IVIL IOI LHC

Uncertainties

Ensembles

Congration

Neural Networks with Calibrated Learned Uncertainties

Tilman Plehn

Universität Heidelberg

Imperial College, February 2025



ML for LHC

Shortest ML-intro ever

Fit-like approximation [2211.01421]

- · approximate $f_{\theta}(x) \approx f(x)$
- · no parametrization, but many θ
- · new representation/latent space θ

Construction and contol

- define loss function
- · minimize loss to find best θ
- · compare $x \to f_{\theta}(x)$ for training vs test data

LHC applications

- · regression $x \to f_{\theta}(x)$
- · classification $x \to f_{\theta}(x) \in [0, 1]$
- · generation $r \sim \mathcal{N} \rightarrow f_{\theta}(r)$
- · conditional generation $r \sim \mathcal{N} \rightarrow f_{\theta}(r|x)$
-
- → Transforming numerical science

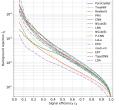


ML for LHC

ML in experiment

Top tagging [classification, 2016-today]

- · 'hello world' of LHC-ML
- end of QCD-taggers
- · ever-improving [Huilin Qu]
- → Driving NN-architectures



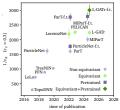
The Machine Learning Landscape of Top Taggers

SciPost Physics

G. Kasiocaka (ed)¹, T. Fishin (ed)², A. Bortier², K. Crazzner², D. Debrach³, B. M. Dilson⁵, M. Barbalm³, D. A. Brozaghy³, W. Federkor³, C. Gay², L. Gouskos³, J. F. Kamenik^{3,5}, P. T. Komiske¹, S. Leise¹, A. Lister³, S. Macsimol³, E. M. Meterdine^{1,8}, L. Mosse^{1,1}, B. Nochman, ^{12,13}, K. Nordstrein^{1,13}, J. Poschen³, B. Qel³, Y. Rath³, M. Reger^{1,9}, D. Shih³ J. M. Thompson², and S. Varma⁵

1 Institut für Experimentalphysik, Universität Hamburg, Germany 8 Center for Cosmology and Particle Physics and Center for Data Science, NYU, USA 4 NHECT, Dept. of Physics and Astronomy, Rutgers, The State University of NJ, USA 5 Josef Stefan Institute, Liubliana, Slovenia

6 Theoretical Particle Physics and Cosmology, King's College London, United Kingdom





ML for LHC

Top tagging

[classification, 2016-today]

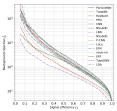
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Driving NN-architectures



The Machine Learning Landscape of Top Taggers

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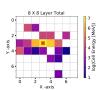
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6 Theoretical Particle Physics and Cosmology, King's College London, United Kingdom 7 Department of Physics and Astronomy, The University of British Columbia, Canada 8 Department of Physics, University of California, Sunta Barbara, USA 9 Faculty of Mathematics and Physics, University of Ljubljana, Ljubljana, Slovenia 10 Center for Theoretical Physics, MIT, Cambridge, USA 11 CP3, Universitéex Catholique de Louvain, Louvain-la-Neuve, Belgium

12 Physics Division, Lawrence Berkeley National Laboratory, Berkeley, USA 13 Simons Inst. for the Theory of Computing, University of California, Berkeley, USA 14 National Institute for Subatomic Physics (NIKHEF), Amsterdam, Netherlands 15 LPTHE, CNRS & Sorboune Université, Paris, France 16 III. Physics Institute A, RWTH Aachen University, Germany

Particle flow [2020-today]

- · mother of jet analyses
- combining detectors with different resolution
- · optimality the key
- → Modern jet analysis basics



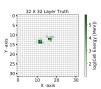
Towards a Computer Vision Particle Flow *

Francesco Armando Di Bello^{k,1}, Sanmay Ganguly^{k,1}, Eilam Gross¹, Marumi Kado^{k,4}, Michael Pitt², Lorenzo Santi ³, Janathan Shlomi Weizmann Institute of Science, Reboyot 76100, Israel

CERN, CH 1211, Geneva 23, Switzerland Università di Roma Sapienza, Piazza Aldo Moro, 2, 00185 Roma, Italy e INFN, Italy ⁶Université Paris-Saclay, CNRS/IN2P3, IJCLab, 91405, Orsay, France

Progress towards an improved particle flow algorithm at CMS with machine learning

> Farouk Mokhtar¹, Joosen Pata², Javier Duarte¹, Eric Wulff², Maurizio Pierini² and Jean-Roch Vlimant⁴ (on behalf of the CMS Collaboration) *Cuiversity of Coldierain San Diege, La Jolla, CA 92093, USA *SNCPR, Rivola pet 30, 2013 Tallian, Solonia *European Congustation for Nuclius Bouncaris (CERN), CR 1211, Graves 23, Switserland Evani decisterbond.eds, josep.patebons.sk, jdoortebond.eds





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ML for LHC

Uncertainties

Bayesian NI

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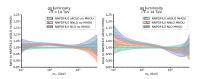
Conciduo

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Parton densities [NNPDF, 2002-today]

- · pdfs without functional bias and full uncertainties
- precision and calibrated uncertainties
- → Drivers of ML-theory

ML in phenomenology



The Path to N¹LO Parton Distributions

The NNPDF Collaboration:

Richard D. Bali¹, Andrea Barentink², Assessable Candids², Stofano Carrazza², Juan Crus-Martinea²,
Luigi Dei Debble³, Soriano Forte², Tamanos Gami², Pikit Helkom^{2,2}, Zahari Kasoloo²,
Niccoli Larrenti², Cincome Magai^{1,5}, Encarele R. Nocera³, Tsajeon R. Hebermansjara^{1,5}, Juan Righ^{1,5}
Cincipiles Stories², Roy Stopmans, and Maria Unide.

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This paper is dedicated to the memory of Stefano Catani,

 $Grand\ Master\ of\ QCD,\ great\ scientist\ and\ human\ being$



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ML for LHC

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Uncertaintie

Ensembles

Amplitudes

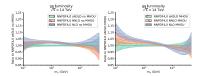
Generation

Amplificati

ML in phenomenology

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Niccolò Laurenta², Giacomo Magni¹, Zimanosle B., Nocera², Teajens B. Hobermanoj ara⁴, June Bojs^{4,3}
Christother Schrau², Ber Stemman, and Maria Unitie².

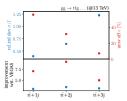
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Grand Master of QCD, great scientist and human being

Ultra-fast event generators [Sherpa, MadNIS, MLHad]

- · event generation modular
- · improve and replace by ML-modules
- → Beat state of the art

Triple-W VBS W+jets tī+jets	$\begin{split} u\bar{d} &\rightarrow W^+W^+W^-\\ uc &\rightarrow W^+W^+ds\\ gg &\rightarrow t\bar{t}+g \end{split}$	$\begin{array}{c} gg \rightarrow W^+ d\bar{u}g \\ gg \rightarrow t\bar{t} + gg \end{array}$	$gg \to W^+ d\bar{u}gg$ $gg \to t\bar{t} + ggg .$	



EdPod Physica Eaberic The MatNIS Reloaded

Theo Heimel², Nazhan Huetach¹, Fabio Maltone^{2,3},
Olivier Mattelaer², Tibron Plehn², and Ramon Winserhalder²

1 Institut für Theocesische Physik, Universität Heidelberg, Germany
2 CF3, Université carbolique de Louvais, Louvais-la-Neure, Belgius
3 Dipartisenson di Fisica e Astronomis, Universitá id Belogas, Alori

December 17, 2024

Abstrac

In pursuit of precise and fast theory predictions for the LPK, we present an implementation of the MARNS method in the MAROSAUS event generator. A series of improvements in MARNS further enhance in efficiency and speed. We validate this implementation for realistic partonic processes and find significant gains from using modern machine learning in event generators.



ML for LHC

Optimizing integration paths [invertible networks]

· find optimal integration paths

· learn variable transformation

→ Theory-integrator

ML in theory





SciPost Phys. 12, 129 (2022)

Targeting multi-loop integrals with neural networks Ramon Winterhalder^{1,2,3}, Vitaly Magerya⁴, Emilio Villa⁴, Stephen R Jones³, Matthias Kerner^{4,6}, Anja Butter^{1,2}, Gudrun Heinrich^{2,4} and Tilman Pleha^{1,2}

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Université catholique de Louvain, Belvium 4 Institut für Theoretische Physik. Karlsruher Institut für Technologie. Germans 5 Institute for Particle Physics Phenomenology, Durham University, UK 6 Institut für Astroteilchenebysik. Karlsruher Institut für Technologie. Germany

Numerical evaluations of Feynman integrals often proceed via a deformation of the integration contour into the complex plane. While valid contours are easy to construct the numerical precision for a multi-loop integral can depend critically on the chosen contour. We present methods to optimize this contour using a combination of optimized, elobal complex shifts and a normalizing flow. They can lead to a significant wain in precision.

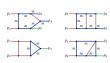


ML for LHC

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Sci Post

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Navigating string landscape [reinforcement learning]

- searching for viable vacua
- high dimensions, unknown global structure
- → Model space sampling





Figure 1: Left: Cluster structure in dimensionally reduced flux samples for RL and 25 GA runs (PCA on all samples of GA and RL). The colors indicate individual GA runs. Right: Dependence on flux (input) values (N3 and N5 respectively) in relation to principal components for a PCA fit of the individual output of GA and RL.

Probing the Structure of String Theory Vacua with Genetic Algorithms and Reinforcement Learning

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rield Center for Theoretical Physics LMU Munich even krippendorf@physik.uni-maenchen.de

Andreas Schachm Centre for Mathematical Sciences

University of Wisconsin-Madison shiu@ohysics.wisc.edu

Abstract

Identifying string theory vacua with desired physical properties at low energies requires searching through high-dimensional solution spaces - collectively referred to as the string landscape. We highlight that this search problem is amenable to able to reveal novel features (suggesting previously unidentified symmetries) in the to identify these features robustly, we combine results from both search methods, which we argue is imperative for reducing sampling bias



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Uncertainties

Ensemble

. . . .

Learned uncertainties

Network training as a fit

- · learn scalar field $f_{\theta}(x) \approx f(x)$
- · statistics: maximize parameter probability given (f_j, σ_j)

$$p(\theta|x) = \frac{p(x|\theta) \ p(\theta)}{p(x)}$$

→ maximize tractable likelihood instead

$$\begin{split} p(x|\theta) &= \prod_{j} \frac{1}{\sqrt{2\pi}\sigma_{j}} \exp\left(-\frac{|f_{j} - f_{\theta}(x_{j})|^{2}}{2\sigma_{j}^{2}}\right) \\ \Rightarrow \qquad \mathcal{L} \equiv -\log p(x|\theta) &= \sum_{j} \frac{|f_{j} - f_{\theta}(x_{j})|^{2}}{2\sigma_{j}^{2}} + \text{const}(\theta) \end{split}$$



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Uncertainties

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Learned uncertainties

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Learned local uncertainty

· Gaussian log-likelihood with normalization

$$\mathcal{L}_{\mathsf{heteroscedastic}} = rac{|f(x) - f_{ heta}(x)|^2}{2\sigma_{ heta}(x)^2} + \log \sigma_{ heta}(x) + \cdots$$

- · if needed replace $\sigma_{\theta}(x)$ by mixture model
- \rightarrow learn $f_{\theta}(x)$ and $\sigma_{\theta}(x)$ together



Bayesian networks

Learned function statistically

· amplitude over phase phase

$$\langle A \rangle = \int dA \ A \ p(A)$$

· internal representation θ of training data T [think Gaussian with mean and width]

$$p(A) = \int d\theta \ p(A|\theta) \ p(\theta|T)$$

 $\rightarrow \theta$ -distribution defining Bayesian NN



Bayesian networks Tilman Plehn

Learned function statistically

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· internal representation θ of training data T [think Gaussian with mean and width] $p(A) = \int d\theta \ p(A|\theta) \ p(\theta|T)$

 $\rightarrow \theta$ -distribution defining Bayesian NN

Variational approximation

definition of training

$$p(A) = \int d\theta \ p(A|\theta) \ p(\theta|T) \approx \int d\theta \ p(A|\theta) \ q(\theta)$$

· similarity through minimal KL-divergence [Bayes' theorem to remove unknown posterior]

$$\begin{split} D_{\mathsf{KL}}[q(\theta),p(\theta|T)] &= \int d\theta \; q(\theta) \; \log \frac{q(\theta)}{p(\theta|T)} \\ &= \int d\theta \; q(\theta) \; \log \frac{q(\theta)p(T)}{p(T|\theta)p(\theta)} \\ &\approx D_{\mathsf{KL}}[q(\theta),p(\theta)] - \int d\theta \; q(\theta) \; \log p(T|\theta) \equiv \mathcal{L} \end{split}$$



→ Two-term loss: likelihood + prior

Statistics vs systematics Tilman Plehn

Statistical network evaluation

· expectation value from network $q(\theta)$

$$\langle A \rangle = \int dA d\theta \ A \ p(A|\theta) \ q(\theta)$$

$$\equiv \int d\theta \ q(\theta) \overline{A}(\theta) \qquad \text{with} \qquad \overline{A}(\theta) = \int dA \ A \ p(A|\theta)$$

corresponding variance

$$\begin{split} \sigma_{\text{tot}}^2 &= \int dA d\theta \ (A - \langle A \rangle)^2 \ p(A|\theta) \ q(\theta) \\ &= \int d\theta \ q(\theta) \left[\overline{A^2}(\theta) - \overline{A}(\theta)^2 + \left(\overline{A}(\theta) - \langle A \rangle \right)^2 \right] \equiv \sigma_{\text{syst}}^2 + \sigma_{\text{stat}}^2 \end{split}$$

Two uncertainties

· statistical — vanishing for perfect training: $q(\theta) \rightarrow \delta(\theta - \theta_0)$

$$\sigma_{\text{stat}}^2 = \int d\theta \ q(\theta) \left[\overline{A}(\theta) - \langle A \rangle \right]^2$$

· systematic — vanishing for perfect data: $p(A|\theta) \rightarrow \delta(A - A_0)$

$$\sigma_{\mathsf{syst}}^2 = \int d\theta \ q(\theta) \left[\overline{A^2}(\theta) - \overline{A}(\theta)^2 \right]$$





Repulsive ensembles

Posterior from network ensemble

OED vs continuity equation

$$\frac{d\theta}{dt} = v(\theta, t) \qquad \Leftrightarrow \qquad \frac{\partial \rho(\theta, t)}{\partial t} = -\nabla_{\theta} \left[v(\theta, t) \rho(\theta, t) \right]$$

· Fokker-Planck equation with stationary $\rho(\theta, t) = \pi(\theta)$

$$\frac{d\theta}{dt} = -\nabla_{\theta} \log \frac{\rho(\theta, t)}{\pi(\theta)}$$

ODE describing training progress

$$\begin{split} \boldsymbol{\theta}^{t+1} - \boldsymbol{\theta}^t &\propto -\nabla_{\boldsymbol{\theta}^t} \left[\log \rho(\boldsymbol{\theta}^t) - \log \pi(\boldsymbol{\theta}^t) \right] \\ &= -\nabla_{\boldsymbol{\theta}^t} \left[\log \sum_j k(\boldsymbol{\theta}^t, \boldsymbol{\theta}_j^t) - \log p(\boldsymbol{\theta}|\mathbf{x}_{\text{train}}^t) \right] \equiv -\nabla_{\boldsymbol{\theta}^t} \mathcal{L}_{\text{RE}} \end{split}$$

→ Joint ensemble training



Incertaintie

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ML for LHC

Uncertaintie Bayesian NN

Ensembles

Generation

Repulsive ensembles

Posterior from network ensemble

OED vs continuity equation

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→ Joint ensemble training

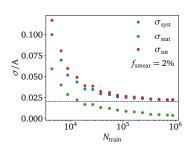
Repulsive ensembles

- · train network ensemble
- apply repulsive force kernel in function space
- → Alternative for statistical uncertainty



Loop amplitude $gg o \gamma \gamma g(g)$ [Bahl, Elmer, Favaro, Haussmann, TP, Winterhalder]

- · regression of exact scalar over phase space [Aylett-Bullock, Badger, Moodie]
- · example systematics: artificial noise
- · assume $\sigma_{\text{stat}} \ll \sigma_{\text{syst}}$

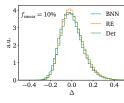


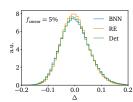


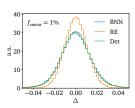
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- · assume $\sigma_{\text{stat}} \ll \sigma_{\text{syst}}$
- · accuracy over phase space

$$\Delta(x) = \frac{A_{\mathsf{NN}}(x) - A_{\mathsf{true}}(x)}{A_{\mathsf{true}}(x)}$$









Amplitudes

Network amplitudes

Loop amplitude $gg o \gamma \gamma g(g)$ [Bahl, Elmer, Favaro, Haussmann, TP, Winterhalder]

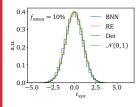
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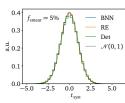
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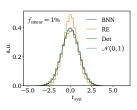
pull over phase space

$$t(x) = \frac{A_{\text{NN}}(x) - A_{\text{true}}(x)}{\sigma(x)}$$

→ calibrated leading systematics









certainties

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ML for LHC

Uncertaintie

Ensembles

Amplitudes

Amplification

Network amplitudes

Loop amplitude $gg o \gamma \gamma g(g)$ [Bahl, Elmer, Favaro, Haussmann, TP, Winterhalder]

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$$\Delta(x) = \frac{A_{\mathsf{NN}}(x) - A_{\mathsf{true}}(x)}{A_{\mathsf{true}}(x)}$$

pull over phase space

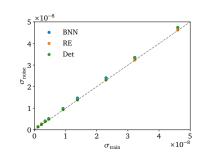
$$t(x) = \frac{A_{\text{NN}}(x) - A_{\text{true}}(x)}{\sigma(x)}$$

→ calibrated leading systematics

Towards zero noise

· extrapolate to zero noise

$$\sigma_{\text{noise}}^2 = \sigma_{\text{syst}}^2 - \sigma_{\text{syst,0}}^2 \approx \sigma_{\text{train}}$$





certainties

Tilman Plehn

ML for LHC

Uncertainties

Amplitudes

Generation

Network amplitudes

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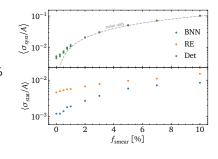
→ calibrated leading systematics

Towards zero noise

· extrapolate to zero noise

$$\sigma_{
m noise}^2 = \sigma_{
m syst}^2 - \sigma_{
m syst,0}^2 pprox \sigma_{
m train}$$

- · systematics plateau $\langle \sigma/A \rangle \sim 0.4\%$
- → Limiting factor??



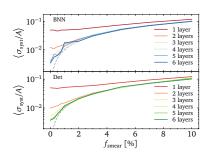


Amplitudes

Improved accuracy

Network expressivity

- · large range of amplitude values
- · resolution of (collinear) peaks
- · network breaks for large amplitudes
- · 3 hidden layers needed
- · activation function machine precision...





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Uncertaintie: Bayesian NN

Amplitudes

Generation

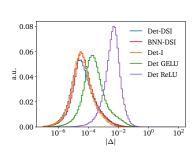
Improved accuracy

Network expressivity

- · large range of amplitude values
- · resolution of (collinear) peaks
- · network breaks for large amplitudes
- · 3 hidden layers needed
- activation function machine precision...

Data pre-processing

- · amplitude from invariants
- · learn Minkowski metric?
- · Deep-sets-invariant network [Heinrich etal] L-GATr transformer





Improved accuracy

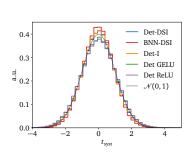
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- amplitude from invariants
- learn Minkowski metric?
- Deep-sets-invariant network [Heinrich etal] L-GATr transformer
- · uncertainty scaling with accuracy pull unit Gaussian
- → Calibrated leading systematics





Amplitudes

Improved accuracy

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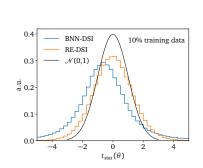
Statistical uncertainties

- · well-defined for $\sigma_{\rm stat} \ll \sigma_{\rm syst}$ calibration of $\sigma_{\mathsf{tot}} \sim \sigma_{\mathsf{syst}}$
- · systematics per network statistics sampled
- · calibration from pull

$$t_{\mathrm{stat}}(x, \theta) = rac{\overline{A}(x, \theta) - \langle A \rangle(x)}{\sigma_{\mathrm{stat}}(x)} \ pprox rac{\overline{A}(x, \theta) - A_{\mathrm{true}}(x)}{\sigma_{\mathrm{stat}}(x)}$$

→ Work to do...

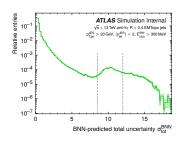


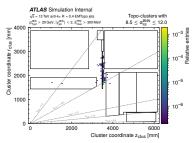


- Energy calibration with uncertainties [ATLAS + Heidelberg, 2412.04370]
 - · interpretable calorimeter phase space x
 - · learned calibration function

$$\mathcal{R}^{\mathsf{BNN}}(x) \pm \Delta \mathcal{R}^{\mathsf{BNN}}(x) pprox rac{\mathsf{E}^{\mathsf{obs}}(x)}{\mathsf{E}^{\mathsf{dep}}(x)}$$

- · systematics: noisy training data ...
- → Understand (simulated) detector







Generation

Generative AI with uncertainties

Generative Ai with uncertainties

Bayesian generative networks [Bellagente, Haussmann, Luchmann, TP]

- network weight distributions for density
- sampling phase space events with error bars on weights
- learned density & uncertainty reflecting network learning?
- → Generative networks like fitted densities



Generative AI with uncertainties

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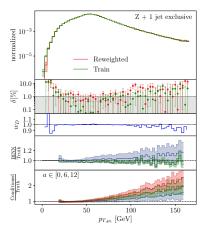
Z+jets events [Heimel, Vent...]

- · per-cent accuracy on density
- · statistical uncertainty from BNN
- · systematics in training data

$$w = 1 + a \left(\frac{p_{T,j_1} - 15 \text{ GeV}}{100 \text{ GeV}} \right)^2$$

sampling a conditionally

→ Precision and uncertainty control





Generation

Controlling generative Al

Compare generated with training data

- · regression accuracy $\Delta = (A_{\text{data}} A_{\theta})/A_{\text{data}}$
- · harder for generation, unsupervised density classify training vs generated events D(x)learned density ratio [Neyman-Pearson]

$$w(x_i) = \frac{D(x_i)}{1 - D(x_i)} = \frac{p_{\text{data}}(x_i)}{p_{\text{model}}(x_i)}$$

→ Test ratio over phase space



certainties

Tilman P

ML for LHC

Uncertaintie Bayesian N

Amplitudes

A----lifi--

Controlling generative Al

Compare generated with training data

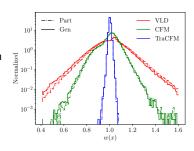
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Progress in NN-unfolding

- · generative network example
- · compare different architectures
- · accuracy from width of weight distribution
- · tails indicating failure mode
- \rightarrow Systematic performance test



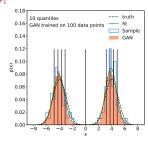


Amplification

Amplification (for Louis)

Improving training data [Butter, Diefenbacher, Kasieczka, Nachman , TP]

- true function known compare GAN vs sampling vs fit
- $-\chi^2$ -sum of quantiles





certainties

Tilman Plehn

Tilman Pie

ML for LHC

Uncertaintie Bayesian NN

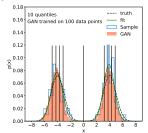
Amplitude

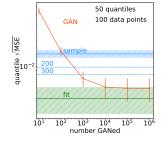
Generation

Amplification (for Louis)

Improving training data [Butter, Diefenbacher, Kasieczka, Nachman , TP]

- true function known compare GAN vs sampling vs fit
- $-\chi^2$ -sum of quantiles
- training and fit with 100 data points
- fit like 500-1000 sampled points
 GAN like 500 sampled points [amplification factor 5]
 requiring 10,000 GANned events
- ⇒ Generative networks like fits







certaintie

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ML for LHC

Uncertainties

Ensembles

Generation

Amplificatio

Modern ML for LHC

Developing ML for the best science

- · just another numerical tool for a numerical field
- · transformative new common language
- · driven by money from data science and medical research
- · be 10000 Einsteins,
 - ...improving established tools
 - ...developing new tools for established tasks
 - ...transforming through new ideas
- → It's the future, let's not miss it

Modern Machine Learning for LHC Physicists

Tilman Plehn^a; Anja Butter^{a,b}, Barry Dillon^a, Theo Heimel^a, Claudius Krause^c, and Ramon Winterhalder^d

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^b LPNHE, Sorbonne Université, Université Paris Cité, CNRS/INZP3, Paris, France
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^d CP3, Université catholique de Louvain, Louvain-la-Neuve, Belgium

March 19, 2024

Abstract

Modern machine learning is transforming particle physics, fat, hallying its way into our numerical tool box. For young researchers it is exceeded to any on up of the development, which means perhips quitting edge methods and books to the first exceeding the continuous of the continuous for machine learning to relevant applications. They start with an LHC openite motivation and associated architecture for machine learning to relevant applications. They start with an LHC openite motivation and associated architecture in toward and accordinate and associated associated



Generative Uncertainties

Unsupervised Bayesian networks

- data: event sample [points in 2D space] learn phase space density standard distribution in latent space [Gaussian] sample from latent space
 - Bayesian version allow weight distributions learn uncertainty map
- 2D wedge ramp

$$p(x) = ax + b = ax + \frac{1 - \frac{a}{2}(x_{\text{max}}^2 - x_{\text{min}}^2)}{x_{\text{max}} - x_{\text{min}}}$$
$$(\Delta p)^2 = \left(x - \frac{1}{2}\right)^2 (\Delta a)^2 + \left(1 + \frac{a}{2}\right)^2 (\Delta x_{\text{max}})^2 + \left(1 - \frac{a}{2}\right)^2 (\Delta x_{\text{min}})^2$$

explaining minimum in $\sigma(x)$



