Feldner⁹, C. Gao¹⁰, L. González¹¹, J. P. Kauer¹², P. T. Komiske¹³, S. Lester¹⁴, A. Lister¹⁵, S. Malhotra¹⁶, E. M. Metodiev¹⁷, L. Moore¹⁸, B. Nachman^{19,20}, K. Nataraj^{21,22}, J. Puelari²³, H. Qu²⁴, Y. Rath²⁵, M. Ringer²⁶, D. Shih²⁷, J. M. Thompson²⁸, and S. Varrault²⁹

- 1 Institut für Experimentalphysik, Universität Hamburg, Germany
- 2 Institut für Theoretische Physik, Universität Heidelberg, Germany
- 3 Center for Cosmology and Particle Physics and Center for Data Science, NYU, USA
- 4 NHBC, Dept. of Physics and Astronomy, Rutgers, The State University of NJ, USA
- 5 Saclay Status Institute, Ljubljana, Slovenia
- 6 Theoretical Particle Physics and Cosmology, King's College London, United Kingdom
- 7 Department of Physics and Astronomy, The University of British Columbia, Canada
- 8 Department of Physics, University of California, Santa Barbara, USA
- 9 Faculty of Mathematics and Physics, University of Ljubljana, Ljubljana, Slovenia
- 10 Center for Theoretical Physics, MIT, Cambridge, USA
- 11 CP3, Université Catholique de Louvain, Louvain-la-Neuve, Belgium
- 12 Physics Division, Lawrence Berkeley National Laboratory, Berkeley, USA
- 13 Simons Inst. for the Theory of Computing, University of California, Berkeley, USA
- 14 National Institute for Subatomic Physics (NIKHEF), Amsterdam, Netherlands
- 15 LPTEP, CNRS & Sorbonne Université, Paris, France
- 16 III. Physikalisches Institut A, RWTH Aachen University, Germany

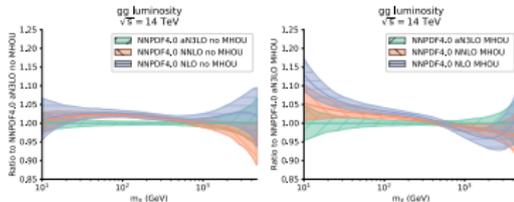


ML in phenomenology

Parton densities [NNPDF, 2002-today]

- LHC-ML classic
- pdfs with uncertainties and without bias

→ Driving precision



The Path to N³LO Parton Distributions

The NNPDF Collaboration:

Richard D. Ball¹, Andrea Barzanti², Alessandro Ciaffaglia^{2,3}, Stefano Carrazza², Juan Cruz-Martinez², Luigi Del Debbio⁴, Stefano Forte⁵, Tommaso Ganti^{6,7}, Felix Hehner^{8,9,10}, Zahed Kanaan⁶, Niccolò Lauretti⁷, Giacomo Maga^{4,5}, Emanuele H. Nieves¹¹, Tiziana R. Saborido-Saiz^{4,5}, Juan Rojo^{4,5}, Christopher Schmidt¹², Roy Stegmann¹³, and Maria Ubiali¹⁴

¹The Hugh Christie for Theoretical Physics, University of Edinburgh, JCMB, KB, Higgsfield BA, Edinburgh ED9 1JZ, Scotland

²FY Lab, Dipartimento di Fisica, Università di Milano and INFN, Sezione di Milano, Via Celoria 16, I-20132 Milano, Italy

³CERN, Theoretical Physics Department, CH-1211 Geneva 23, Switzerland

⁴Department of Physics and Astronomy, Uppsala University, SE-0801 SV Astroteknik

⁵HEP&T Theory Group, Science Park 105, 1099 30 Amsterdam, The Netherlands

⁶University of Jyväskylä, Department of Physics, P.O. Box 25, FI-00014 University of Jyväskylä, Finland

⁷Helsinki Institute of Physics, P.O. Box 64, FI-00014 University of Helsinki, Finland

⁸DMTPP, University of Cambridge, Wilberforce Road, Cambridge, CB3 0WA, United Kingdom

⁹Dipartimento di Fisica, Università degli Studi di Torino and INFN, Sezione di Torino, Via Pevero 13/A, I-10133 Torino, Italy

¹⁰Universität Würzburg, Institut für Theoretische Physik und Astrophysik, 97074 Würzburg, Germany

¹¹Università di Bari, Dipartimento di Fisica, I-70126 Bari, Italy

¹²University of Cambridge, Cavendish Laboratory, Cambridge, CB3 0HE, United Kingdom

¹³University of Cambridge, Cavendish Laboratory, Cambridge, CB3 0HE, United Kingdom

¹⁴University of Cambridge, Cavendish Laboratory, Cambridge, CB3 0HE, United Kingdom

This paper is dedicated to the memory of Stefano Citani, Grand Master of QCD, great scientist and human being

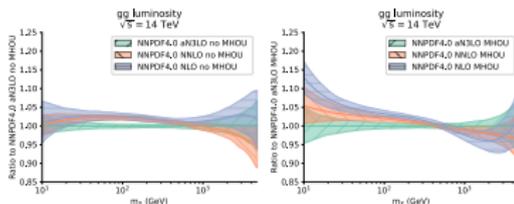


ML in phenomenology

Parton densities [NNPDF, 2002-today]

- LHC-ML classic
- pdfs with uncertainties and without bias

→ Driving precision



The Path to N³LO Parton Distributions

The NNPDF Collaboration:

Richard D. Ball¹, Andrea Bottino², Alessandro Castellani³, Stefano Carrazza², Juan Cruz-Martinez⁴, Luigi Del Debbio⁵, Stefano Di Bartolomeo⁶, Felix Dulat^{7,8}, Zoltan Kunszt⁹, Niccolò Lauret², Giacomo Magre¹⁰, Emanuele M. Remmen¹¹, Tarjona R. S. Baranowski^{12,13}, Juan Rojo¹⁴, Christopher Schwan¹⁵, Roy Stogner¹⁶, and Maria Ubald¹⁷

¹The High Centre for Theoretical Physics, University of Edinburgh, JCRC, KB, Higgsfield BA, Edinburgh EH9 1JZ, Scotland

²FY Lab, Dipartimento di Fisica, Università di Milano and INFN, Sezione di Milano, Via Celoria 16, I-20132 Milano, Italy

³CERN, Theoretical Physics Department, CH-1211 Geneva 23, Switzerland

⁴Department of Physics and Astronomy, York University, 4700 Keele Street, Toronto, Ontario M3J 1P3, Canada

⁵RIKEN Theory Group, Science Park 191, 351-0199 Saitama, The Netherlands

⁶University of Jyväskylä, Department of Physics, P.O. Box 25, FI-00014 University of Jyväskylä, Finland

⁷Helsinki Institute of Physics, P.O. Box 64, FI-00014 University of Helsinki, Finland

⁸DAMTP, University of Cambridge, Wilberforce Road, Cambridge, CB3 0WA, United Kingdom

⁹Dipartimento di Fisica, Università degli Studi di Torino and INFN, Sezione di Torino, Via Pietro Giuria 1, I-10125 Torino, Italy

¹⁰Universität Würzburg, Institut für Theoretische Physik und Astrophysik, 97074 Würzburg, Germany

¹¹This paper is dedicated to the memory of Stefano Catani, Grand Master of QCD, great scientist and human being

¹²This paper is dedicated to the memory of Stefano Catani, Grand Master of QCD, great scientist and human being

¹³This paper is dedicated to the memory of Stefano Catani, Grand Master of QCD, great scientist and human being

¹⁴This paper is dedicated to the memory of Stefano Catani, Grand Master of QCD, great scientist and human being

¹⁵This paper is dedicated to the memory of Stefano Catani, Grand Master of QCD, great scientist and human being

¹⁶This paper is dedicated to the memory of Stefano Catani, Grand Master of QCD, great scientist and human being

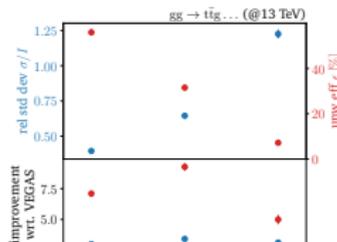
¹⁷This paper is dedicated to the memory of Stefano Catani, Grand Master of QCD, great scientist and human being

Ultra-fast simulations [Sherpa, MadNIS, MLHad]

- event generation modular
- better ML-modules

→ MadNIS → MadGraph7

Triple-W	$u\bar{d} \rightarrow W^+W^+W^-$		
VBS	$uc \rightarrow W^+W^+ds$		
W+jets	$g\bar{g} \rightarrow W^+d\bar{u}$	$g\bar{g} \rightarrow W^+d\bar{g}$	$g\bar{g} \rightarrow W^+d\bar{g}\bar{g}$
$t\bar{t}$ +jets	$g\bar{g} \rightarrow t\bar{t}+g$	$g\bar{g} \rightarrow t\bar{t}+g\bar{g}$	$g\bar{g} \rightarrow t\bar{t}+g\bar{g}\bar{g}$



SciPost Physics

Submission

The MadNIS Reloaded

Theo Heinzel¹, Nathan Huetich¹, Fabian Maltoni^{2,3}, Olivier Mattelaer¹, Tilman Plehn¹, and Ramon Winterhiller²

¹Institut für Theoretische Physik, Universität Heidelberg, Germany
²CPH, Université catholique de Louvain, Louvain-la-Neuve, Belgium
³Dipartimento di Fisica e Astronomia, Università di Bologna, Italy

December 17, 2024

Abstract

In pursuit of precise and fast theory predictions for the LHC, we present an implementation of the MadNIS method in the MadGraph5 event generator. A series of improvements in MadNIS further enhance its efficiency and speed. We validate this implementation for realistic partonic processes and find significant gains from using modern machine learning in event generators.

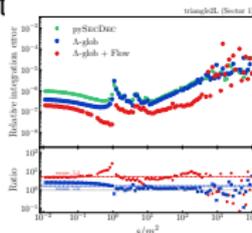
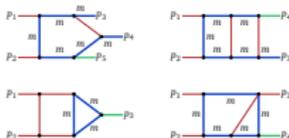


ML in theory

Optimizing integration paths [invertible networks]

- compute Feynman integrals
- learn optimal integration pat^t

→ To be implemented...



Targeting multi-loop integrals with neural networks

Emmanuel Winterhager^{1,2,3}, Vinay Meghria⁴, Emilio Villa⁵, Stephen F. Jones⁶, Matthias Kerner^{6,8}, Anja Rottler^{2,3}, Gudrun Heinrich^{6,8} and Tilman Plehn^{1,2}

¹ Institut für Theoretische Physik, Universität Heidelberg, Germany

² HEKA - Heidelberg Karlsruhe Strategic Partnership, Heidelberg University, Karlsruhe Institute of Technology (KIT), Germany

³ Centre for Cosmology, Particle Physics and Phenomenology (CP3), Université catholique de Louvain, Belgium

⁴ Institut für Theoretische Physik, Karlsruher Institut für Technologie, Germany

⁵ Institute for Particle Physics Phenomenology, Durham University, UK

⁶ Institute für Astroteilchenphysik, Karlsruher Institut für Technologie, Germany

Abstract

Numerical evaluations of Feynman integrals often proceed via a deformation of the integration contour into the complex plane. While valid contours are easy to construct, the numerical precision for a multi-loop integral can depend critically on the chosen contour. We present methods to optimize this contour using a combination of optimized, global complex shifts and a normalizing flow. They can lead to a significant gain in precision.

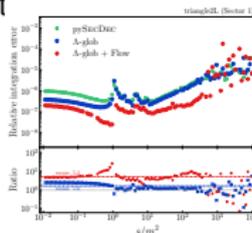
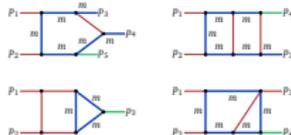


ML in theory

Optimizing integration paths [invertible networks]

- compute Feynman integrals
- learn optimal integration paths

→ To be implemented...



String landscape [reinforcement learning]

- searching for viable vacua
- high dimensions, unknown global structure

→ Life on Mars?

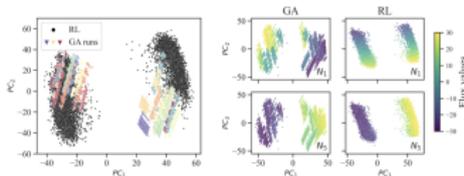


Figure 1: Left: Cluster structure in dimensionally reduced flux samples for RL and 25 GA runs (PCA on all samples of GA and RL). The colors indicate individual GA runs. Right: Dependence on flux (input) values (N_1 and N_2 respectively) in relation to principal components for a PCA fit of the individual output of GA and RL.

Targeting multi-loop integrals with neural networks

Emanuele Winterhager^{1,2,3}, Vinay Megaraj⁴, Emilio Villa⁵, Stephen F. Jones⁶, Matthias Kerner⁶, Anja Rottler^{2,3}, Gudrun Heinrich^{6,4} and Tilman Plehn^{1,2}

- 1 Institut für Theoretische Physik, Universität Heidelberg, Germany
- 2 HEKA - Heidelberg Karlsruhe Strategic Partnership, Heidelberg University, Karlsruhe Institute of Technology (KIT), Germany
- 3 Centre for Cosmology, Particle Physics and Phenomenology (CP3), Université catholique de Louvain, Belgium
- 4 Institut für Theoretische Physik, Karlsruher Institut für Technologie, Germany
- 5 Institute for Particle Physics Phenomenology, Durham University, UK
- 6 Institute für Astronomiephysik, Karlsruher Institut für Technologie, Germany

Abstract

Numerical evaluations of Feynman integrals often proceed via a deformation of the integration contour into the complex plane. While valid contours are easy to construct, the numerical precision for a multi-loop integral can depend critically on the chosen contour. We present methods to optimize this contour using a combination of optimized, global complex shifts and a normalizing flow. They can lead to a significant gain in precision.

Probing the Structure of String Theory Vacua with Genetic Algorithms and Reinforcement Learning

Alex Cole
University of Amsterdam
a.w.cole@uva.nl

Sven Krippendorff
Arnold Sommerfeld Center for Theoretical Physics
LMU Munich
sven.krippendorff@physik.uni-muenchen.de

Andreas Schachner
Center for Mathematical Sciences
University of Cambridge
as207@cam.ac.uk

Gary Shiu
University of Wisconsin-Madison
shiu@physics.wisc.edu

Abstract

Identifying string theory vacua with desired physical properties at low energies requires searching through high-dimensional solution spaces – collectively referred to as the string landscape. We highlight that this search problem is amenable to reinforcement learning and genetic algorithms. In the context of flux vacua, we are able to reveal novel features (suggesting previously overlooked symmetries) in the string theory solutions required for properties such as the string cosmology. In order to identify these features robustly, we combine results from both search methods, which we argue is imperative for reducing sampling bias.



Statistics and systematics

Statistical approach [Yarin Gal (2016)]

- expectation value with internal representation θ

$$\langle A \rangle = \int dA A p(A|x) = \int dA A \int d\theta p(A|\theta) p(\theta|A_{\text{train}})$$

- training a generalization

$$\int d\theta p(A|\theta) p(\theta|A_{\text{train}}) \approx \int d\theta p(A|\theta) q(\theta)$$



Statistical approach [Yarin Gal (2016)]

- expectation value with internal representation θ

$$\langle A \rangle = \int dA A p(A|x) = \int dA A \int d\theta p(A|\theta) p(\theta|A_{\text{train}})$$

- training a generalization

$$\int d\theta p(A|\theta) p(\theta|A_{\text{train}}) \approx \int d\theta p(A|\theta) q(\theta)$$

- similarity from minimal KL-divergence

$$\begin{aligned} D_{\text{KL}}[q(\theta), p(\theta|A_{\text{train}})] &\equiv \int d\theta q(\theta) \log \frac{q(\theta)}{p(\theta|A_{\text{train}})} \\ &= \int d\theta q(\theta) \log \frac{q(\theta)p(A_{\text{train}})}{p(A_{\text{train}}|\theta)p(\theta)} \\ &= - \int d\theta q(\theta) \log p(A_{\text{train}}|\theta) + \int d\theta q(\theta) \log \frac{q(\theta)}{p(\theta)} + \dots \end{aligned}$$

- regularized likelihood loss

$$\mathcal{L} = - \int d\theta q(\theta) \log p(A_{\text{train}}|\theta) + D_{\text{KL}}[q(\theta), p(\theta)]$$



Statistical approach [Yarin Gal (2016)]

- expectation value with internal representation θ

$$\langle A \rangle = \int dA A p(A|x) = \int dA A \int d\theta p(A|\theta) p(\theta|A_{\text{train}})$$

- training a generalization

$$\int d\theta p(A|\theta) p(\theta|A_{\text{train}}) \approx \int d\theta p(A|\theta) q(\theta)$$

- similarity from minimal KL-divergence

$$\begin{aligned} D_{\text{KL}}[q(\theta), p(\theta|A_{\text{train}})] &\equiv \int d\theta q(\theta) \log \frac{q(\theta)}{p(\theta|A_{\text{train}})} \\ &= \int d\theta q(\theta) \log \frac{q(\theta)p(A_{\text{train}})}{p(A_{\text{train}}|\theta)p(\theta)} \\ &= - \int d\theta q(\theta) \log p(A_{\text{train}}|\theta) + \int d\theta q(\theta) \log \frac{q(\theta)}{p(\theta)} + \dots \end{aligned}$$

- regularized likelihood loss

$$\mathcal{L} = - \int d\theta q(\theta) \log p(A_{\text{train}}|\theta) + D_{\text{KL}}[q(\theta), p(\theta)]$$

→ Variance

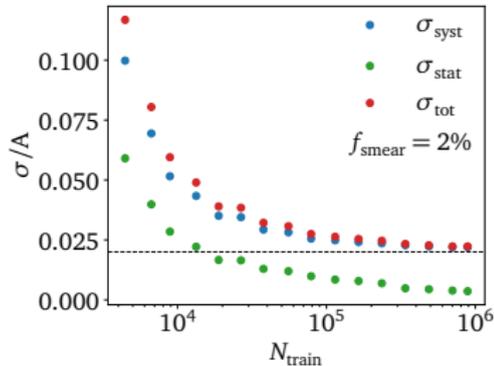
$$\sigma^2 = \int dA \int d\theta (A - \langle A \rangle)^2 p(A|\theta) q(\theta) \equiv \sigma_{\text{syst}}^2 + \sigma_{\text{stat}}^2$$



Amplitudes with calibrated uncertainties

Loop amplitude $gg \rightarrow \gamma\gamma g(g)$ over phase space

- systematics: **artificial noise**
- statistics plateau

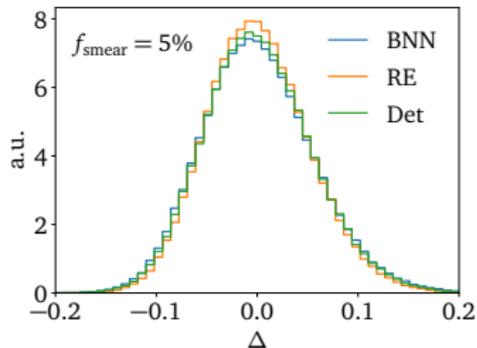
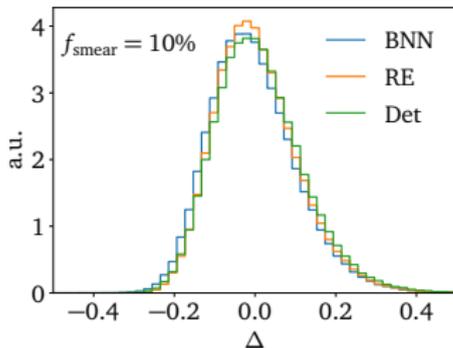


Amplitudes with calibrated uncertainties

Loop amplitude $gg \rightarrow \gamma\gamma g(g)$ over phase space

- systematics: **artificial noise**
- statistics plateau
- accuracy over phase space

$$\Delta(x) = \frac{A_{\text{NN}}(x) - A_{\text{true}}(x)}{A_{\text{true}}(x)}$$



Amplitudes with calibrated uncertainties

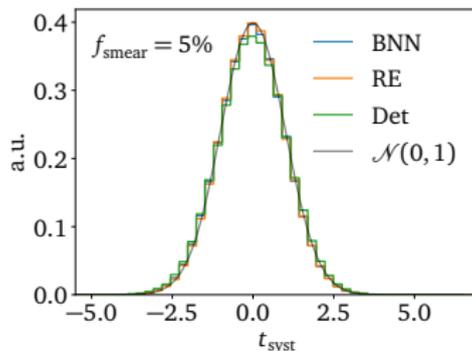
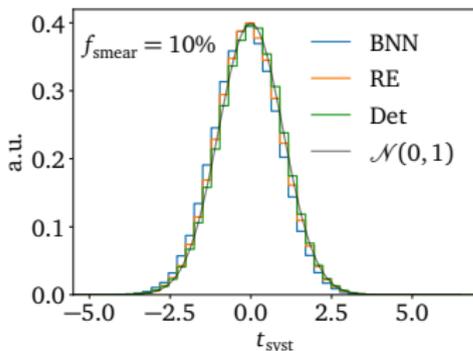
Loop amplitude $gg \rightarrow \gamma\gamma g(g)$ over phase space

- systematics: **artificial noise**
- statistics plateau
- accuracy over phase space

$$\Delta(x) = \frac{A_{\text{NN}}(x) - A_{\text{true}}(x)}{A_{\text{true}}(x)}$$

- pull over phase space

$$t_{\text{syst}}(x) = \frac{A_{\text{NN}}(x) - A_{\text{true}}(x)}{\sigma_{\text{syst}}(x)}$$



Amplitudes with calibrated uncertainties

Loop amplitude $gg \rightarrow \gamma\gamma g(g)$ over phase space

- systematics: **artificial noise**
- statistics plateau
- accuracy over phase space
- pull over phase space

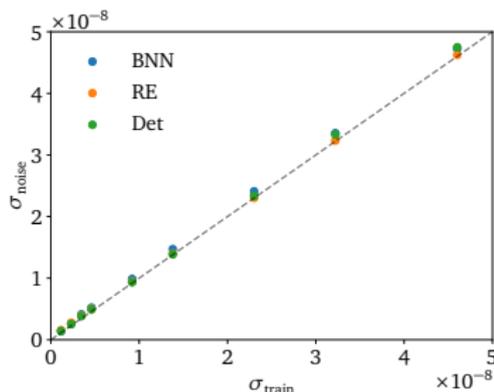
$$\Delta(x) = \frac{A_{\text{NN}}(x) - A_{\text{true}}(x)}{A_{\text{true}}(x)}$$

$$t_{\text{syst}}(x) = \frac{A_{\text{NN}}(x) - A_{\text{true}}(x)}{\sigma_{\text{syst}}(x)}$$

Towards zero noise

- scaling

$$\sigma_{\text{syst}}^2 - \sigma_{\text{syst},0}^2 \approx \sigma_{\text{train}}^2$$



Amplitudes with calibrated uncertainties

Loop amplitude $gg \rightarrow \gamma\gamma g(g)$ over phase space

- systematics: **artificial noise**
- statistics plateau
- accuracy over phase space
- pull over phase space

$$\Delta(x) = \frac{A_{\text{NN}}(x) - A_{\text{true}}(x)}{A_{\text{true}}(x)}$$

$$t_{\text{syst}}(x) = \frac{A_{\text{NN}}(x) - A_{\text{true}}(x)}{\sigma_{\text{syst}}(x)}$$

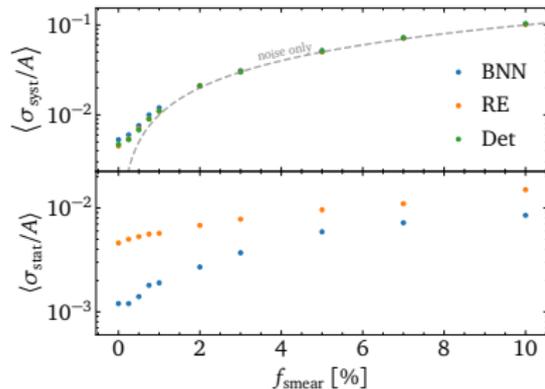
Towards zero noise

- scaling

$$\sigma_{\text{syst}}^2 - \sigma_{\text{syst},0}^2 \approx \sigma_{\text{train}}^2$$

- plateau $\langle \sigma_{\text{syst}}/A \rangle \sim 0.4\%$

→ **Limiting factor??**



Amplitudes with calibrated uncertainties

Loop amplitude $gg \rightarrow \gamma\gamma g(g)$ over phase space

- systematics: **artificial noise**
- statistics plateau
- accuracy over phase space

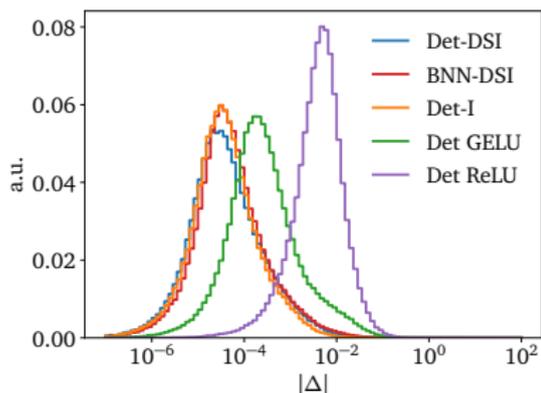
- pull over phase space

$$\Delta(x) = \frac{A_{\text{NN}}(x) - A_{\text{true}}(x)}{A_{\text{true}}(x)}$$

$$t_{\text{syst}}(x) = \frac{A_{\text{NN}}(x) - A_{\text{true}}(x)}{\sigma_{\text{syst}}(x)}$$

Data pre-processing

- amplitude from invariants
- learn Minkowski metric
- Deep-sets-invariant network
L-GATr transformer



Amplitudes with calibrated uncertainties

Loop amplitude $gg \rightarrow \gamma\gamma g(g)$ over phase space

- systematics: **artificial noise**
- statistics plateau
- accuracy over phase space

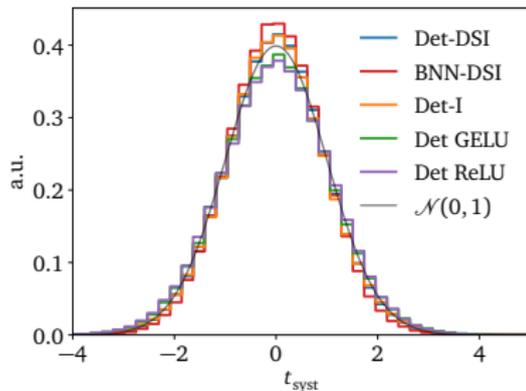
- pull over phase space

$$\Delta(x) = \frac{A_{\text{NN}}(x) - A_{\text{true}}(x)}{A_{\text{true}}(x)}$$

$$t_{\text{syst}}(x) = \frac{A_{\text{NN}}(x) - A_{\text{true}}(x)}{\sigma_{\text{syst}}(x)}$$

Data pre-processing

- amplitude from invariants
 - learn Minkowski metric
 - Deep-sets-invariant network
L-GATr transformer
- **Calibrated systematics**



ATLAS calibration

Energy calibration with uncertainties [ATLAS + Heidelberg]

- interpretable calorimeter phase space x
- learned calibration function

$$\mathcal{R}_{\text{NN}}(x) \pm \Delta \mathcal{R}_{\text{NN}}(x) \approx \frac{E^{\text{obs}}(x)}{E^{\text{dep}}(x)}$$

- **systematics:** noise in data
network expressivity
data representation ...



ATLAS calibration

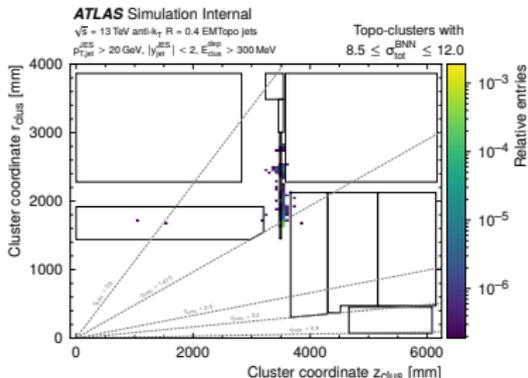
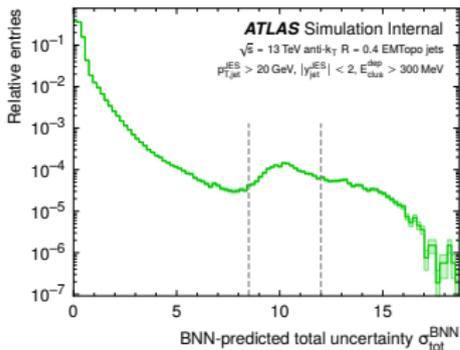
Energy calibration with uncertainties [ATLAS + Heidelberg]

- interpretable calorimeter phase space x
- learned calibration function

$$\mathcal{R}_{NN}(x) \pm \Delta \mathcal{R}_{NN}(x) \approx \frac{E^{obs}(x)}{E^{dep}(x)}$$

- **systematics:** noise in data
network expressivity
data representation ...

→ Understand (simulated) detector

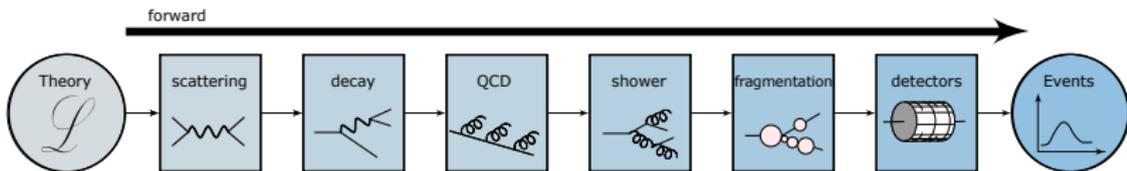
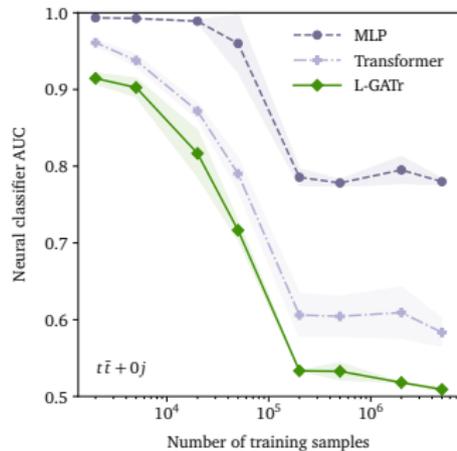


Generative AI

Simulations, MadNIS, calorimeters,...

- learn phase space density
fast sampling Gaussian \rightarrow phase space
Bayesian generative network \rightarrow uncertainties
- Variational Autoencoder
 \rightarrow low-dimensional physics
- Generative Adversarial Network
 \rightarrow generator trained by classifier
- Normalizing Flow/Diffusion
 \rightarrow (bijective) mapping
- JetGPT, ViT
 \rightarrow non-local structures
- Equivariant L-GATr
 \rightarrow Lorentz symmetry for efficiency

\rightarrow Equivariant transformer CFM...



Controlling generative AI

Compare generated with training data

- generation: unsupervised density
- classify training vs generated events $D(x)$
learned density ratio [Neyman-Pearson]

$$w(x_i) = \frac{D(x_i)}{1 - D(x_i)} = \frac{p_{\text{data}}(x_i)}{p_{\text{model}}(x_i)}$$

→ Test ratio over phase space



Controlling generative AI

Compare generated with training data

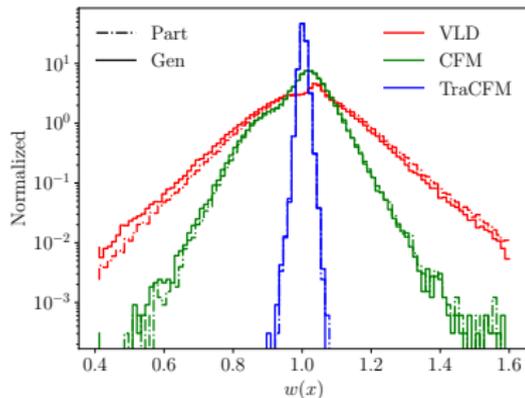
- generation: unsupervised density
- classify training vs generated events $D(x)$
learned density ratio [Neyman-Pearson]

$$w(x_i) = \frac{D(x_i)}{1 - D(x_i)} = \frac{\rho_{\text{data}}(x_i)}{\rho_{\text{model}}(x_i)}$$

→ Test ratio over phase space

Testing NN-generators

- accuracy from width of weight distribution
 - tails indicating failure mode
- Systematic performance test



Transforming LHC physics

Number of searches

- optimal inference: signal and background simulations
- CPU-limitation for many signals?

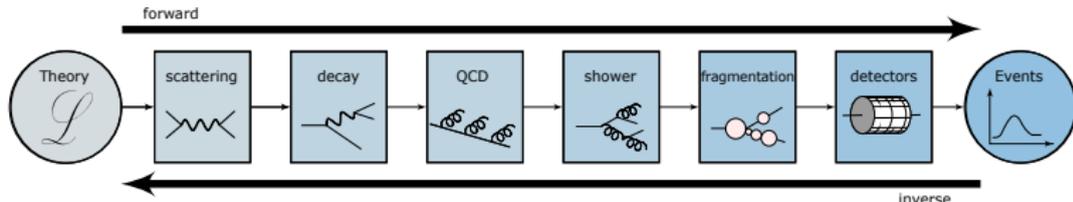
Optimal analyses

- theory limiting many analyses
- include predictions not in event generators

Public LHC data

- common lore:
LHC data too complicated for amateurs
- in truth:
hard scattering and decay simulations public
BSM physics not in hadronization and detector

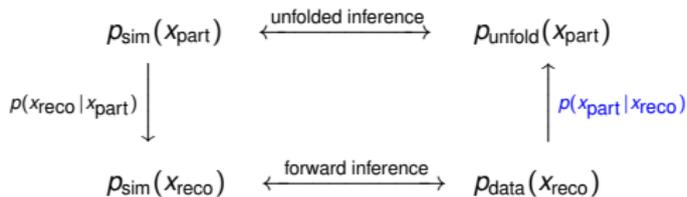
→ **Unfold to suitable level**



ML-Unfolding

View as generative inference [Köthe etal, Macke etal]

- four phase space distributions



- learn conditional probabilities from $(X_{\text{part}}, X_{\text{reco}})$ [forward-inverse symmetric]

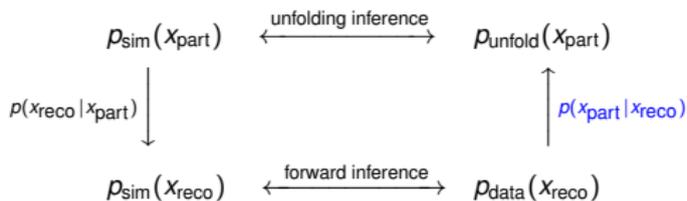
→ Unbinned and high-dimensional unfolding



ML-Unfolding

View as generative inference [Köthe et al, Macke et al]

- four phase space distributions

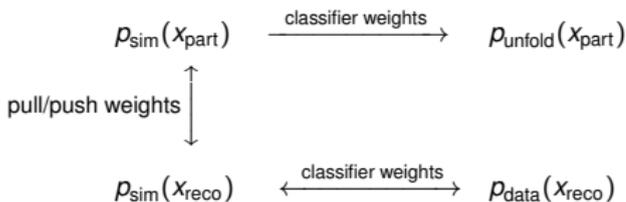


- learn conditional probabilities from $(x_{\text{part}}, x_{\text{reco}})$ [forward-inverse symmetric]

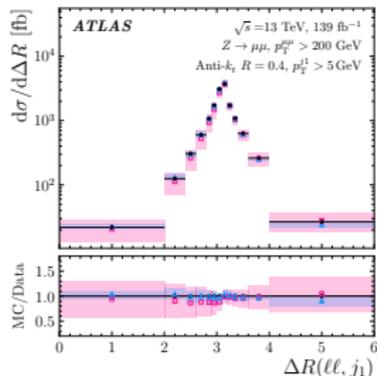
→ Unbinned and high-dimensional unfolding

OmniFold

- learn $\rho_{\text{sim}}(x_{\text{reco}}) \leftrightarrow \rho_{\text{data}}(x_{\text{reco}})$
- reweight $\rho_{\text{sim}}(x_{\text{part}}) \rightarrow \rho_{\text{unfold}}(x_{\text{part}})$



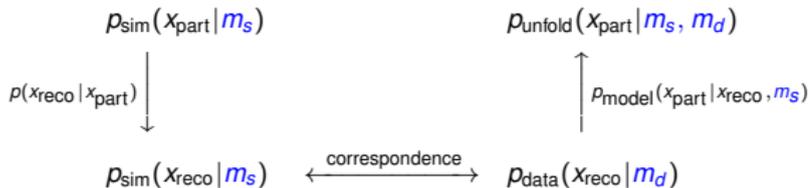
→ Z+jets in 24D [ATLAS]



Unfolding top decays

Top mass as high school project [Heidelberg + CMS]

- first measure m_t in unfolded data
then unfold full kinematics
- simulation m_s vs data m_d [too bad to reweight]



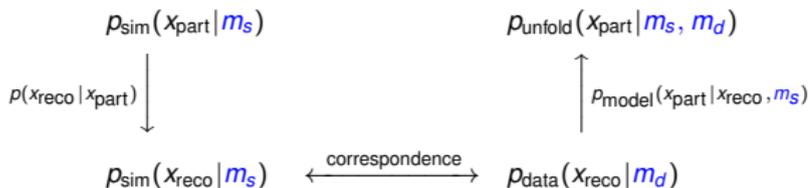
→ train on m_s -range
include batch-wise $M_{jjj} \in x_{\text{reco}}$



Unfolding top decays

Top mass as high school project [Heidelberg + CMS]

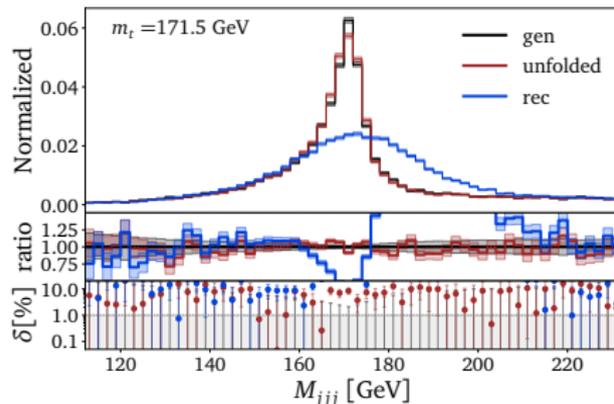
- first measure m_t in unfolded data
then unfold full kinematics
- simulation m_s vs data m_d [too bad to reweight]



- train on m_s -range
include batch-wise $M_{jjj} \in x_{\text{reco}}$

Preliminary unfolding results [TraCFM]

- 4D for m_t -measurement
uncloning m_W -calibration
- 12D published data
- CMS data next



ML for LHC Theory

Developing ML for the best science

- 1 just another numerical tool for a numerical field
- 2 completely transformative new language
 - driven by money from data science and medical research
 - physics should be leading scientific AI
 - 10000 Einsteins...
 - ...improving established tools
 - ...developing new tools for established tasks
 - ...transforming through new ideas

→ Complexity your friend

Modern Machine Learning for LHC Physicists

Tilman Plehn^a, Anja Butter^{a,b}, Barry Dillon^a,
Theo Heime^a, Claudius Krause^c, and Ramon Winterhalder^d

^a Institut für Theoretische Physik, Universität Heidelberg, Germany

^b LPNHE, Sorbonne Université, Université Paris Cité, CNRS/IN2P3, Paris, France

^c HEPHY, Austrian Academy of Sciences, Vienna, Austria

^d CP3, Université catholique de Louvain, Louvain-la-Neuve, Belgium

March 19, 2024

Abstract

Modern machine learning is transforming particle physics fast, bullying its way into our numerical tool box. For young researchers it is crucial to stay on top of this development, which means applying cutting-edge methods and tools to the full range of LHC physics problems. These lecture notes lead students with basic knowledge of particle physics and significant enthusiasm for machine learning to relevant applications. They start with an LHC-specific motivation and a non-standard introduction to neural networks and then cover classification, unsupervised classification, generative networks, and inverse problems. Two themes defining much of the discussion are well-defined loss functions and uncertainty-aware networks. As part of the applications, the notes include some aspects of theoretical LHC physics. All examples are chosen from particle physics publications of the last few years.¹



Generative AI with uncertainties

Bayesian generative networks

- encoding phase space probabilities
 - events with error bars on weights
 - learned density & uncertainty reflecting network learning
- **Generative networks like fitted densities**



Generative AI with uncertainties

Bayesian generative networks

- encoding phase space probabilities
- events with error bars on weights
- learned density & uncertainty reflecting network learning

→ Generative networks like fitted densities

Z+jets events

- per-cent accuracy on density
- statistical uncertainty from BNN
- systematics in training data

$$w = 1 + a \left(\frac{p_{T,j_1} - 15 \text{ GeV}}{100 \text{ GeV}} \right)^2$$

- training with condition a
 - sampling including a
- Precision and uncertainty control

